



Venezia
2018 Quark Matter

The 27th International Conference
on Ultrarelativistic
Nucleus-Nucleus Collisions

14-19 May Palazzo del Cinema
Lido di Venezia, Italy

Have we seen the QGP?

Thanks to QM2018

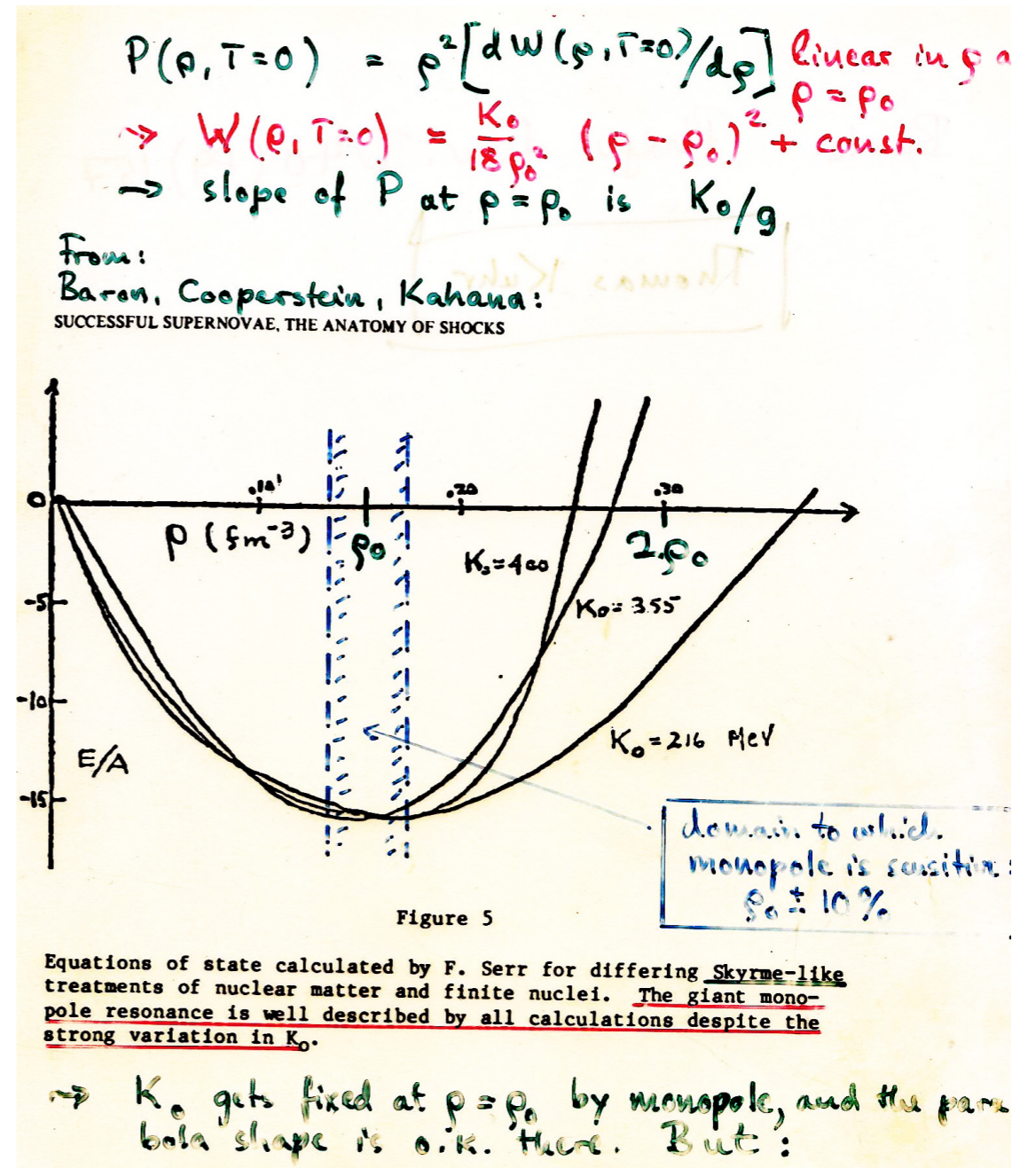
In-Kwon Yoo (Pusan Nat'l Univ.)

Issues

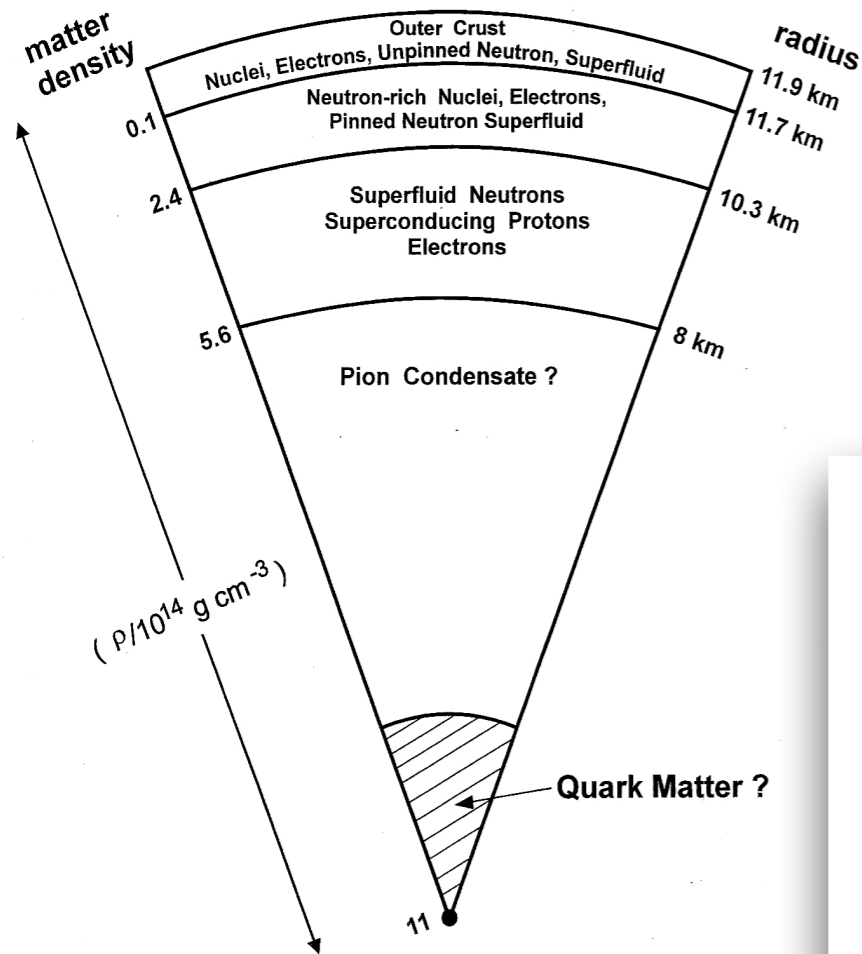
- What is QGP?
 - Historical Overview
- How can we know it?
 - Signatures of its existence
 - Its properties
- Where are we?
- Where to go?

Historical Overview

- R. Stock's lecture at QM2018
- Nuclear/Hadronic Matter
- EoS / Neutron Stars / Supernovae
- Nuclear Physics Community at LBL Bevalac (1974 - 1985)
- Advent of QCD in AA collisions



Structure of neutron stars



as of 1974
with later addition of
Quark Matter Core
Hypothesis

Oppenheimer Volkov
Hydro-static equilibrium 1956

Nuclear Shock Waves in Heavy-Ion Collisions

Werner Scheid, Hans Müller, and Walter Greiner

Institut für Theoretische Physik der Universität Frankfurt, Frankfurt am Main, Germany
(Received 19 November 1973)

It is shown that nuclear matter is compressed during the encounter of heavy ions. If the relative velocity of the nuclei is larger than the velocity of first sound in nuclear matter (compression sound for isospin $T=0$), nuclear shock waves occur. They lead to densities which are 3-5 times higher than the nuclear equilibrium density ρ_0 , depending on the energy of the nuclei. The implications of this phenomenon are discussed.

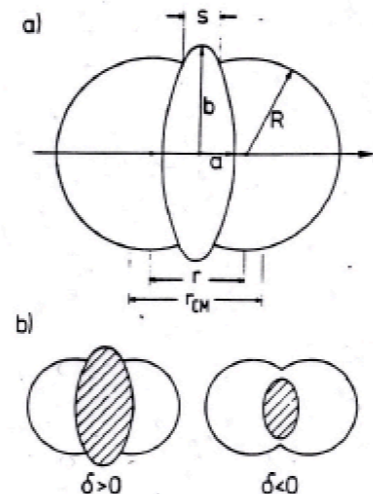


FIG. 1. (a) Geometric parameters of the model. (b) Two cases $\delta > 0$ and $\delta < 0$. The unphysical situation $\delta < 0$ is excluded by forces of constraints.

PRL 1974

The paper that started
nuclear
hydrodynamics

„Hugoniot-Rankine“ shock
compression as precursor of
hydrodynamics

Initial idea:
Glassgold, Heckrotte and
Watson,
Ann. Phys. 6(1959) 1

p+A !

A + A

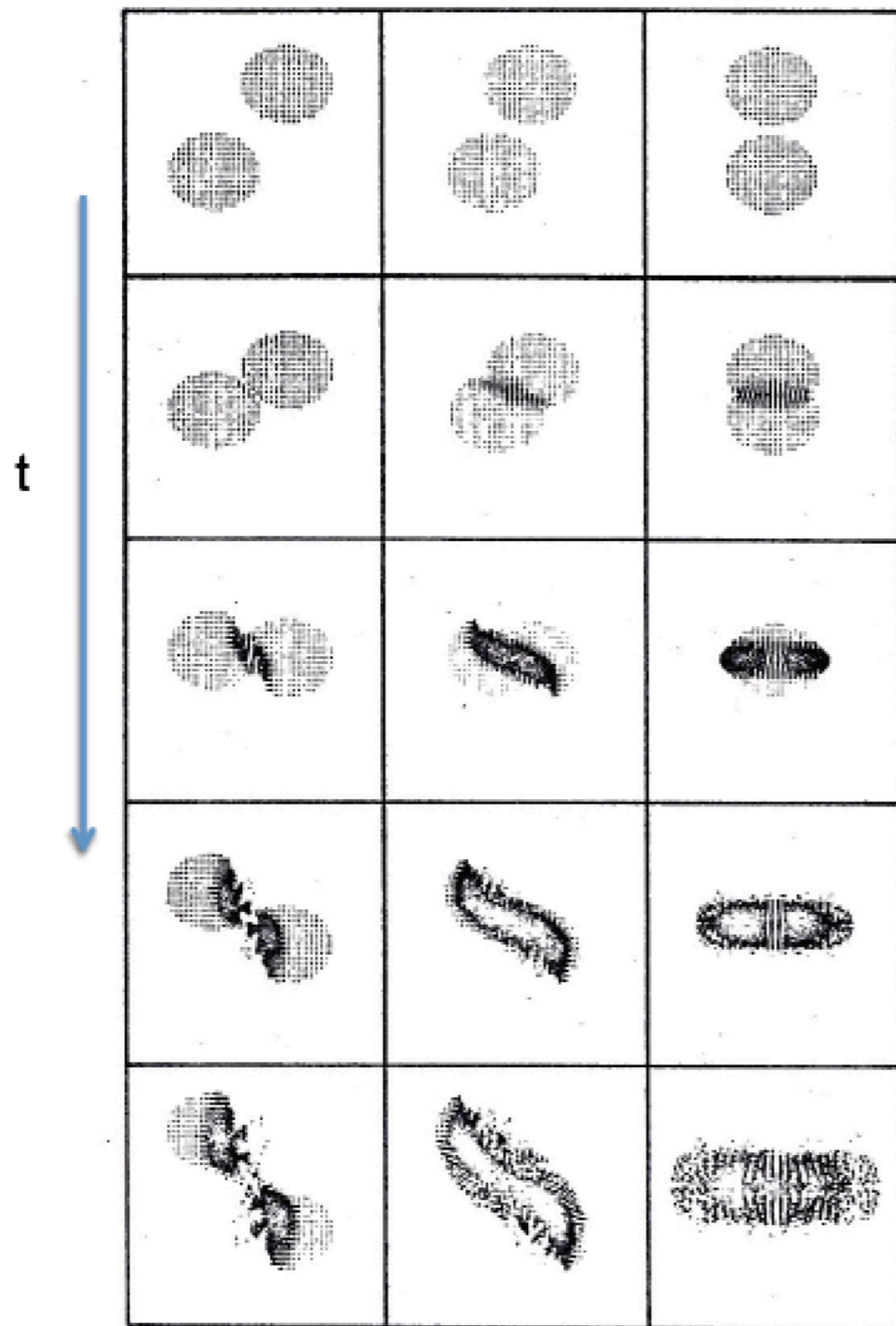
182 MeV/A

CM

$b = 0.8 b_{max}$

$b = 0.4 b_{max}$

$b = 0$

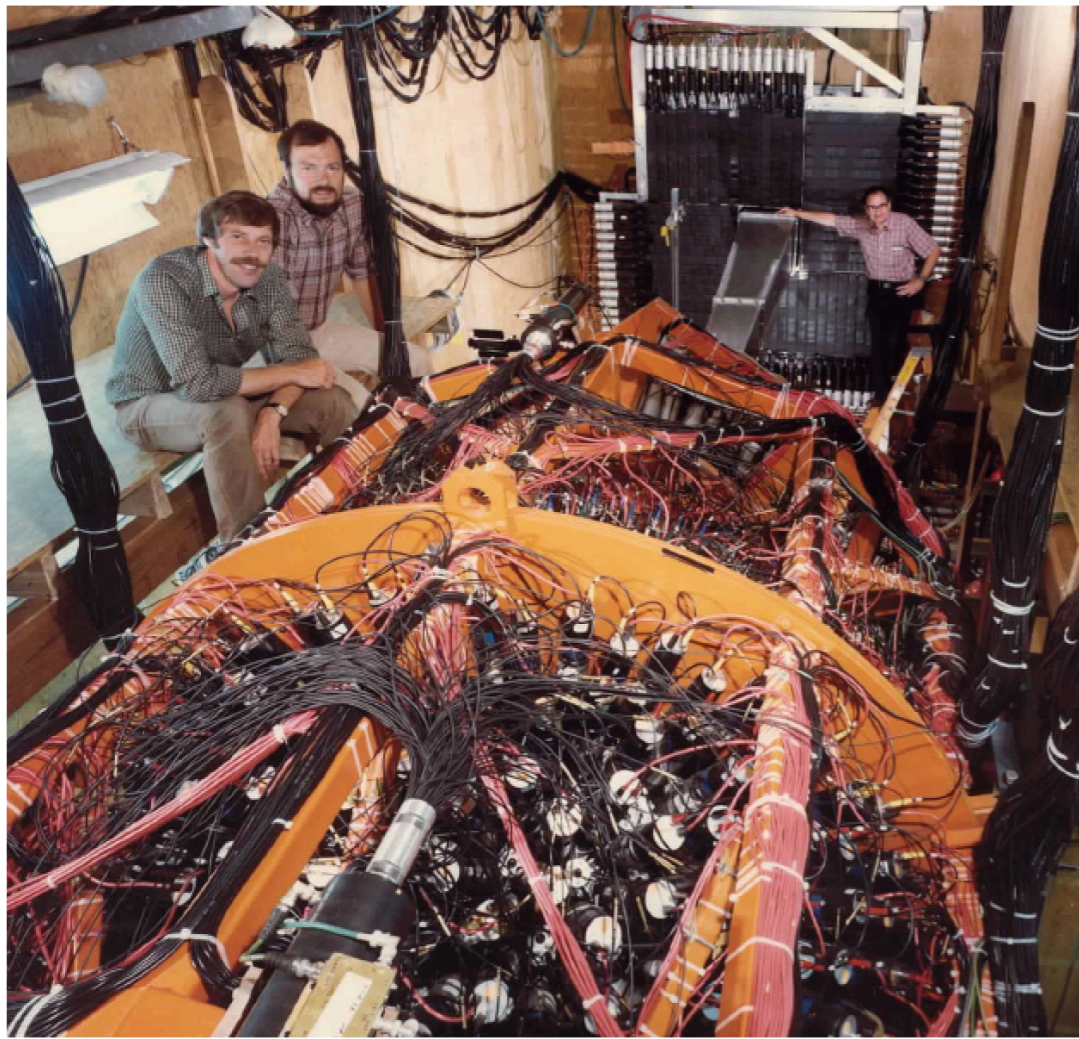


HYDRODYNAMICS!

R. Nix, Los Alamos Bomb Code

$b \rightarrow 0$:
radial flow expansion

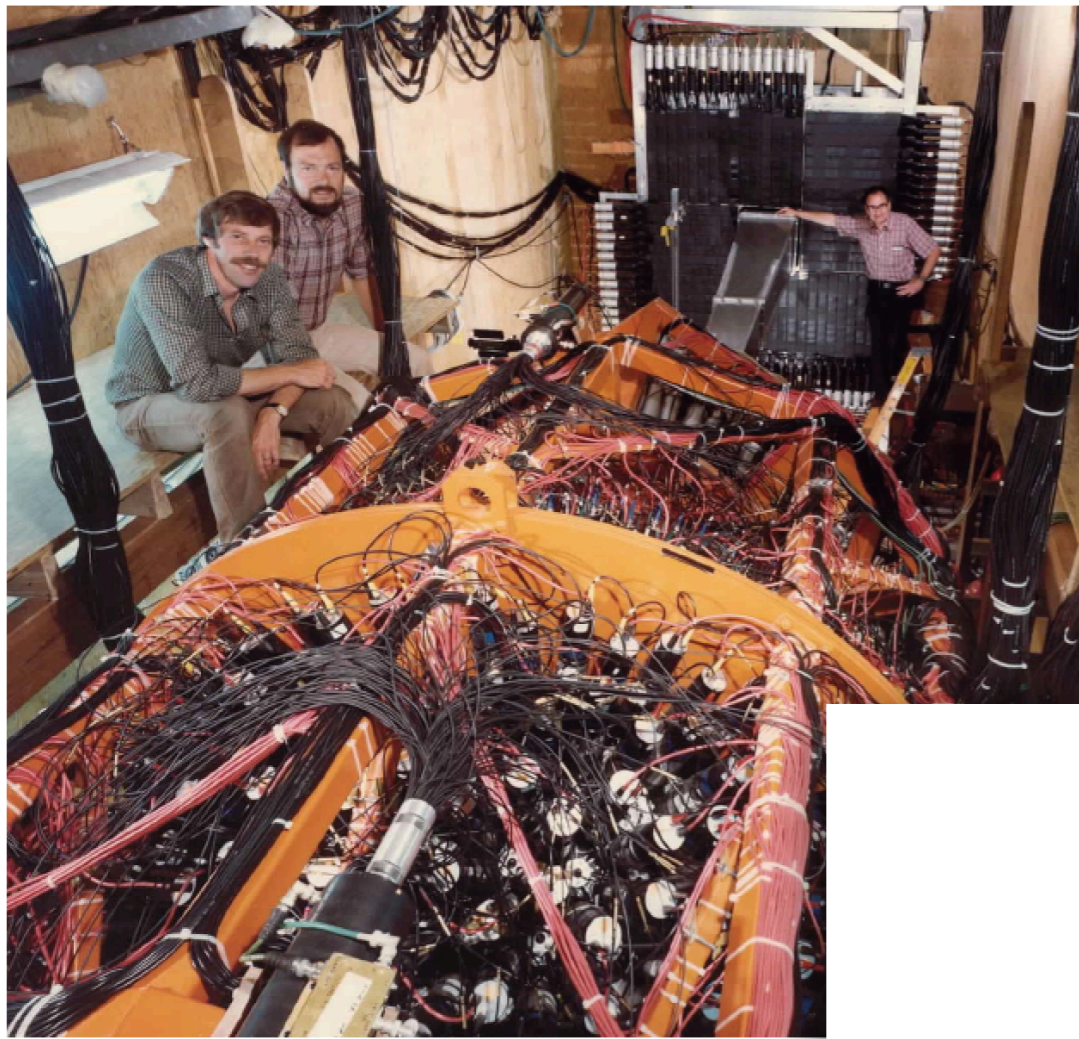
Intermediate b :
sideward flow, now v_1



Plastic Ball

850x
dE/dx-E

Gutbrod
Poskanzer
Ritter



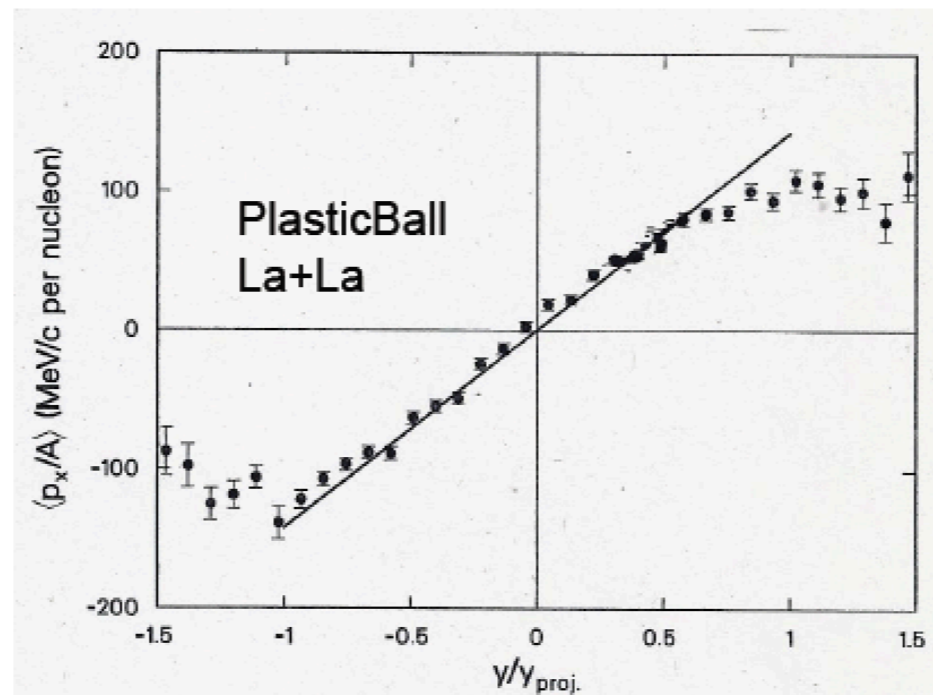
Plastic Ball

850x
dE/dx-E

Gutbrod

Deduce EoS from
HYDRODYNAMICS!

$F(y) = \text{Slope at mid-rap.}$

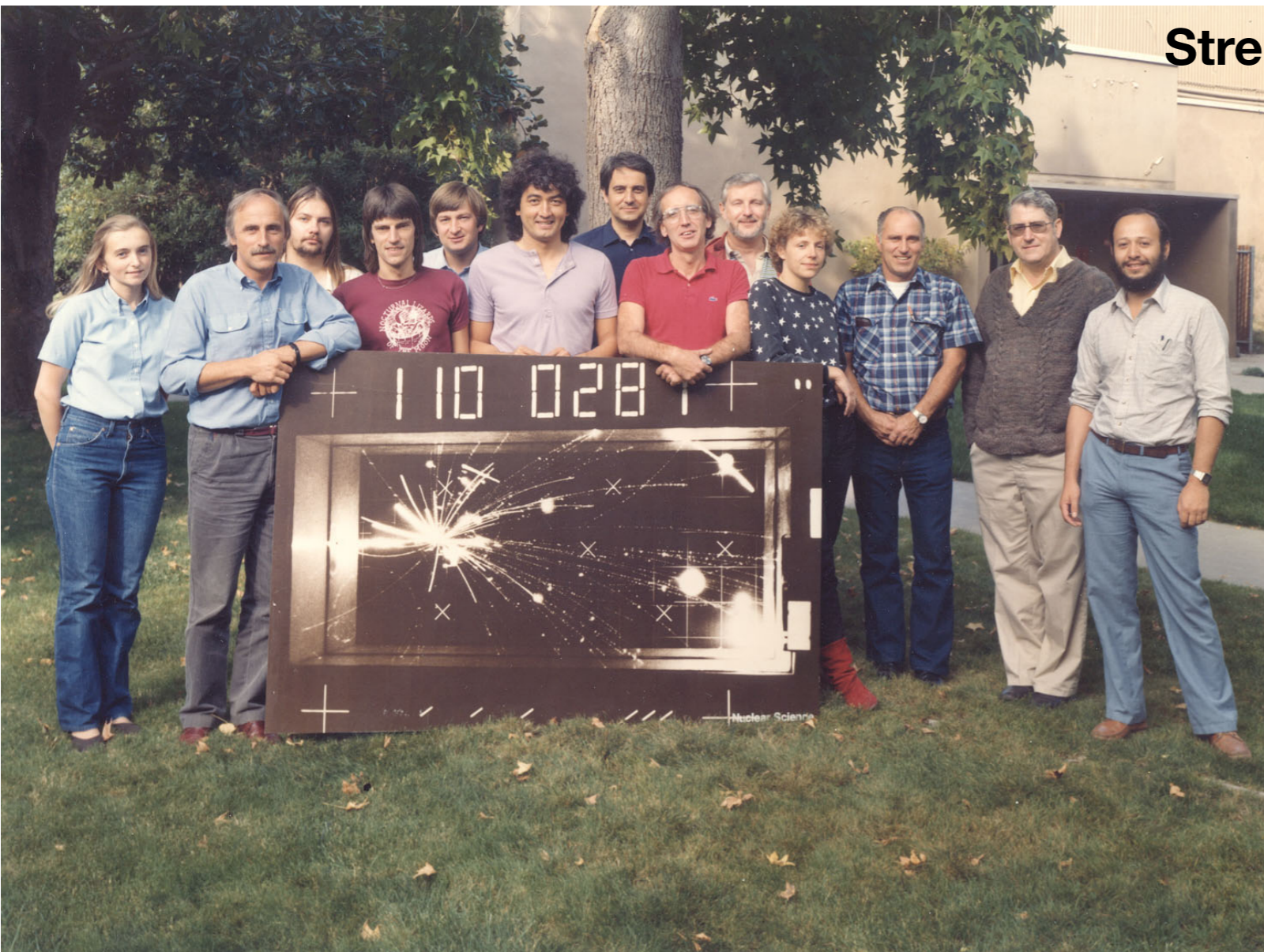


$$F(y) = P_{eff} \times S \times t_{pass}$$

↓
EOS

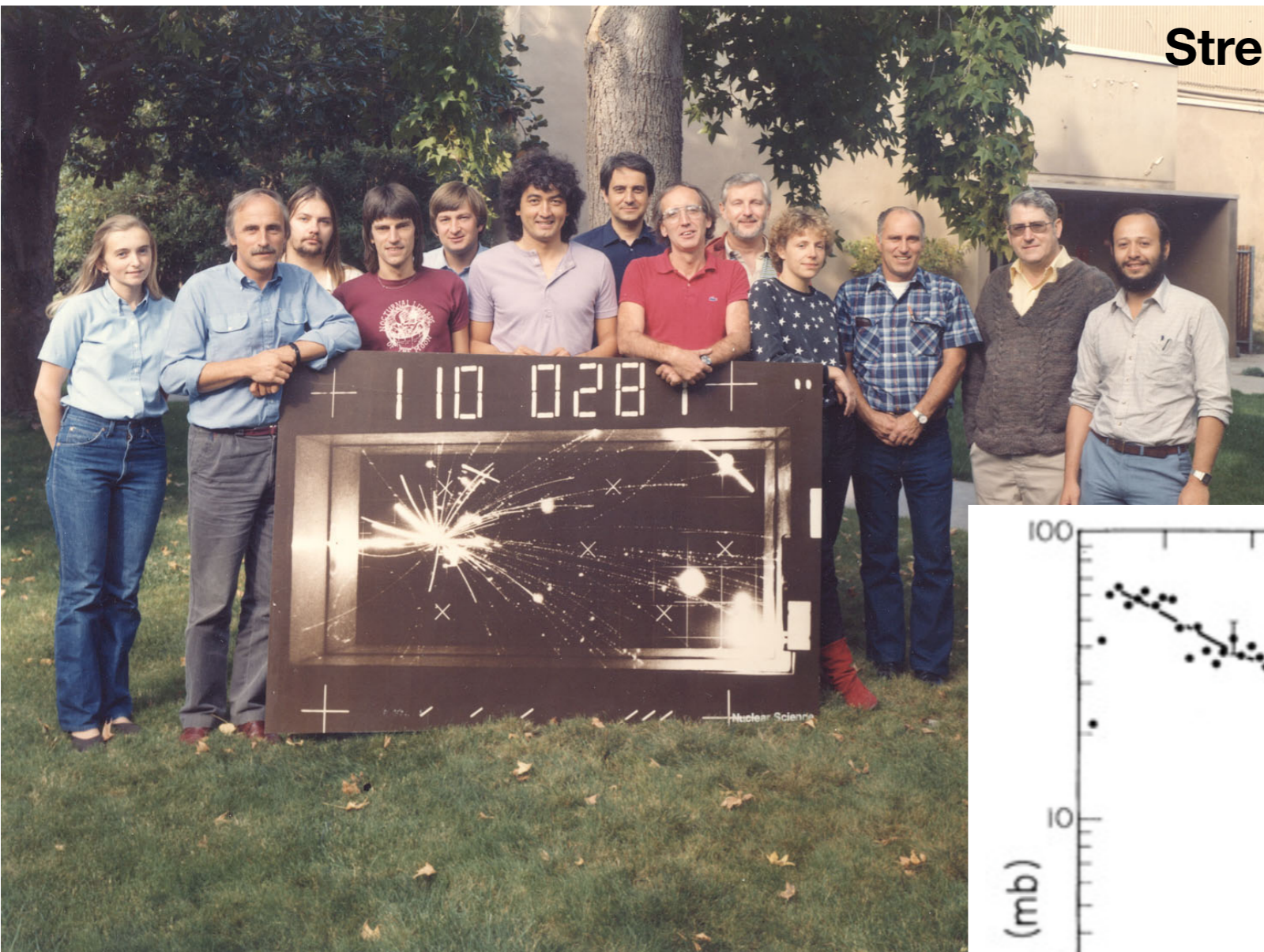
A long story since 1989!

Streamer Chamber Collab. at LBL 1983

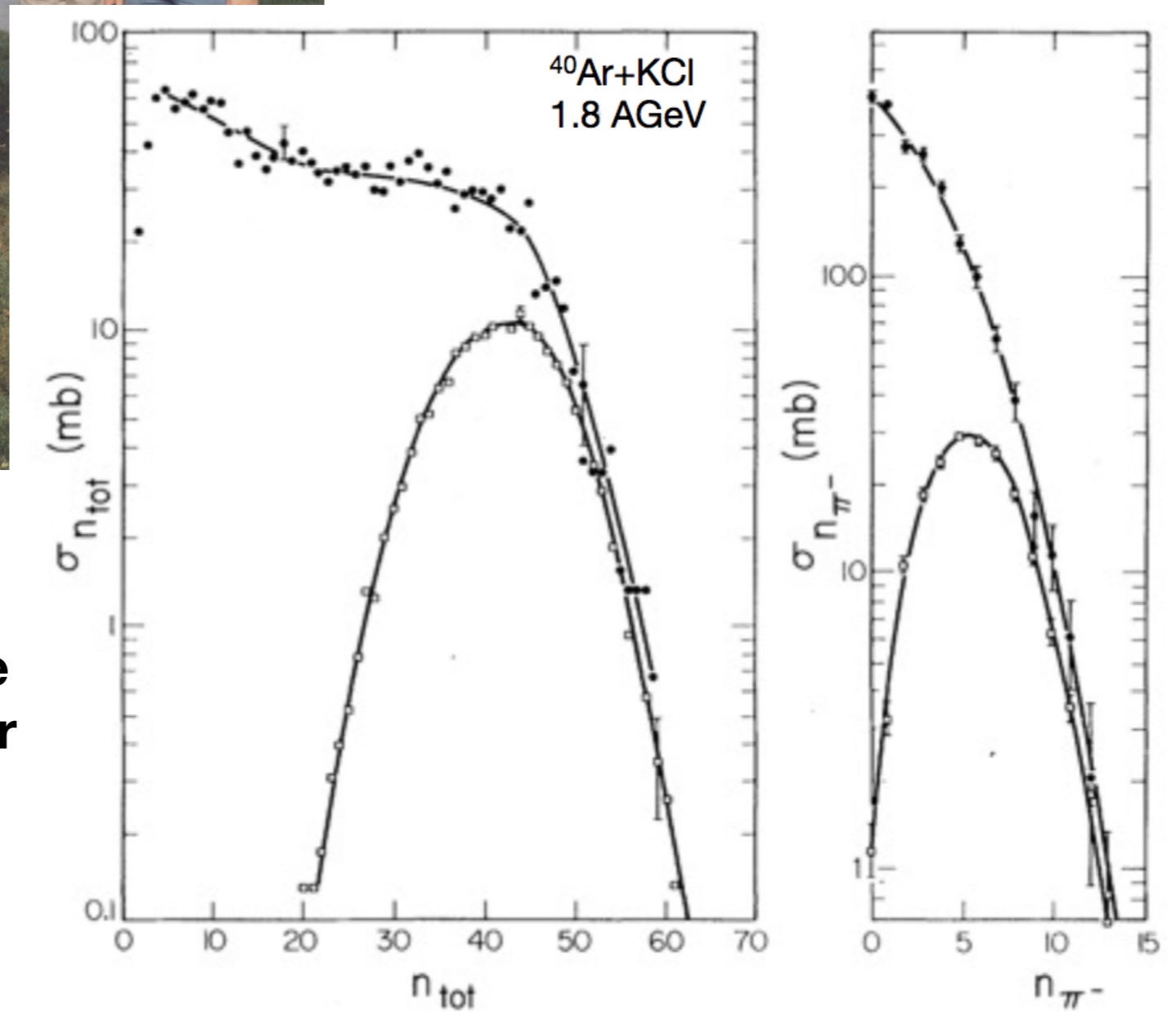


**Centrality Selection and Event plane
with the historical streamer chamber**

Streamer Chamber Collab. at LBL 1983



**Centrality Selection and Event plane
with the historical streamer chamber**



Phys. Rev. Lett. 45 (1980) 874

Part II:

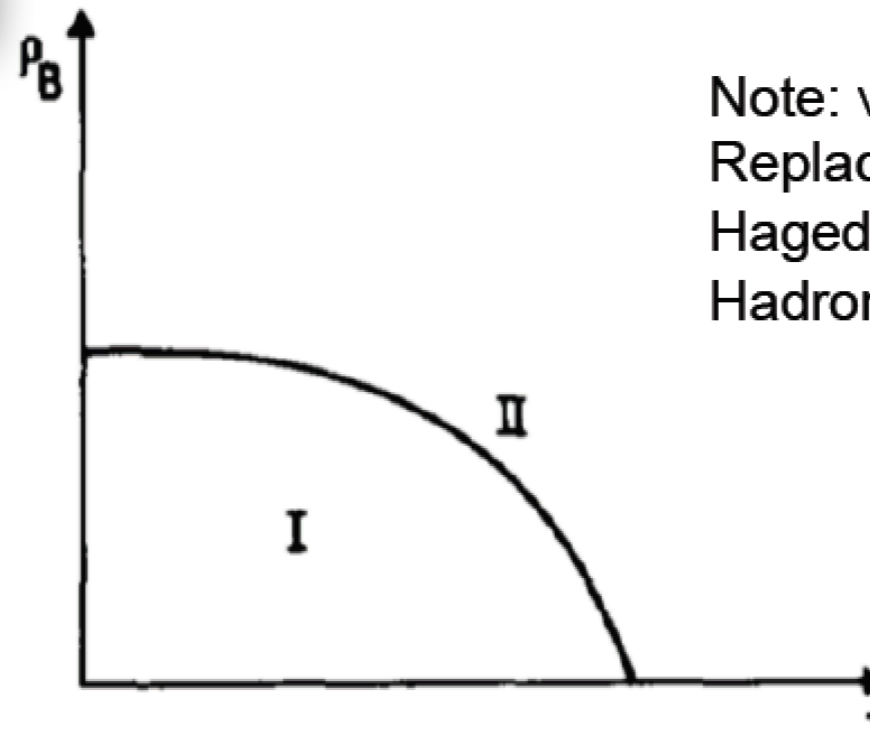
The Advent of QCD in A+A Collisions

Hagedorns limiting hadronic temperature T_H
as interpreted by Cabibbo and Parisi 1975
as indication of a second order phase transition to QGP
Phys.Lett. B59 (1975) 67

The currently accepted interpretation of the properties of hadronic matter is based on the “realistic” quark model where quarks are permanently confined in hadrons. We expect models of this kind to give rise to a phase transition at a temperature $kT \approx m_\pi$, the high temperature phase being one where quarks can move freely in space.

-> Deconfinement at T_H to a new QCD phase

The first QCD Phase Diagram

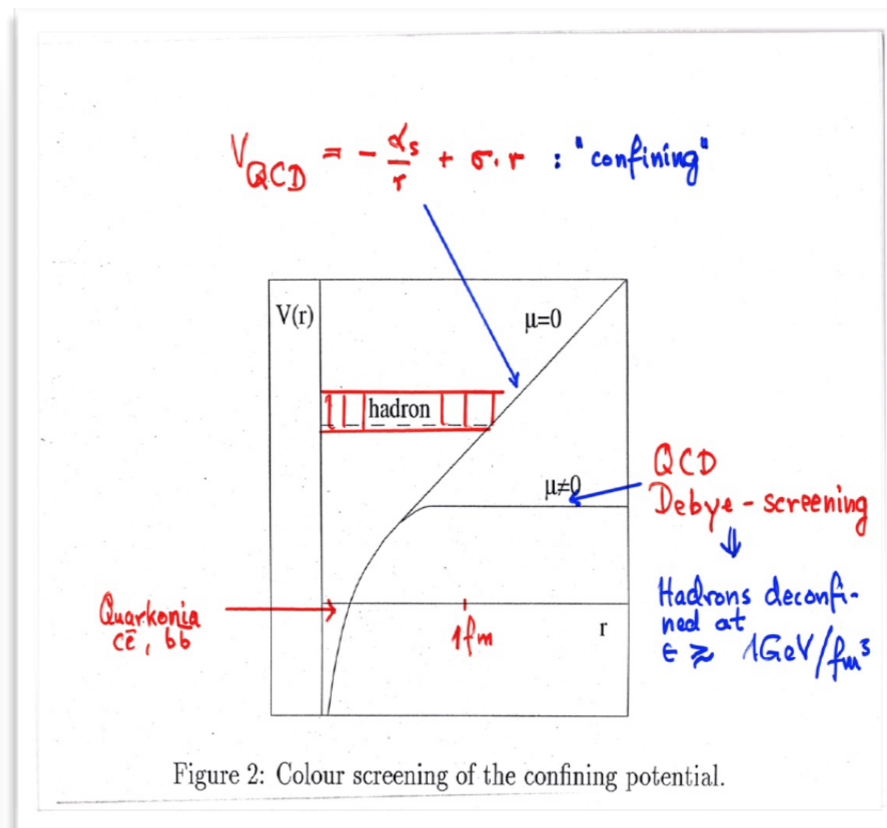


Note: variables here are (T, ρ_B) .
Replaced later by (T, μ_B) in the
Hagedorn-inspired Statistical
Hadronization Model

Fig. 1. Schematic phase diagram of hadronic matter. ρ_B is the density of baryonic number. Quarks are confined in phase I and unconfined in phase II.

First view of deconfinement in relativistic A+A collisions

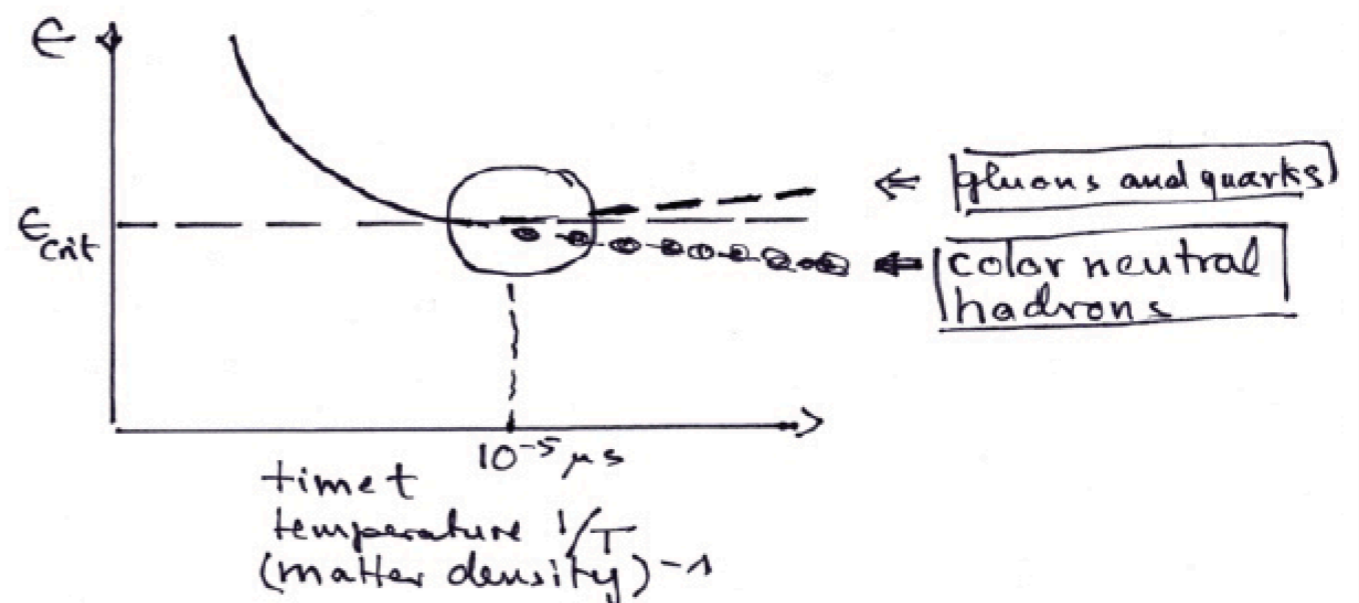
- about 1980 T.D.Lee: „Distributing enough energy density in space to melt the physical vacuum“
- „Physical Vacuum“ \equiv Color fields confined in hadrons
- after „melting“ \equiv Color fields permeate all space: QCD Plasma
- Estimate: $\epsilon(c)$ about 1-2 GeV/fm³ in the „fireball“



Hadronization: a phase transition in the cosmic expansion

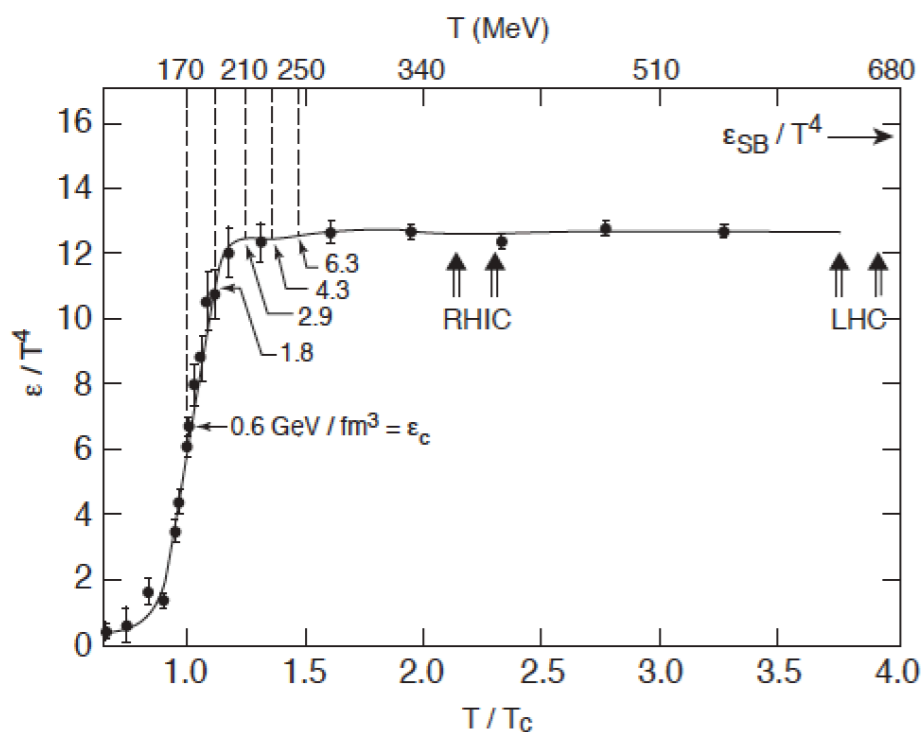
at $t \approx 10^{-5} s$ where $T \approx 2 \cdot 10^{12} K$ and $R \approx 10^{11} km$
 Matter/energy density $\epsilon = 1 GeV/fm^3 \hat{=} 2 \cdot 10^{18} kg/m^3$

The expanding {quark+gluon} matter hits a "critical point", at $\epsilon \approx 1 GeV/fm^3, T = 170 MeV$



Entry of Lattice QCD

Lattice QCD with three flavors(2000)



F. Karsch, S. Hands 2001

- Energy density in Lattice QCD, year 2000
- At $3 \text{ GeV}/\text{fm}^3$ we are at $T = 210 \text{ MeV}$
- Note: RHIC and LHC estimates!
- Note: $\epsilon \approx 0.6 \text{ GeV}/\text{fm}^3$ at $T_c \approx 170 \text{ MeV}$ as it is in the Stat. Hadronization Model

Experiment Proposals!

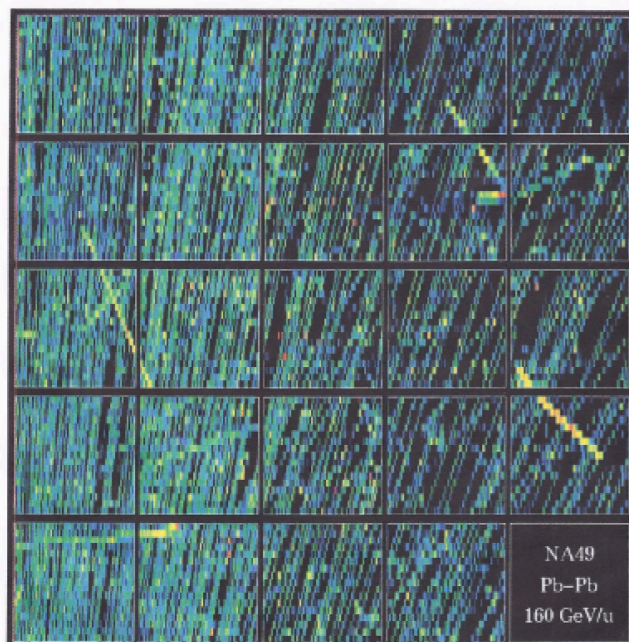
Upgrades:

- $^4\text{He} + ^4\text{He}$ @ CERN ISR $\sqrt{s} = 30 \text{ GeV}$
- $^{16}\text{O}, ^{32}\text{S} \dots ^{208}\text{Pb}$ @ AGS, SPS $\sqrt{s} \rightarrow 20 \text{ GeV}$

New Constructions:

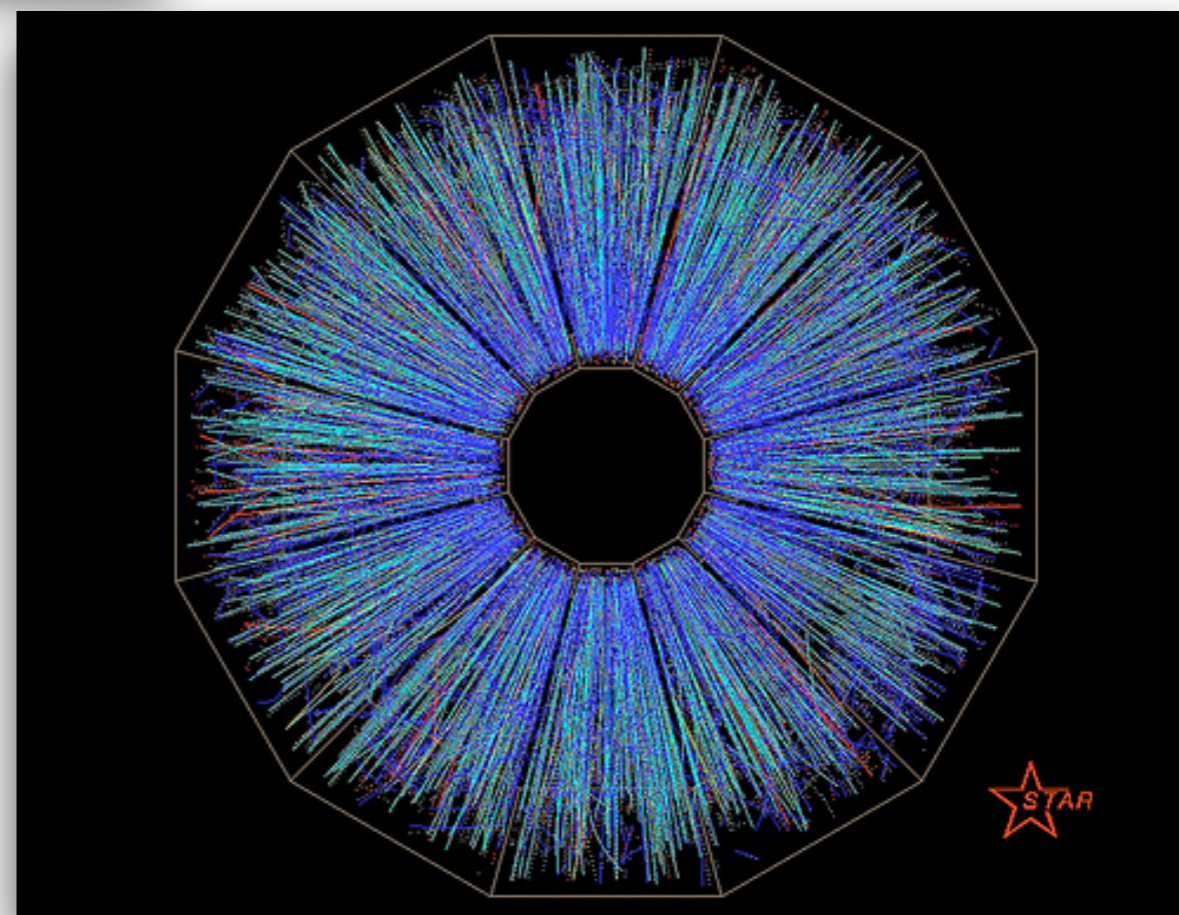
- RHIC @ BNL and LHC @ CERN

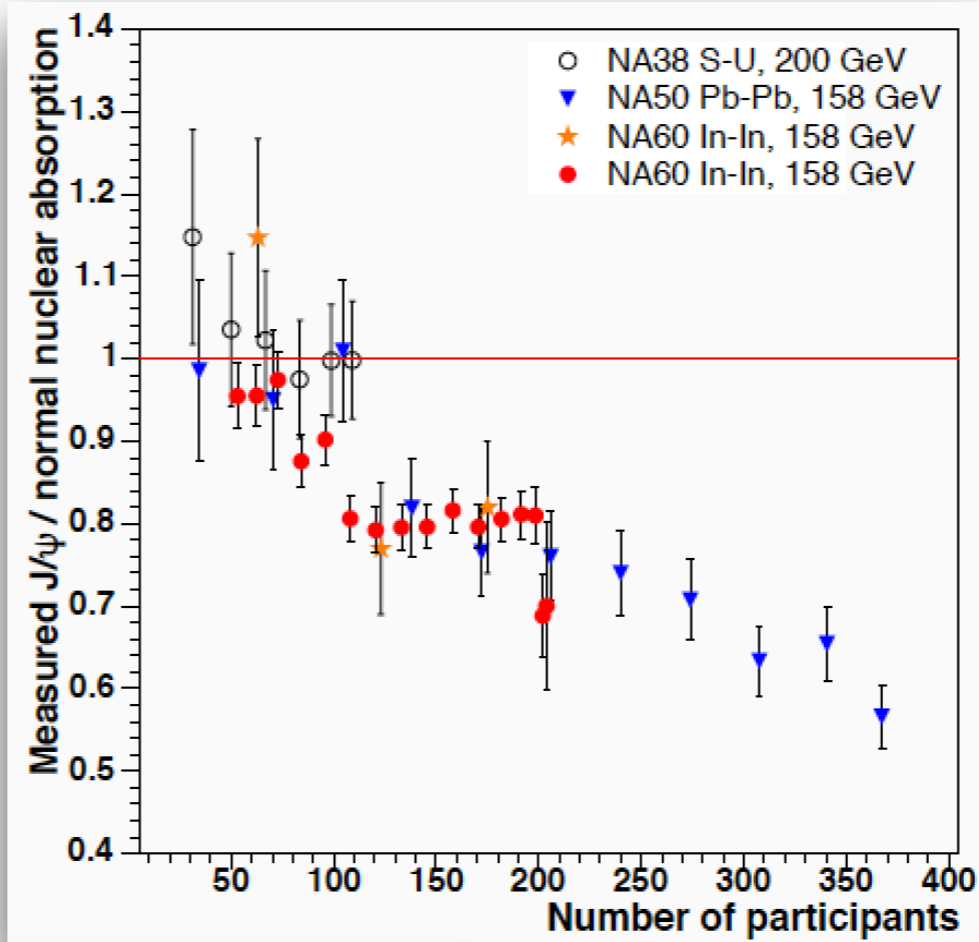
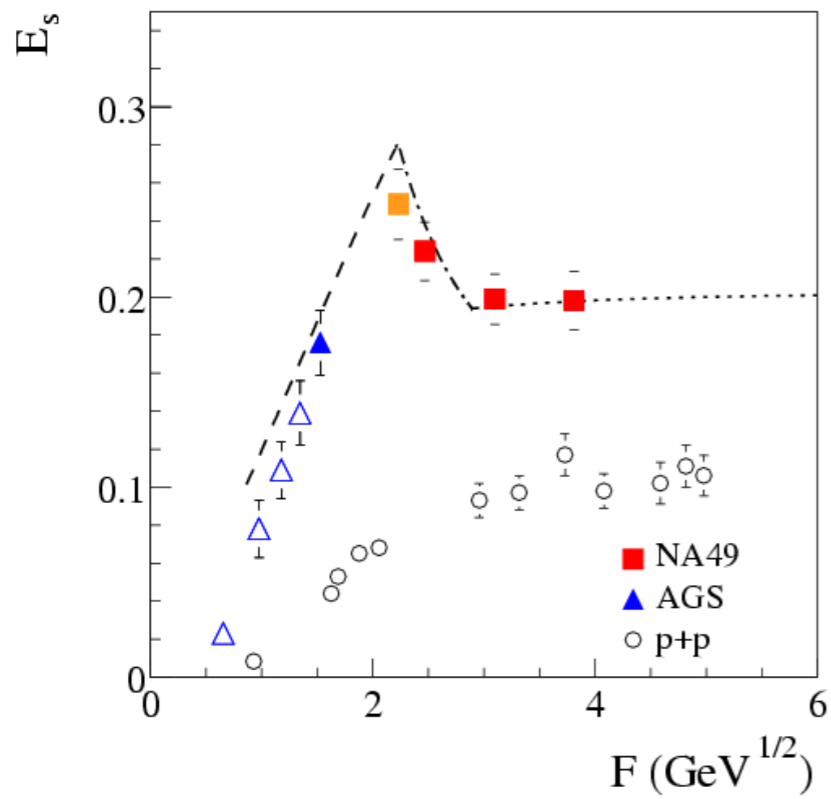
TPC Tracking 1996



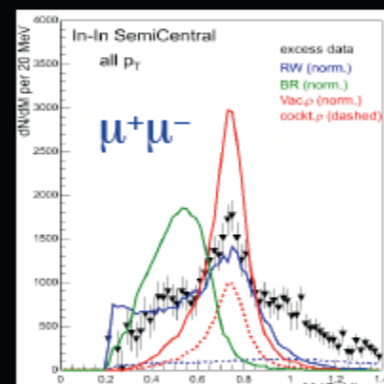
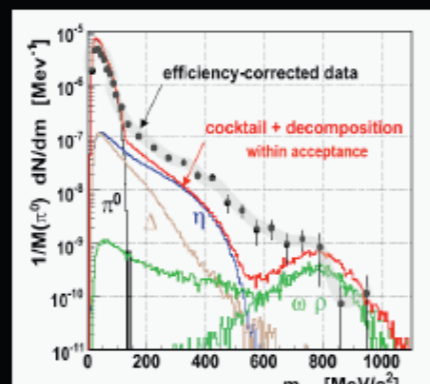
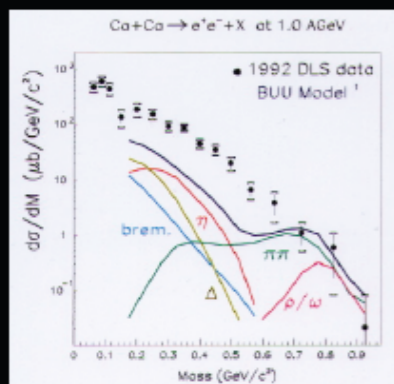
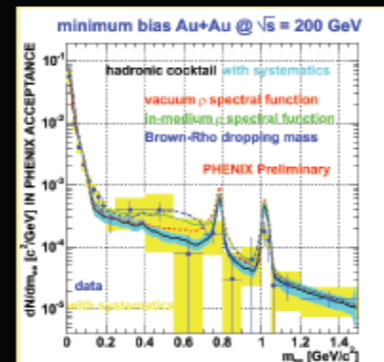
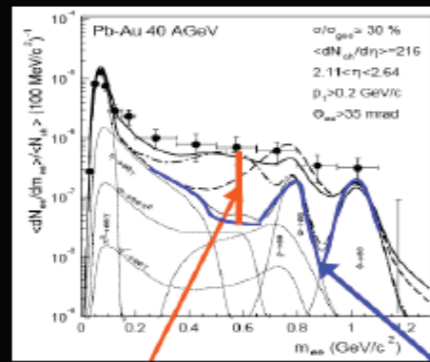
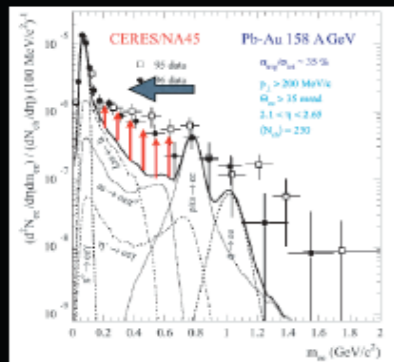
- Total charged particle multiplicity about 1000
- inner segment of TPC drift volume

Data rate from NA49 to ALICE: Factor > 1000 !



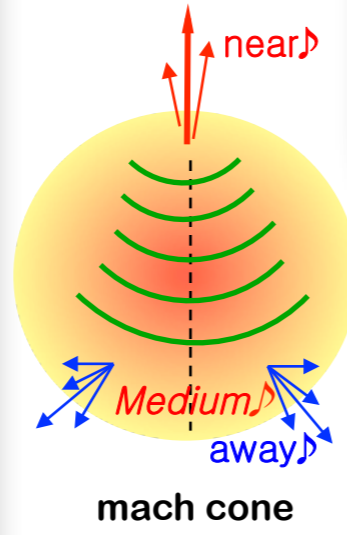
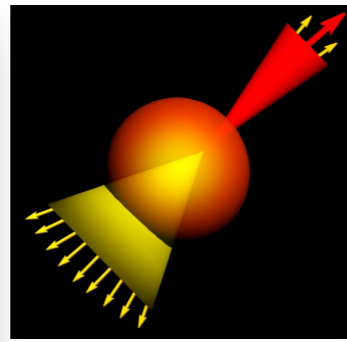
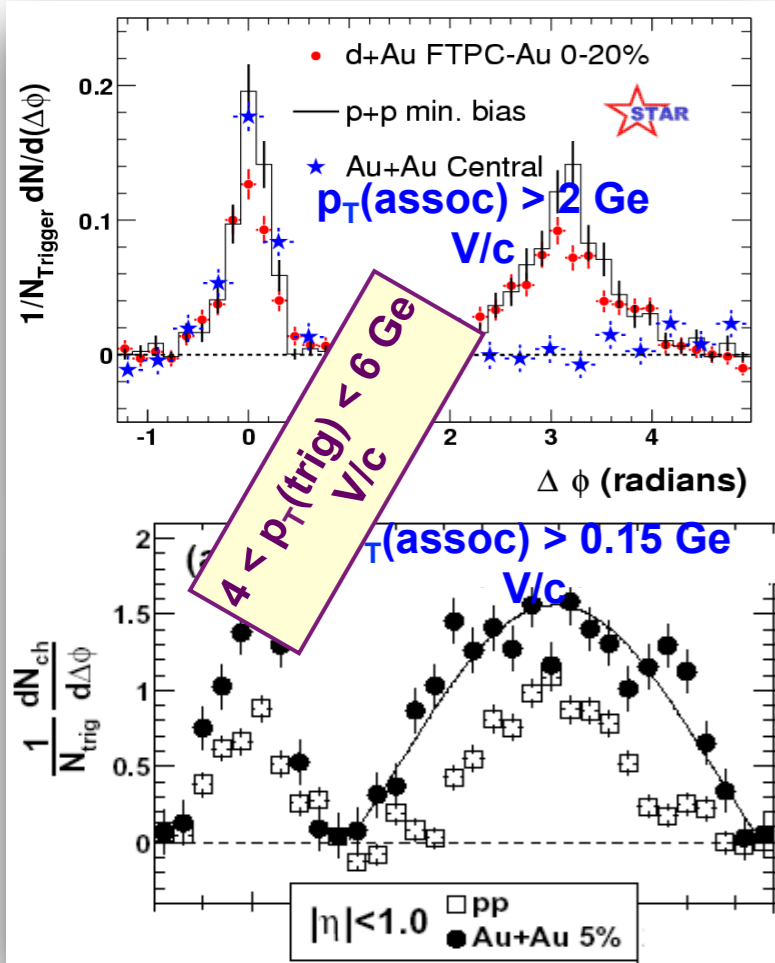


HI results (see V. Metag's talk)



Clear excess of dileptons observed. NA60: $\Gamma \nearrow$, no ΔM

SPS-Era



The RHIC Revolution

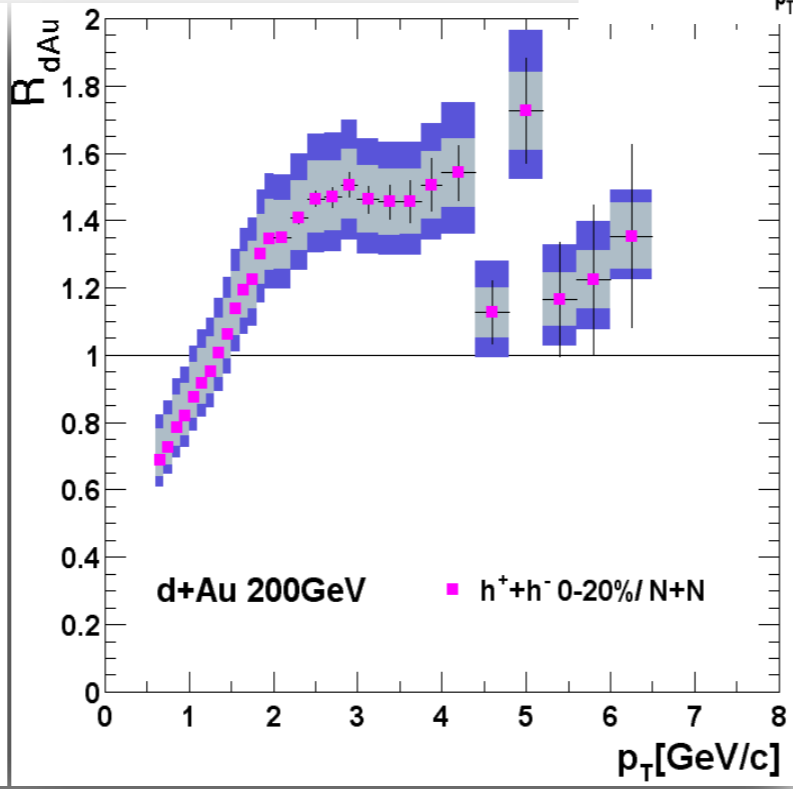
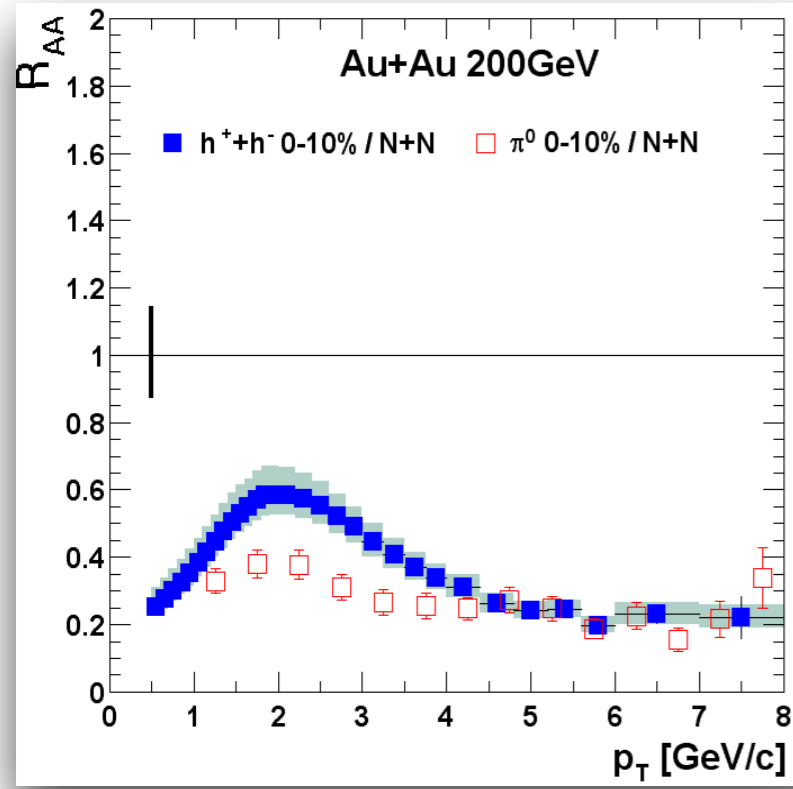
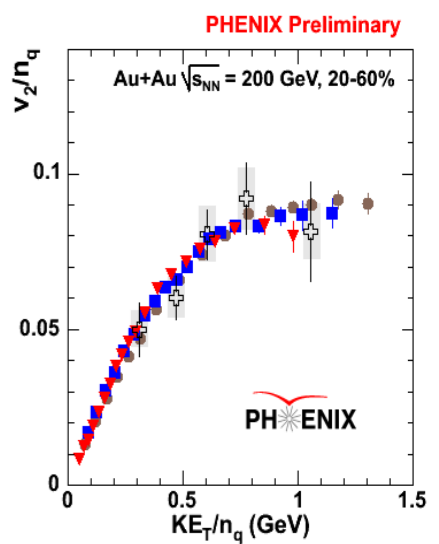
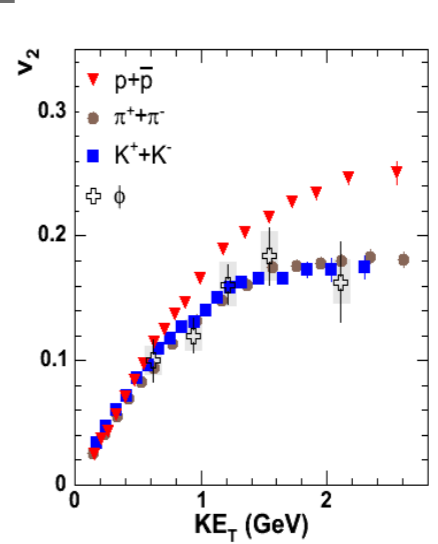
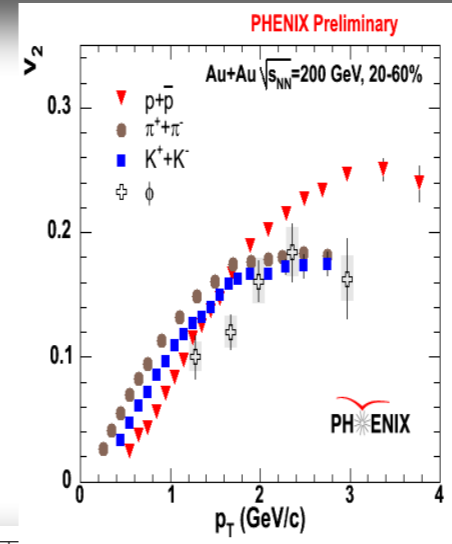
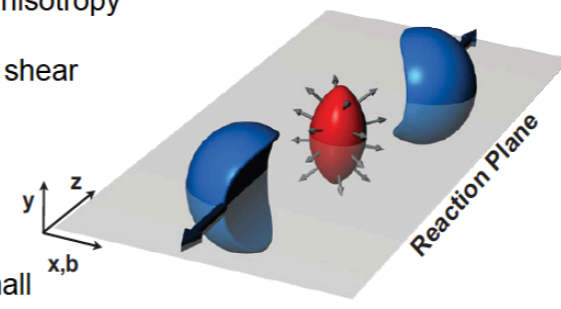
Elliptic Flow v_2 generated by primordial source anisotropy

Sensitive to scaled shear viscosity η/s

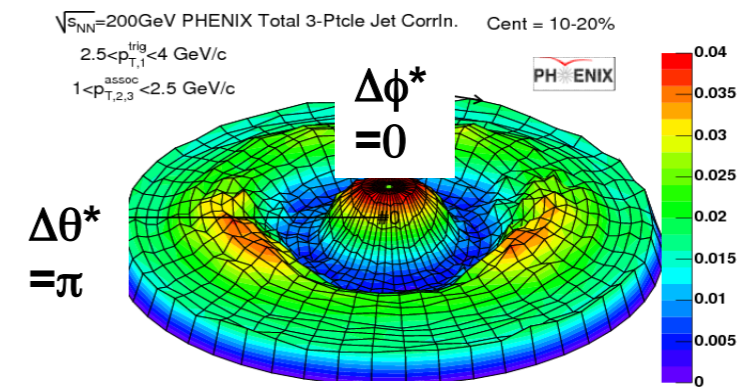
Viscous relativistic hydrodynamics

BUT: η/s is very small

QGP is a nearly ideal fluid

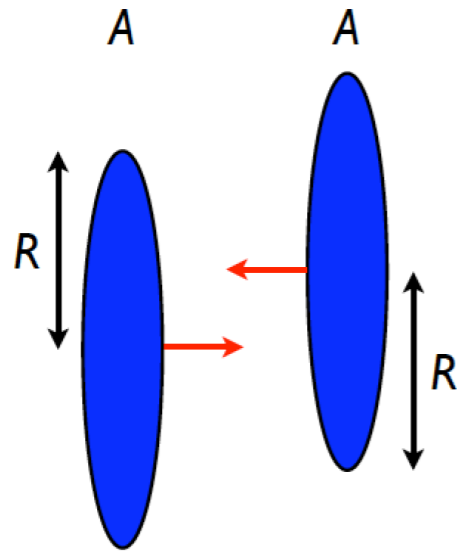


RHIC-Era



PHENIX Preliminary

Early time resolution as a consequence of short interpenetration time



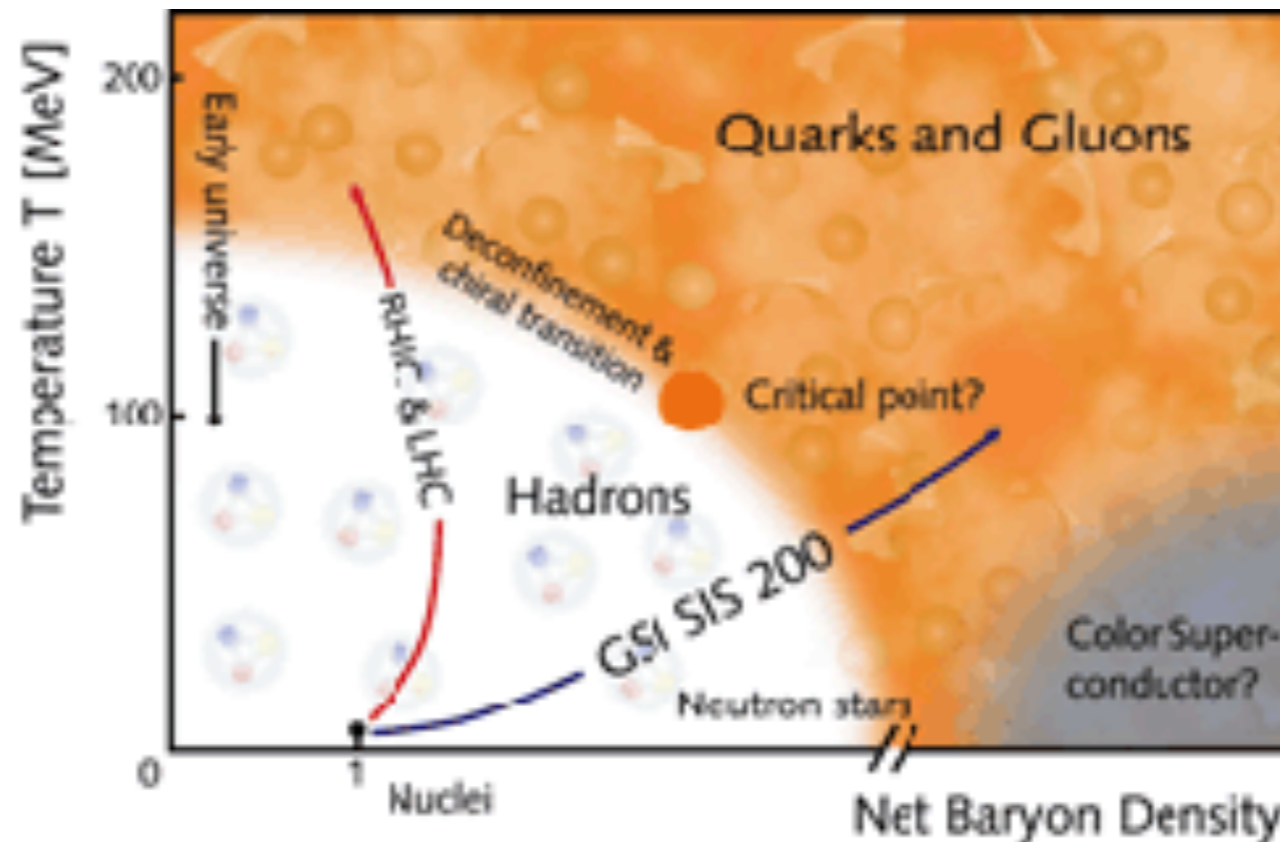
$$\tau = 2R/\gamma$$

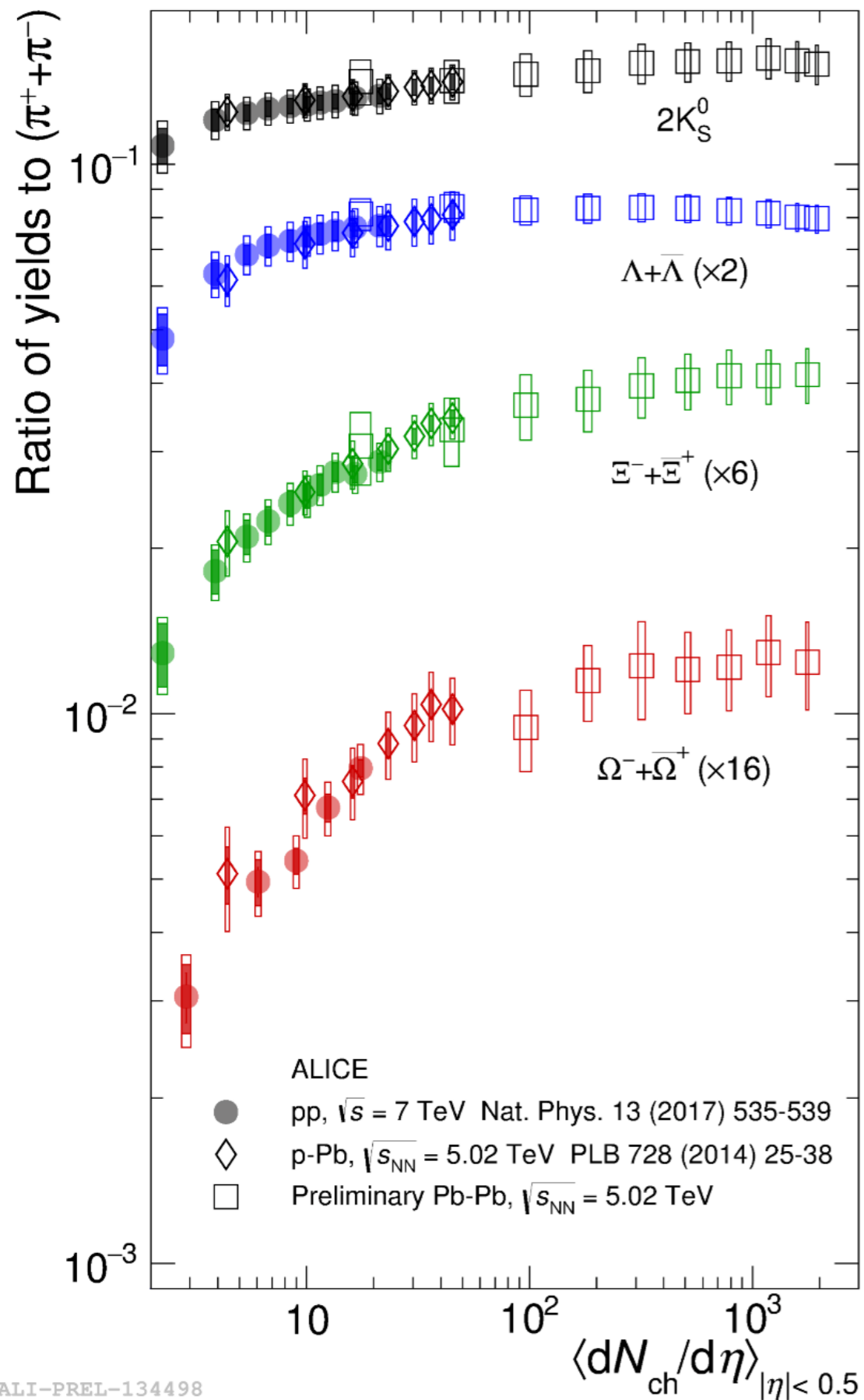
= 1.5 fm/c at top SPS energy
= 0.14 fm/c at top RHIC energy

= 0.06 fm/c at LHC

Hot matter study

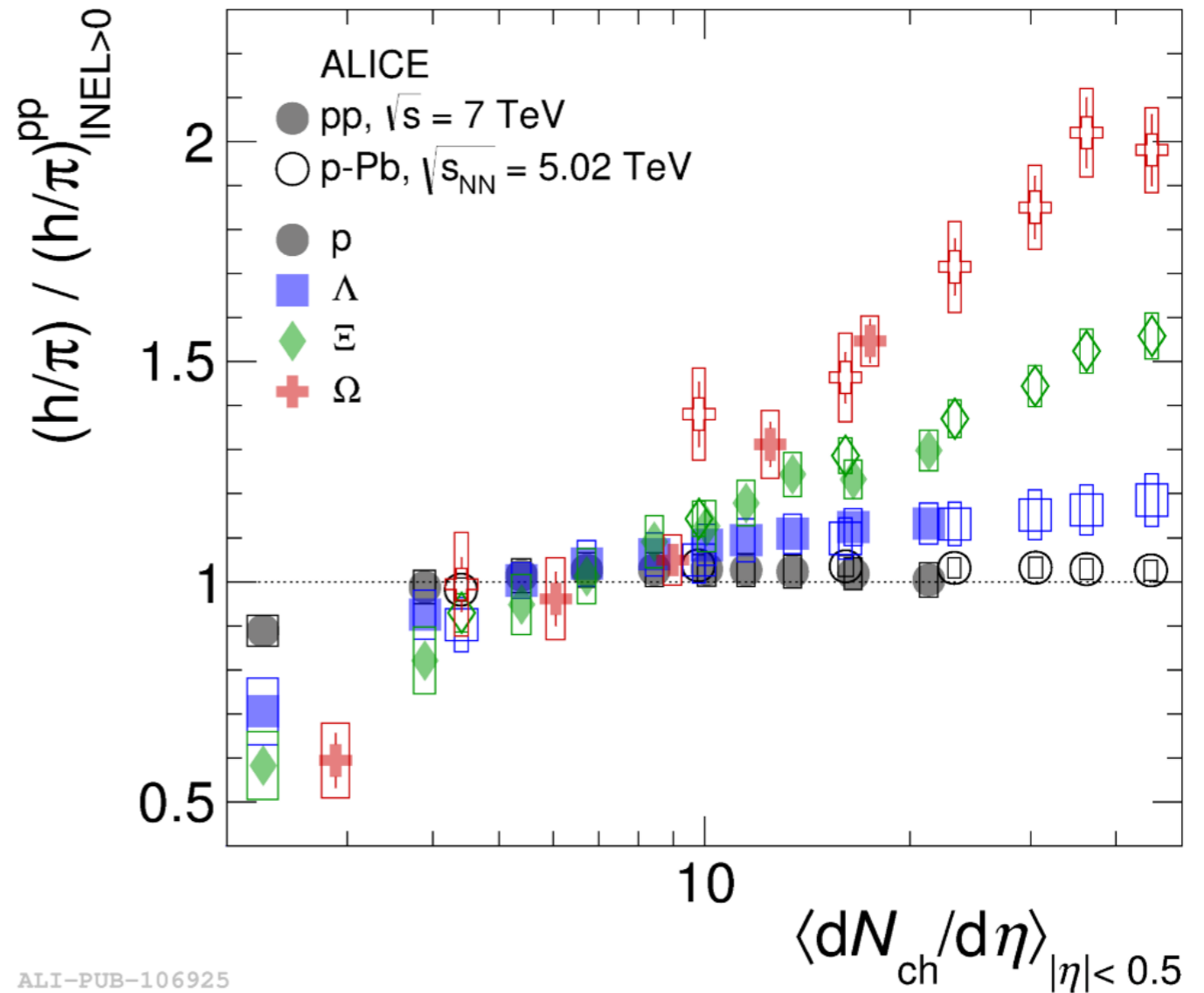
- strongly interacting hot matter
- Properties: viscosity
- how opaque?
- is it QGP?
- ➔ Hard probes: HF
- ➔ How about LF?



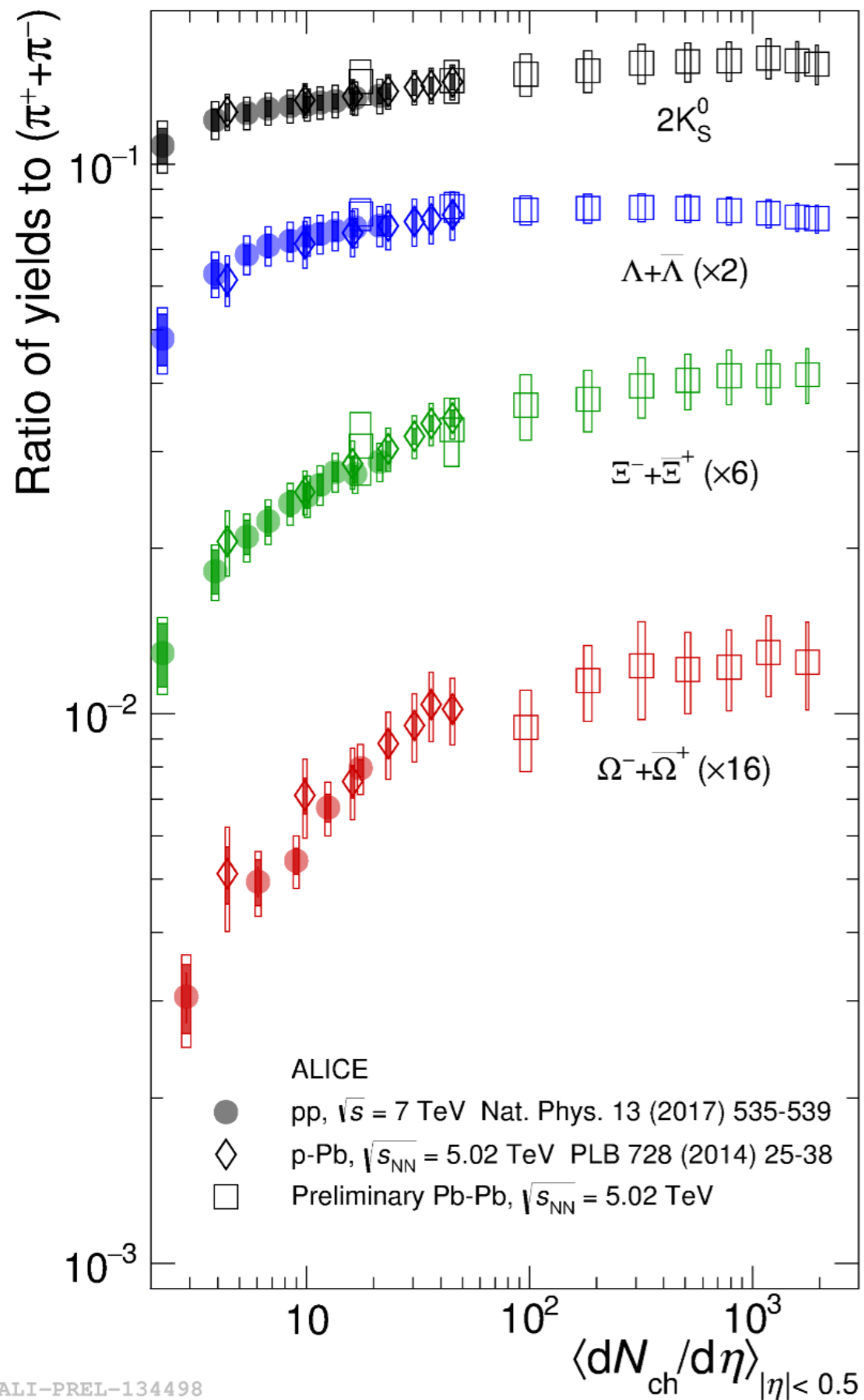


ALI-PREL-134498

Nat. Phys. 13 (2017)535-539

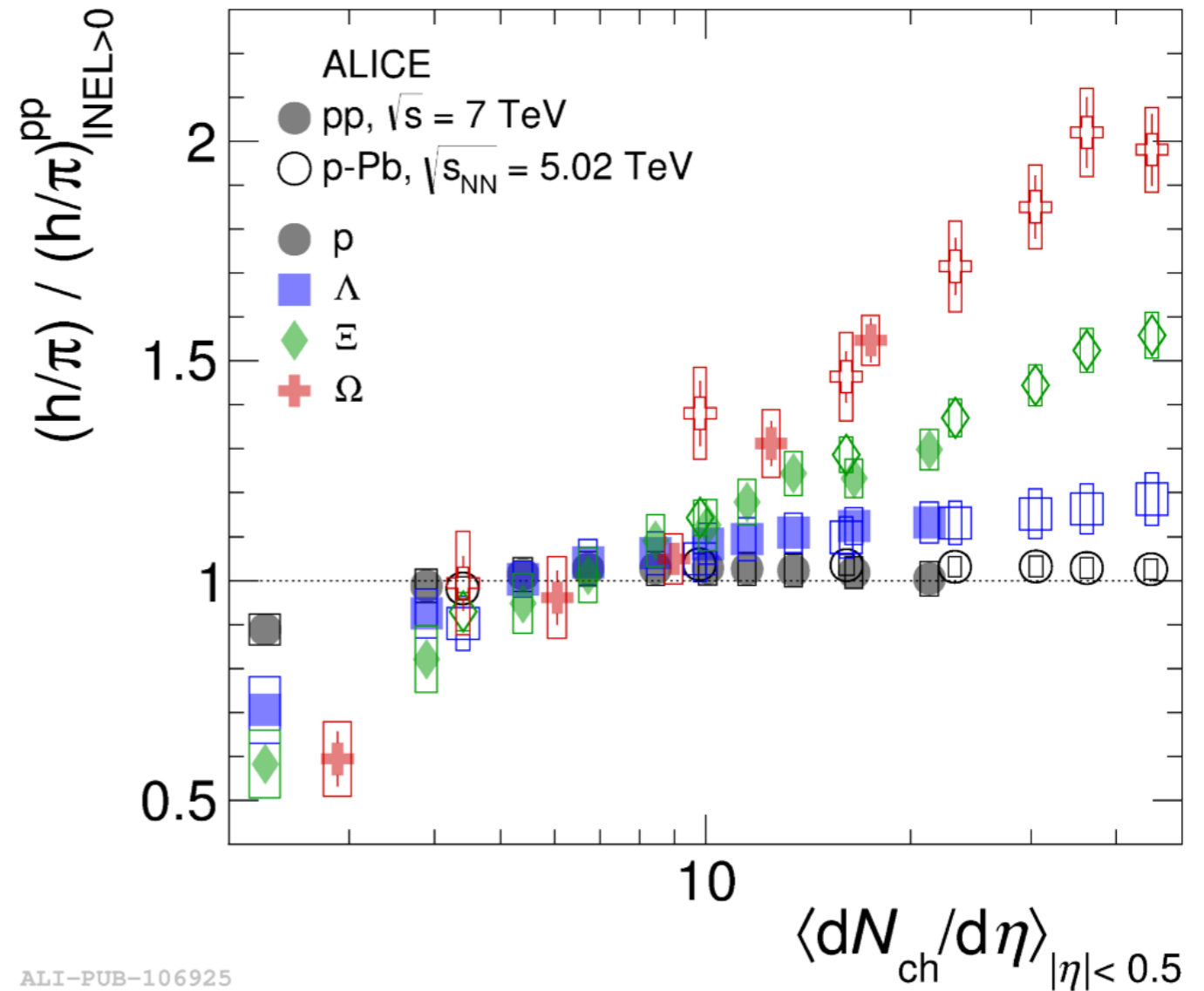


ALI-PUB-106925



ALI-PREL-134498

Nat. Phys. 13 (2017)535-539

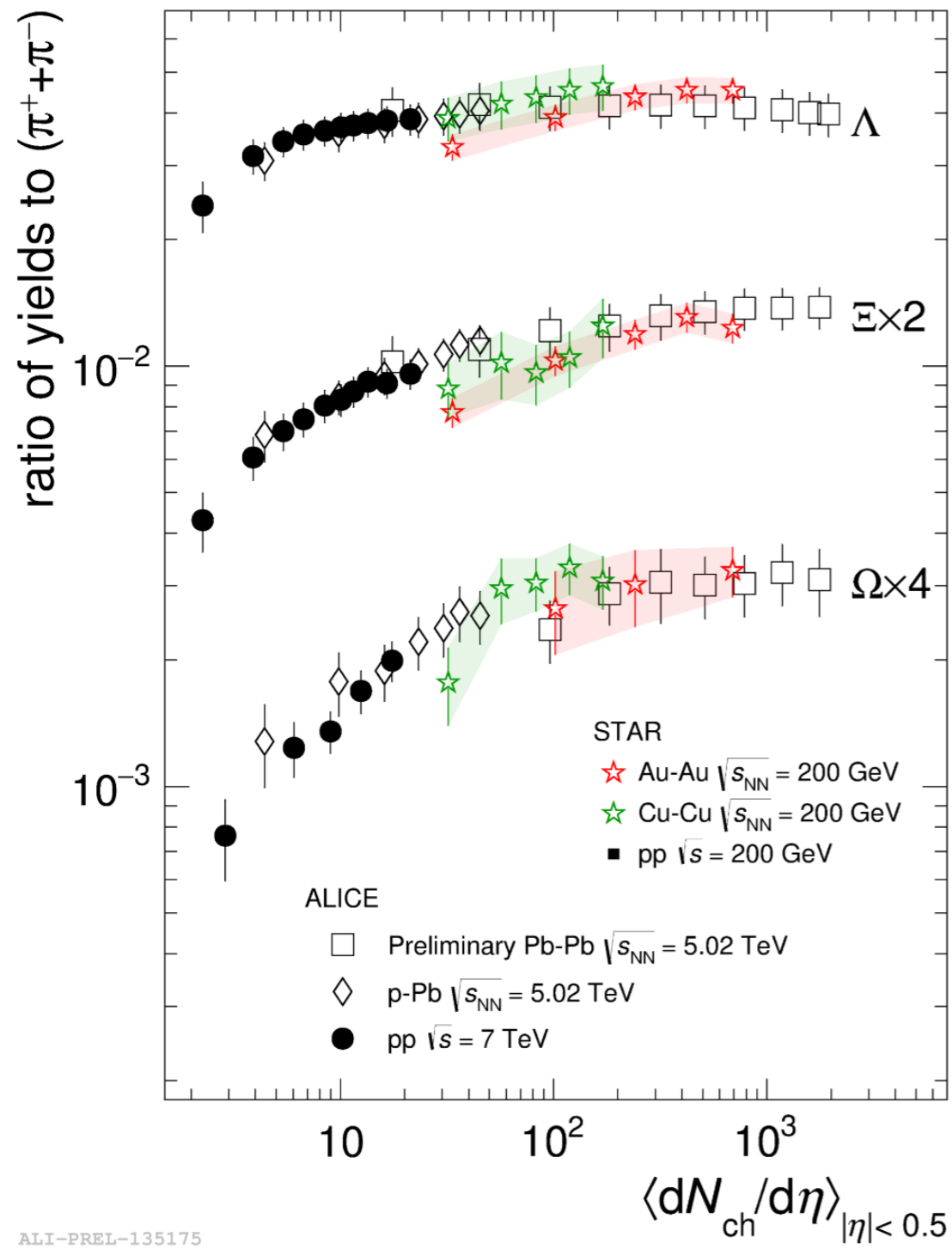


ALI-PUB-106925

Multiplicity = Universal Variable for the collisions

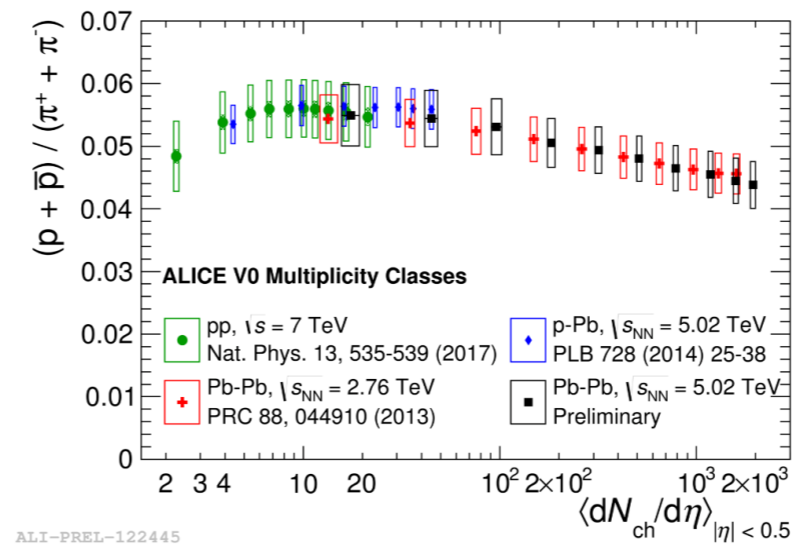
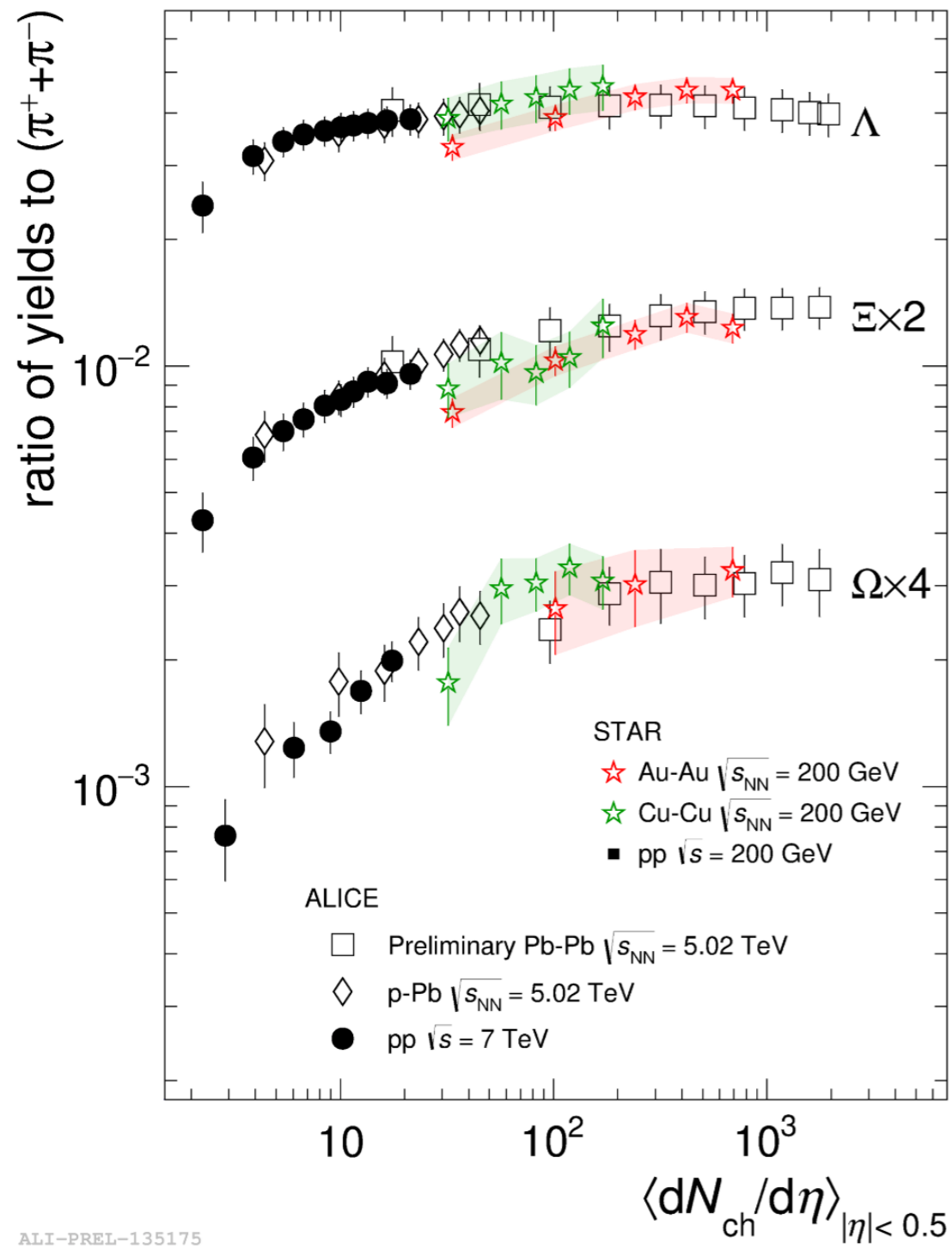
- Energy, System don't matter!

➡ Attention to the small system!

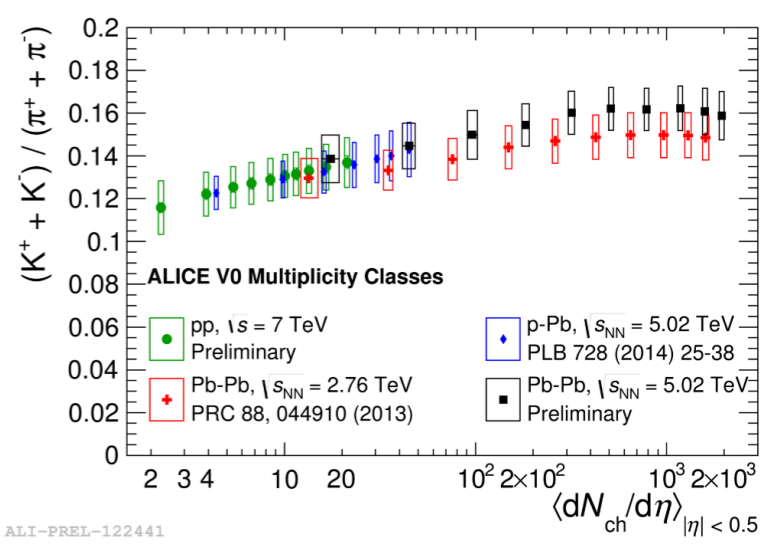
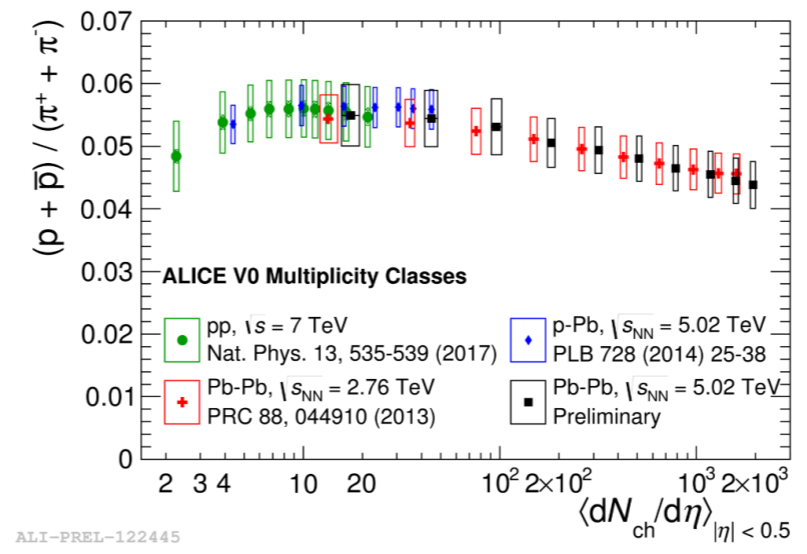
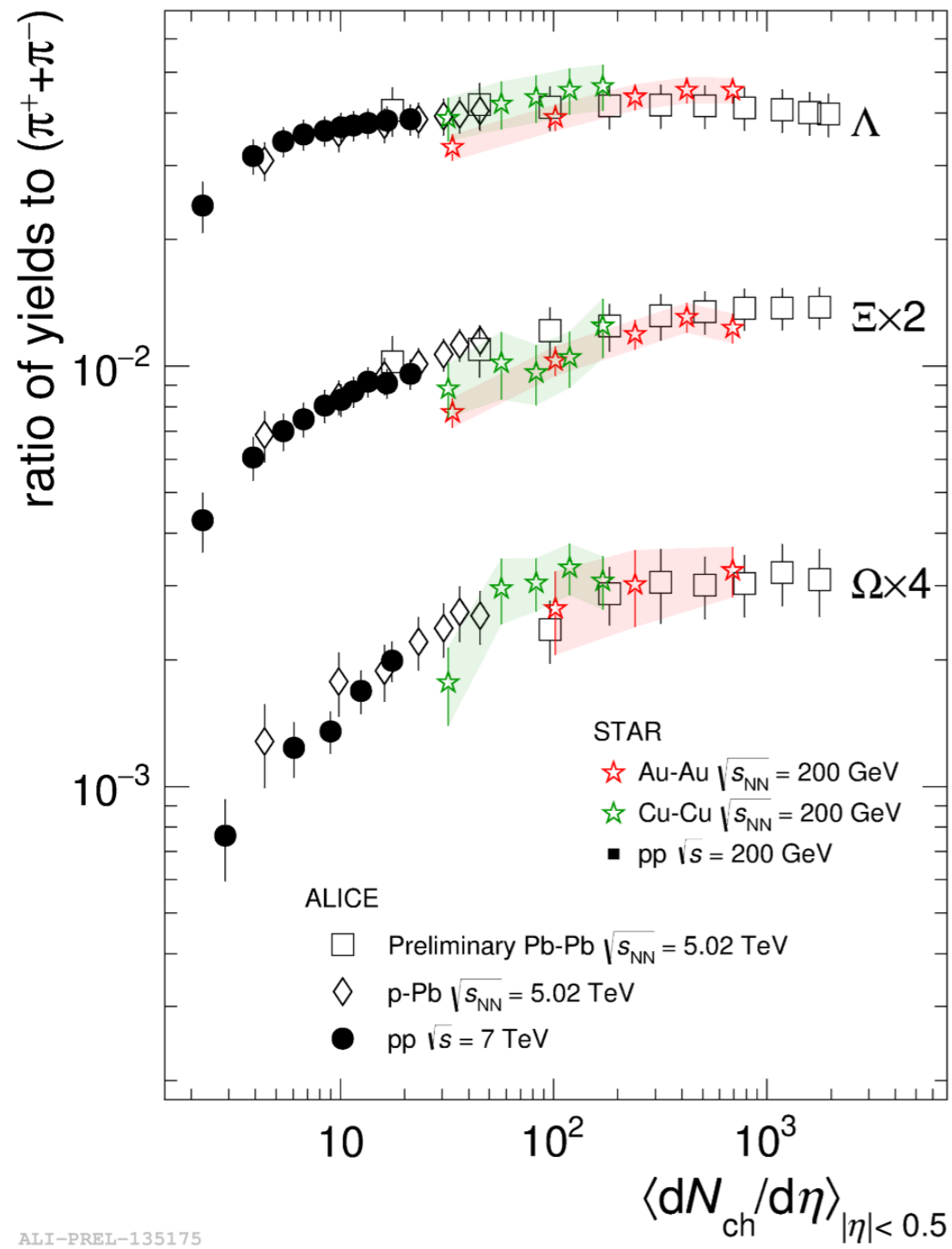


ALI-PREL-135175

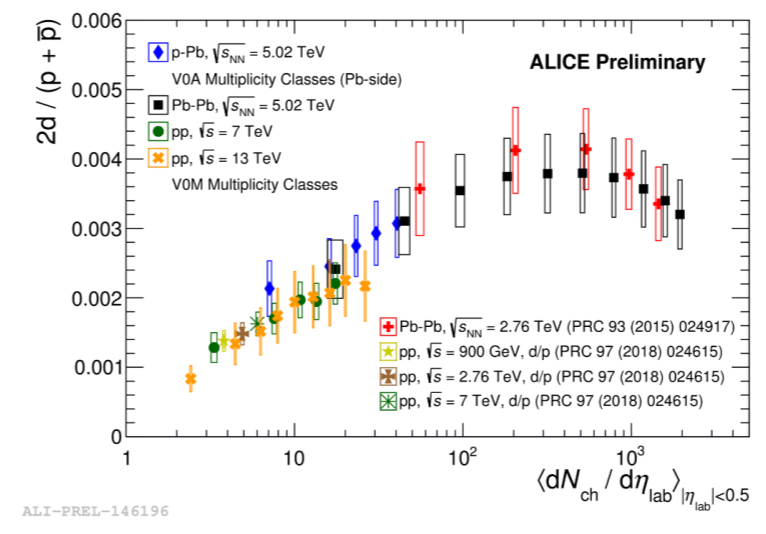
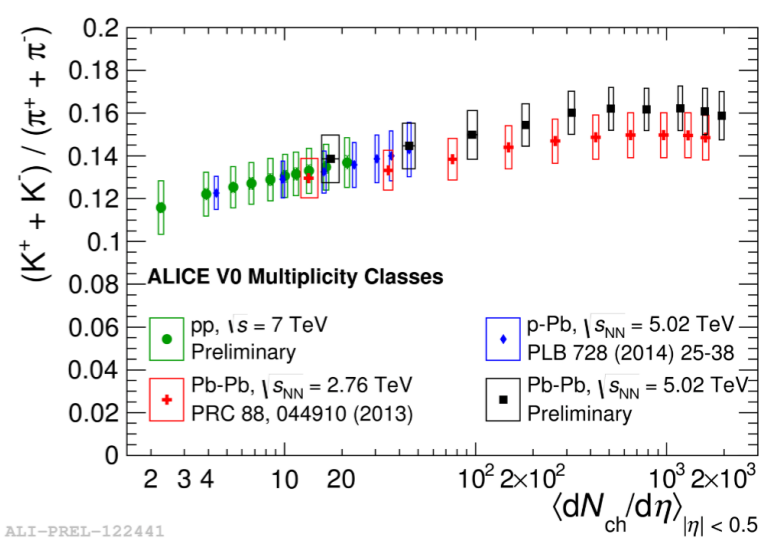
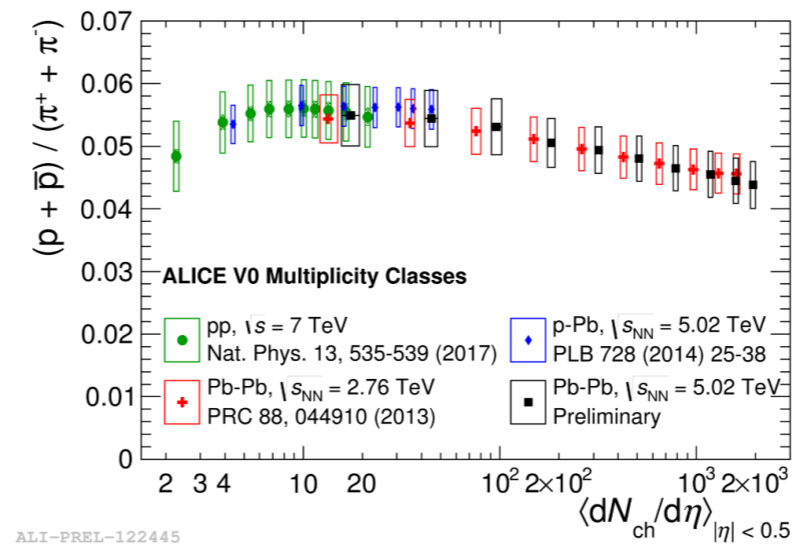
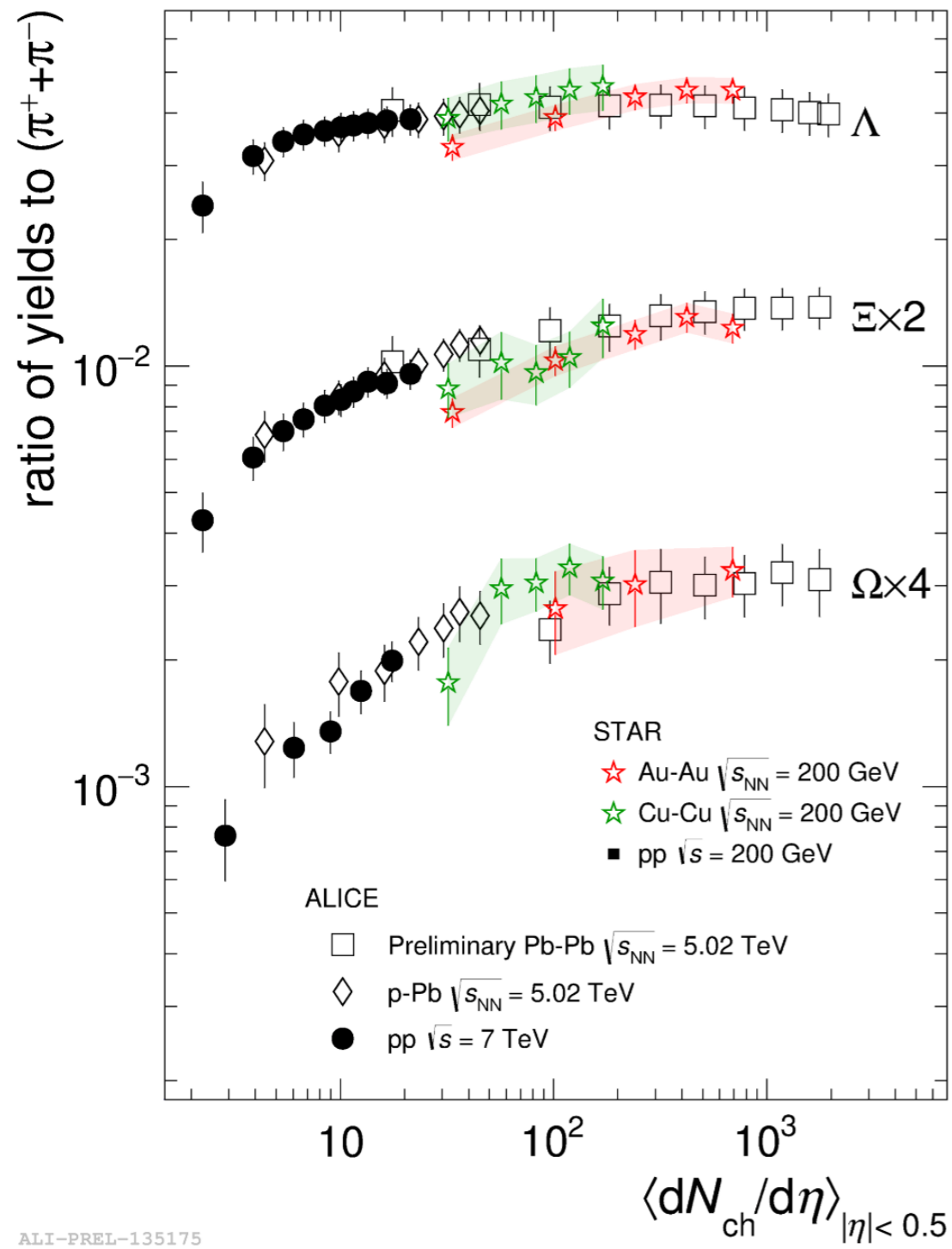
Even also for RHIC data!



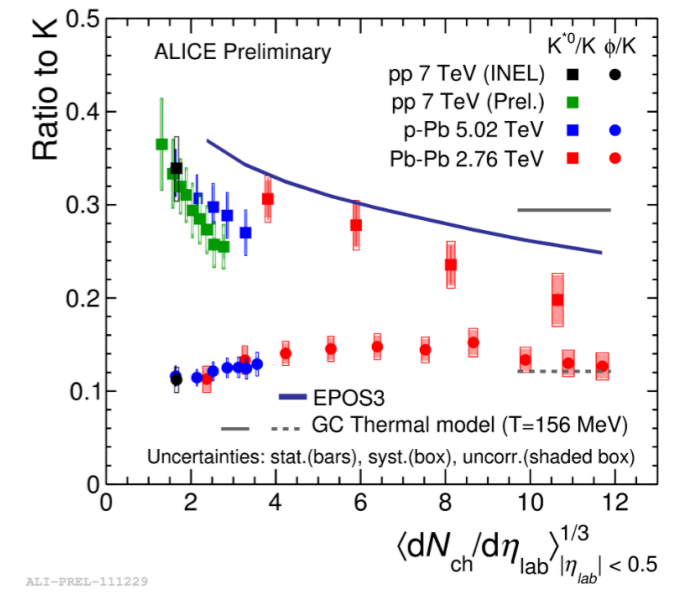
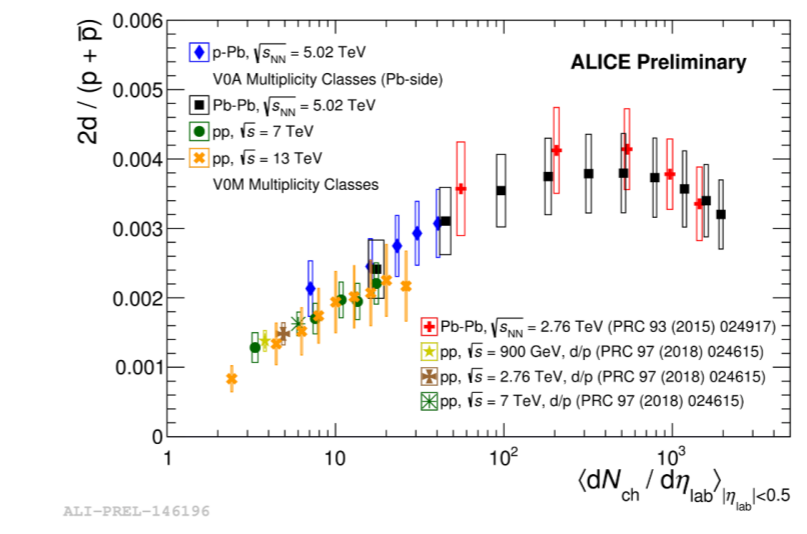
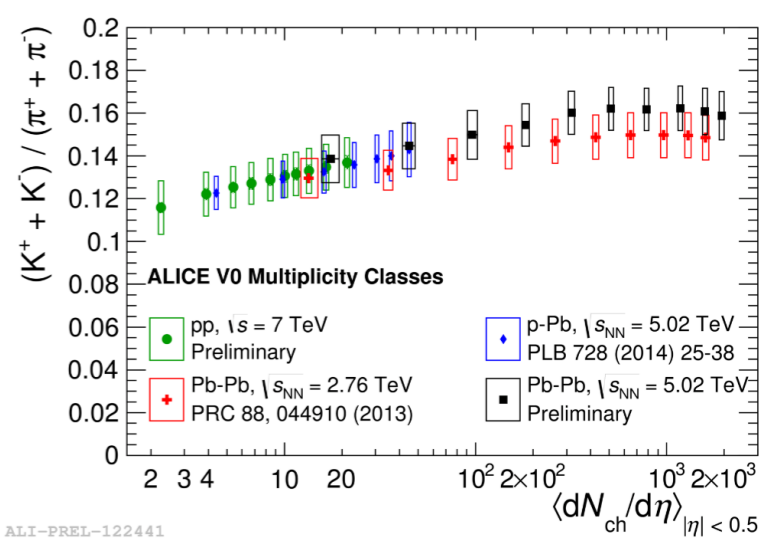
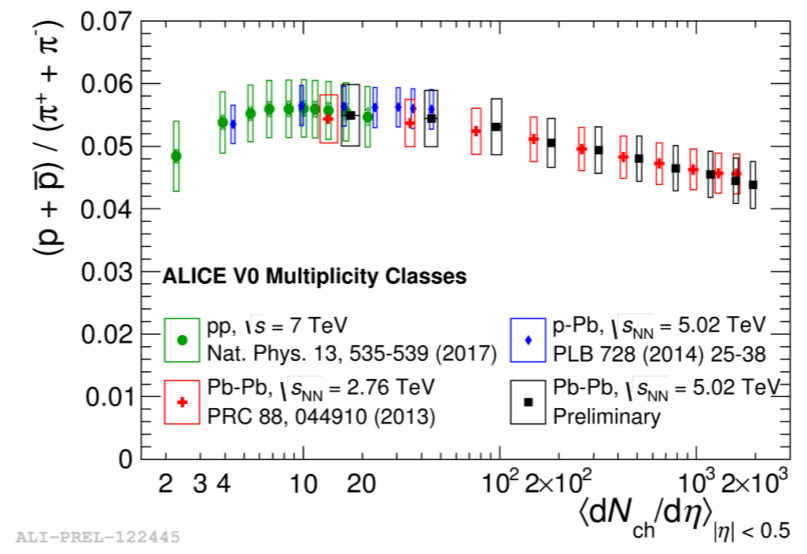
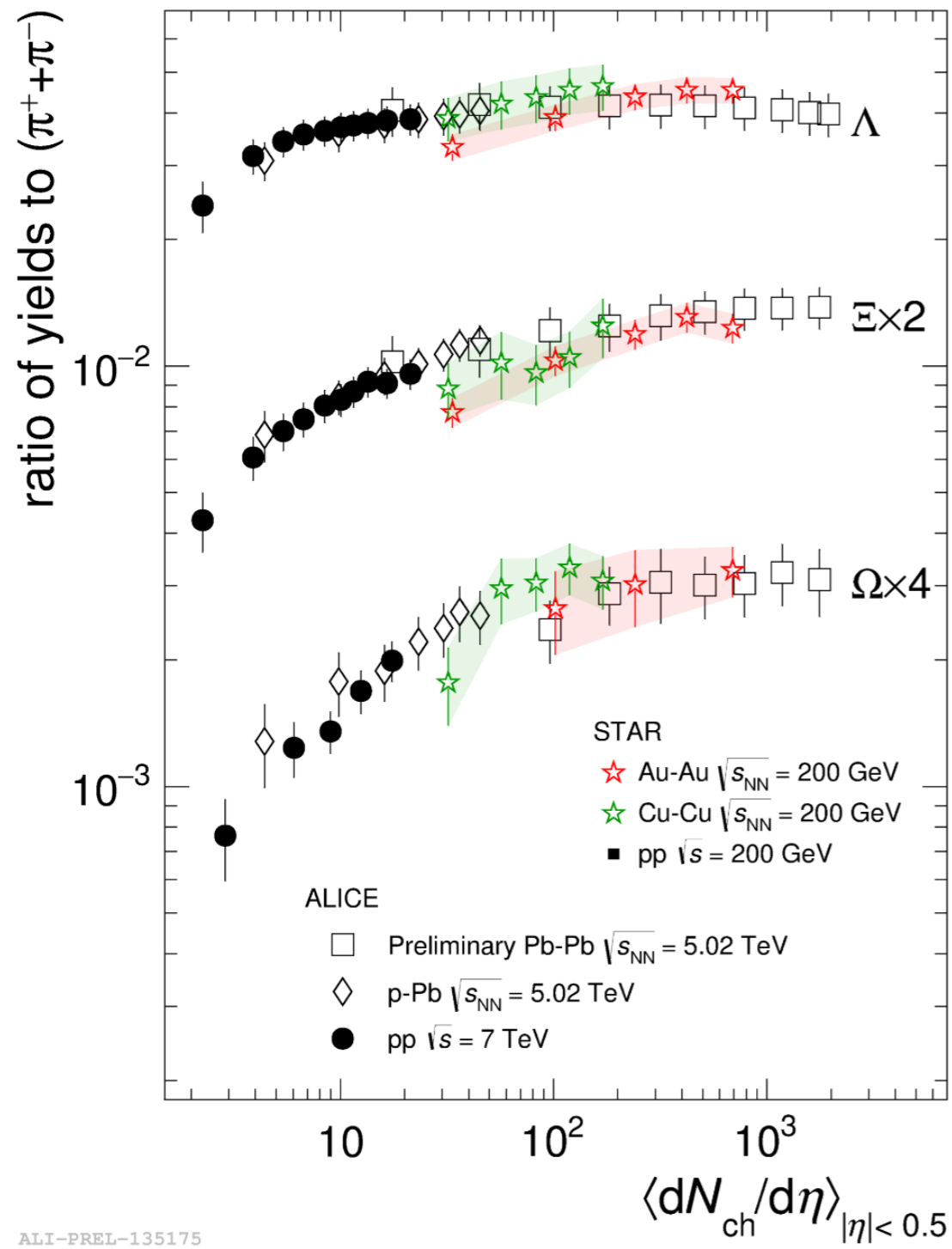
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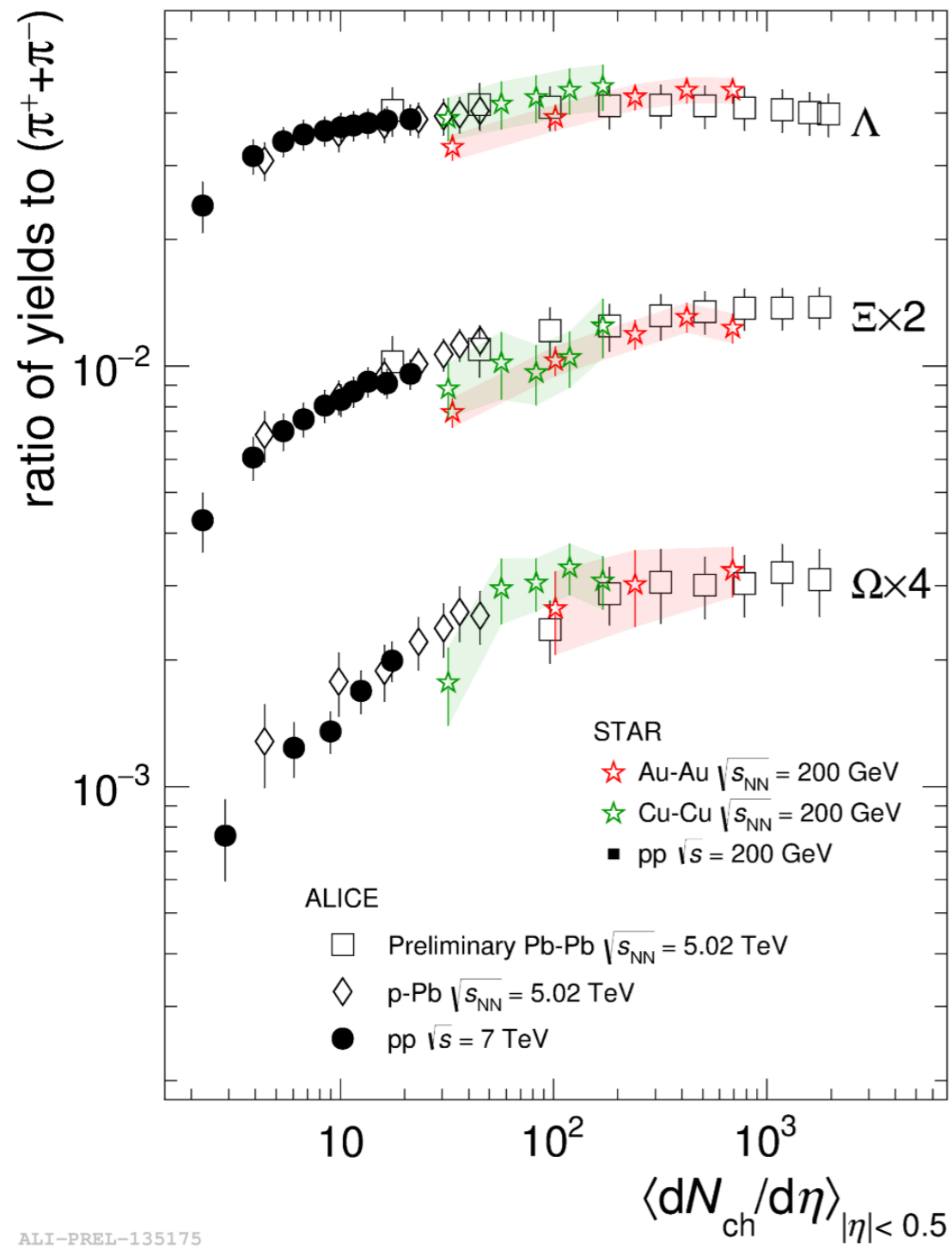
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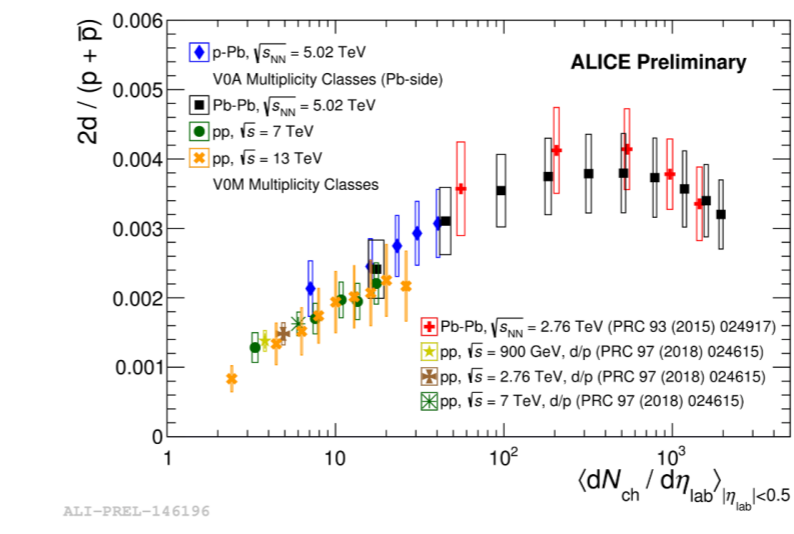
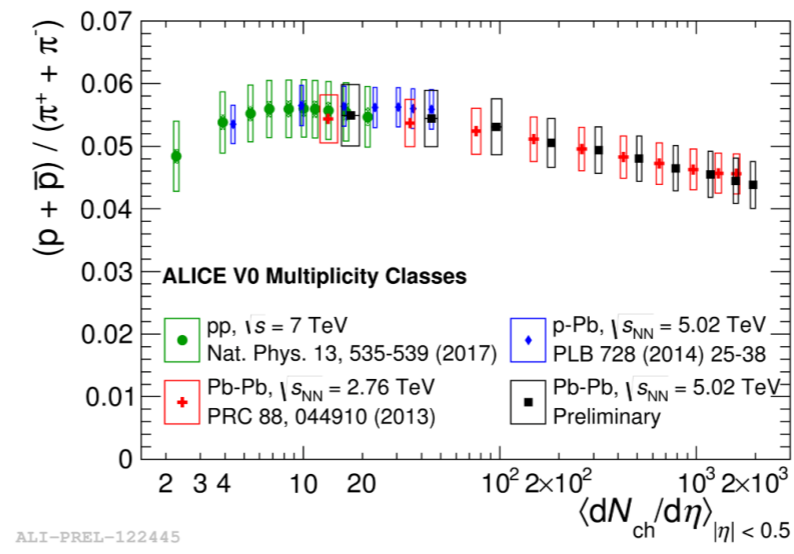
Even also for RHIC data!



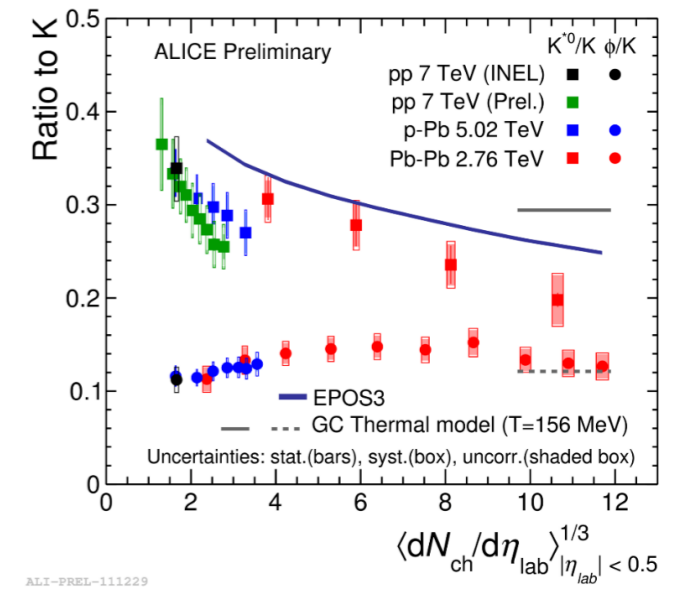
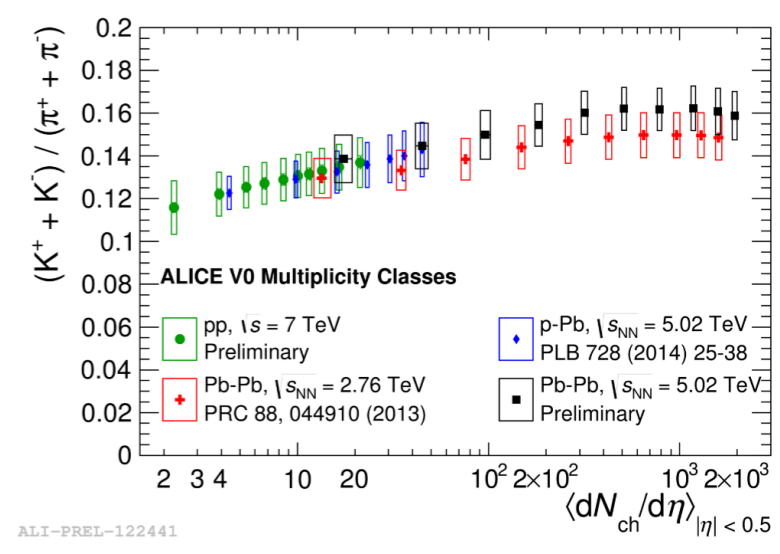
Even also for RHIC data!



Even also for RHIC data!



Even for all LFs, not only for Strangeness



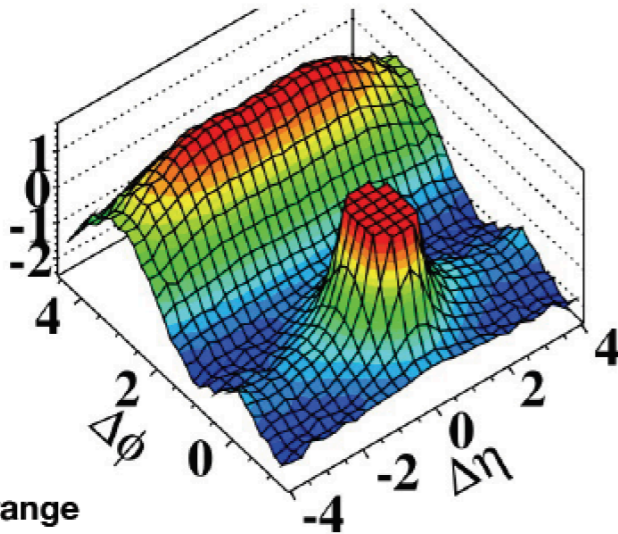
Ridge in High Multiplicity pp

CMS, JHEP 09 (2010) 091

(d) CMS $N \geq 110$, $1.0 \text{ GeV}/c < p_T < 3.0 \text{ GeV}/c$

Long-range
Away-side jet

$R(\Delta\eta, \Delta\phi)$



Long-range
Ridge

Short-range
Near-side Jet

Collectivity in small systems

~ year 2005: sQGP signature

~ year 2010: sQGP in pp as well?

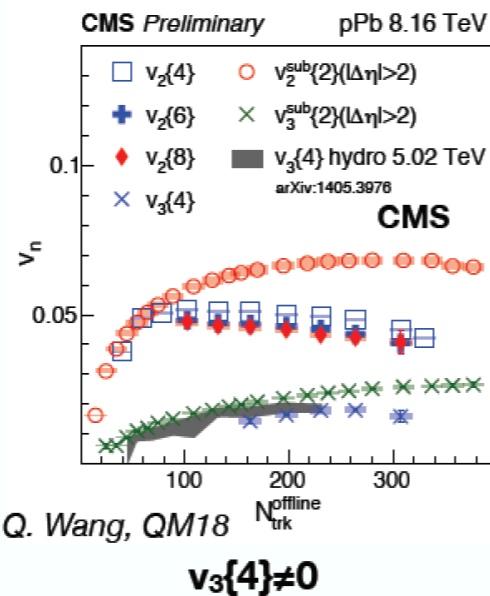
ridge includes v_2 + higher orders

Bounty Evidence of Collectivity at LHC

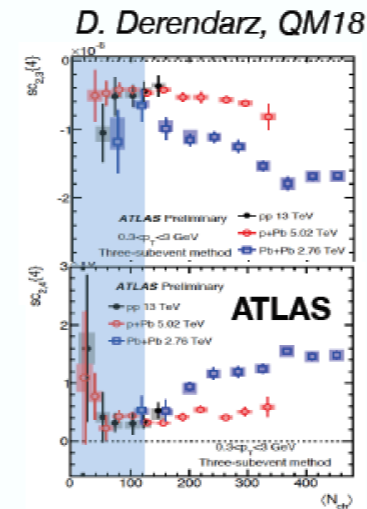
Selected few QM results

Li Yi's talk at QM2018

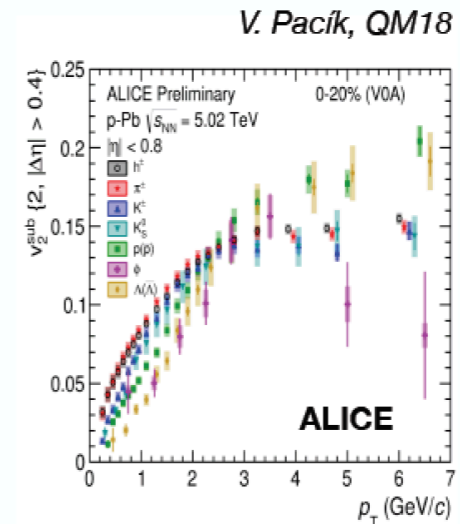
Multi-particle v_n



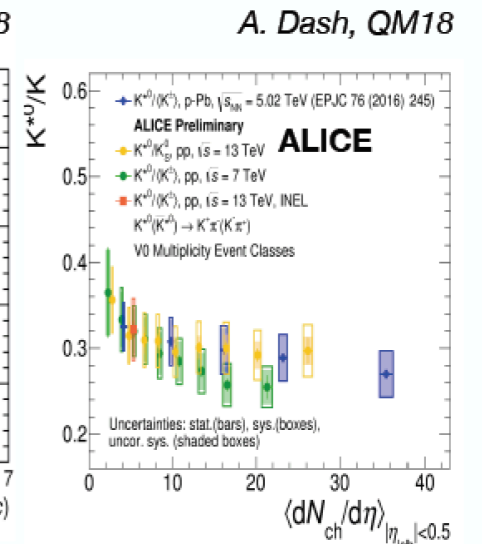
Sub-event cumulants



PID v_2



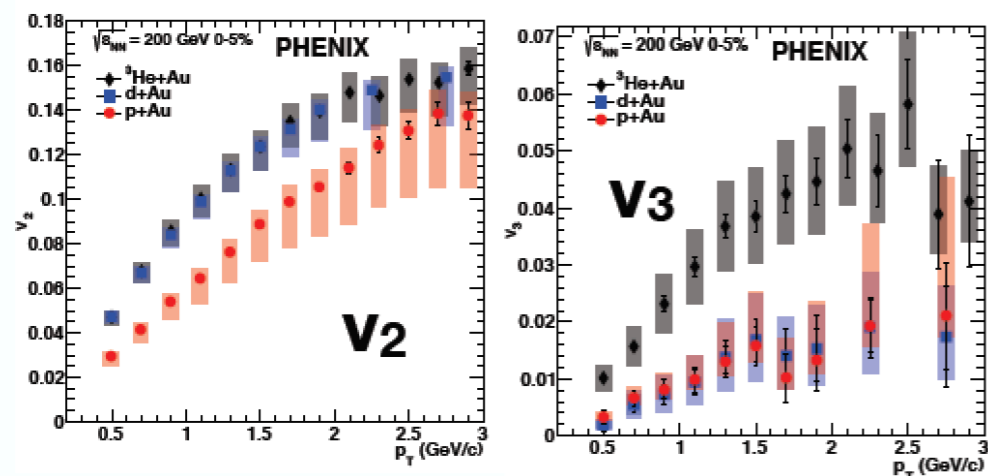
Resonance



Evidences of Collectivity at RHIC

Selected few QM results

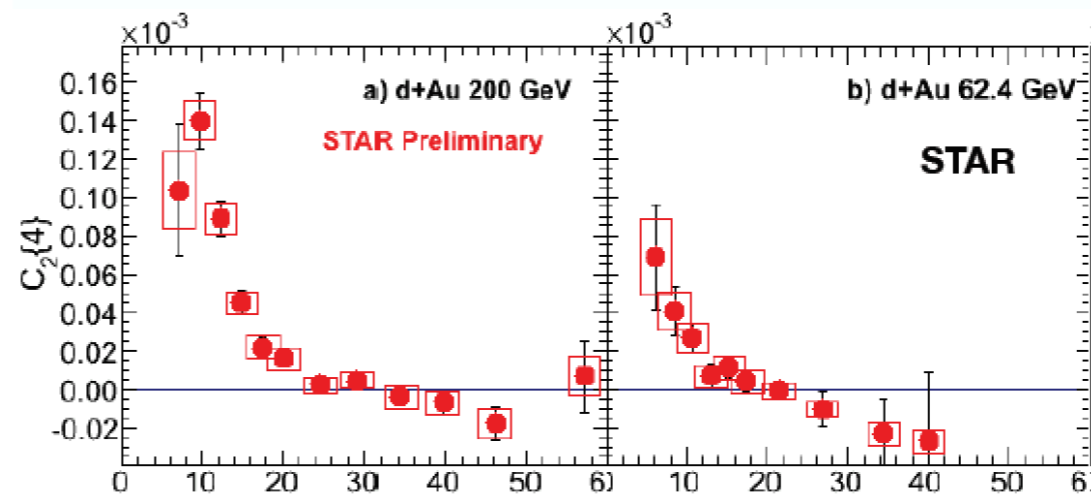
System shape scan



S. Morrow, QM18

v_n ordering, sensitive to initial geometry
Small QGP droplet?

4-particle cumulants

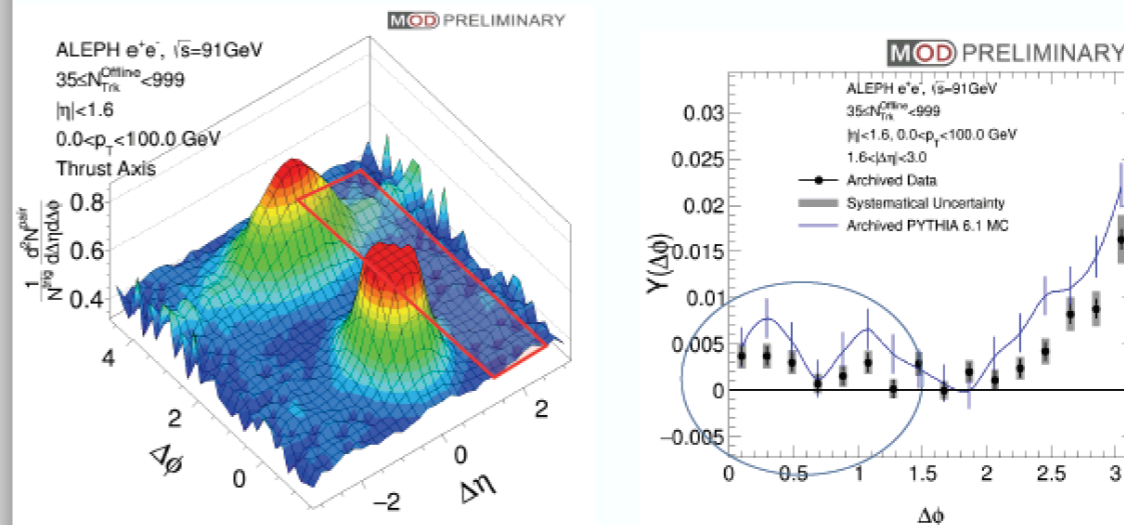


S. Huang, QM18

Negative $c_2\{4\}$

Even Smaller System

e^+e^- Y. Lee, QM18



Long-range correlation in e^+e^-

PYTHIA seems to be able to describe it

Where does Collectivity Come from?

Next talk! M.Strickland

Hydro

High density system evolution

Smallest QGP droplet?

Initial **spatial** anisotropy, driven by **pressure gradient**

Low density

AMPT, “escape model”; Kinetic transport; ..

L. He, et al, PLB 753, 506

A. Kurkela, QM18

Initial **spatial** anisotropy

...

CGC

Initial **momentum** anisotropy

C. Bierlich, QM18

B. Blok, QM18

...

MPI + CR, ..

How can we distinguish the models?

Where does Collectivity Come from?

Next talk! M.Strickland

Hydro

High density system evolution

Smallest QGP droplet?

Initial spatial anisotropy, driven by pressure gradient

Low density

AMPT, “escape model”; Kinetic transport; ..

L. He, et al, PLB 753, 506

A. Kurkela, QM18

Initial spatial anisotropy

...

CGC

Initial momentum anisotropy

C. Bierlich, QM18

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MPI + CR, ..

How can we distinguish the models?

- What is QG-P?
- How can we be sure for QGP?
- Why no quenching? Not observed, yet?

Where does Collectivity Come from?

Next talk! M.Strickland

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Li Yi (Shandong University)

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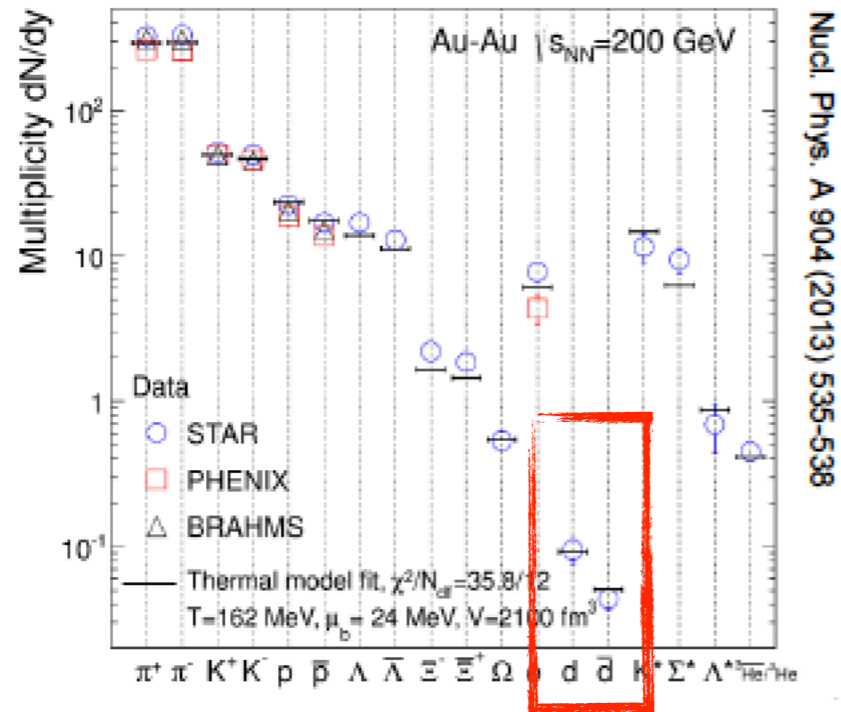
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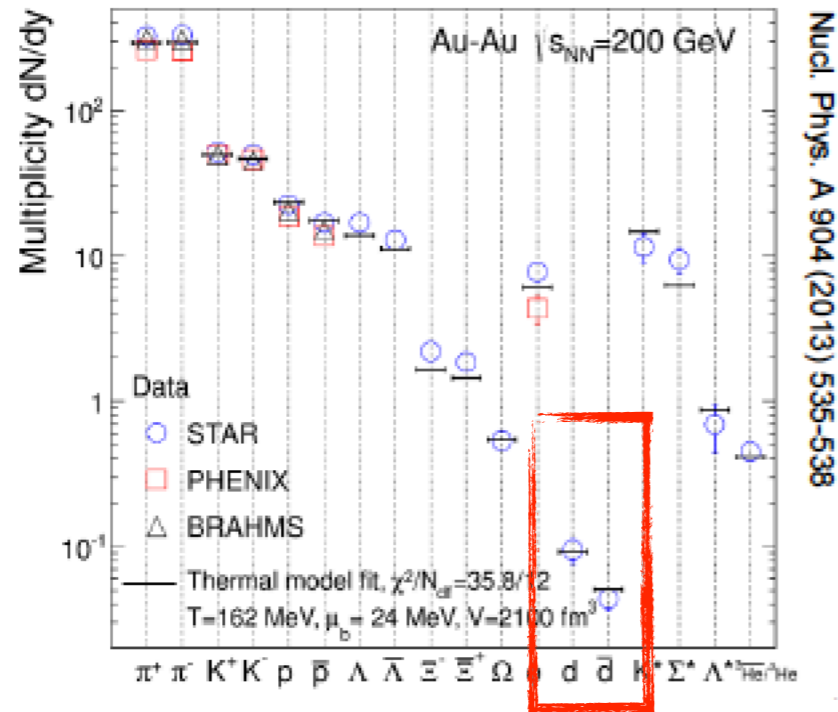


Always-on-switch

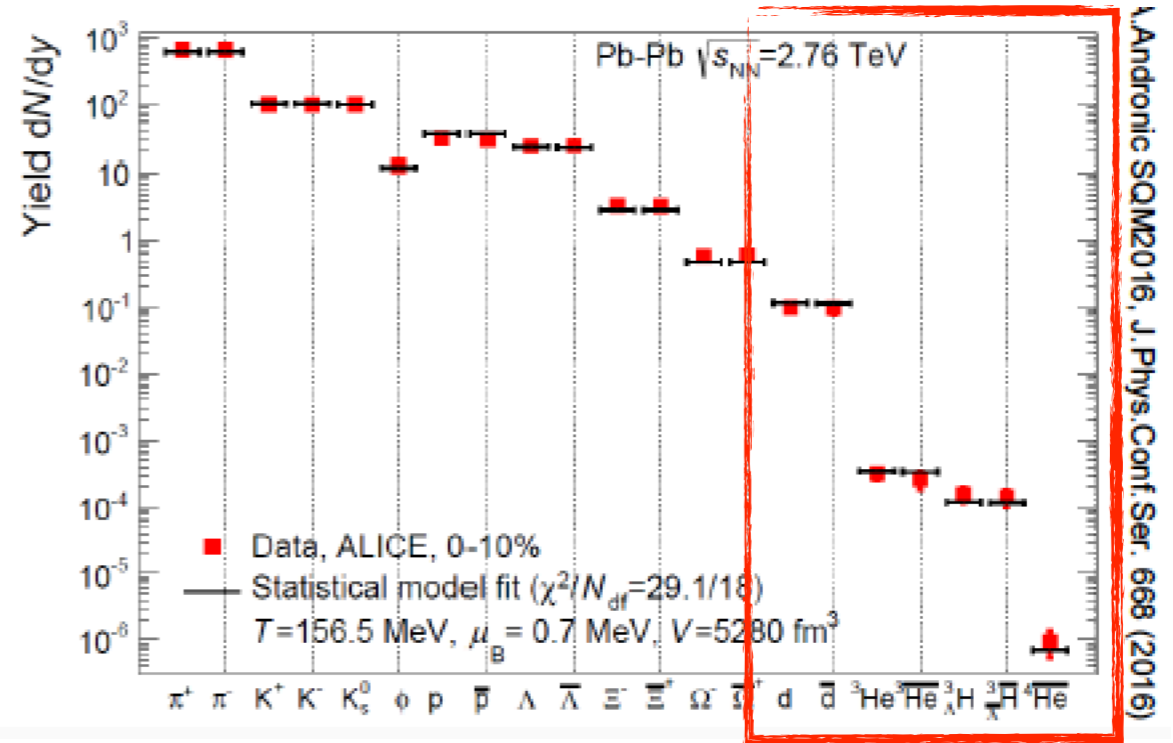
Yesterday's mystery



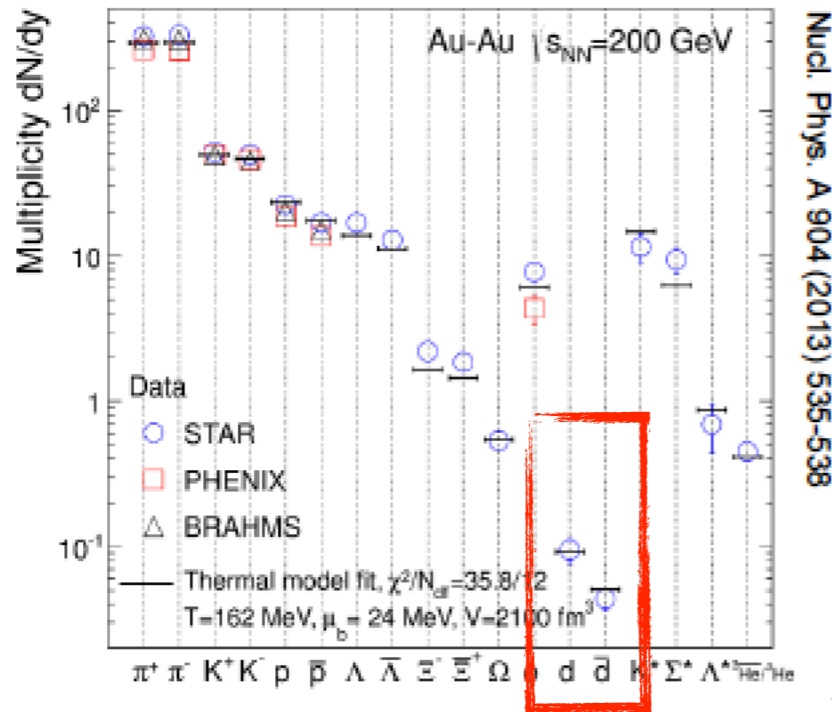
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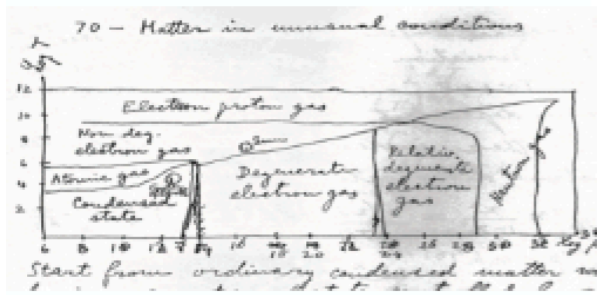
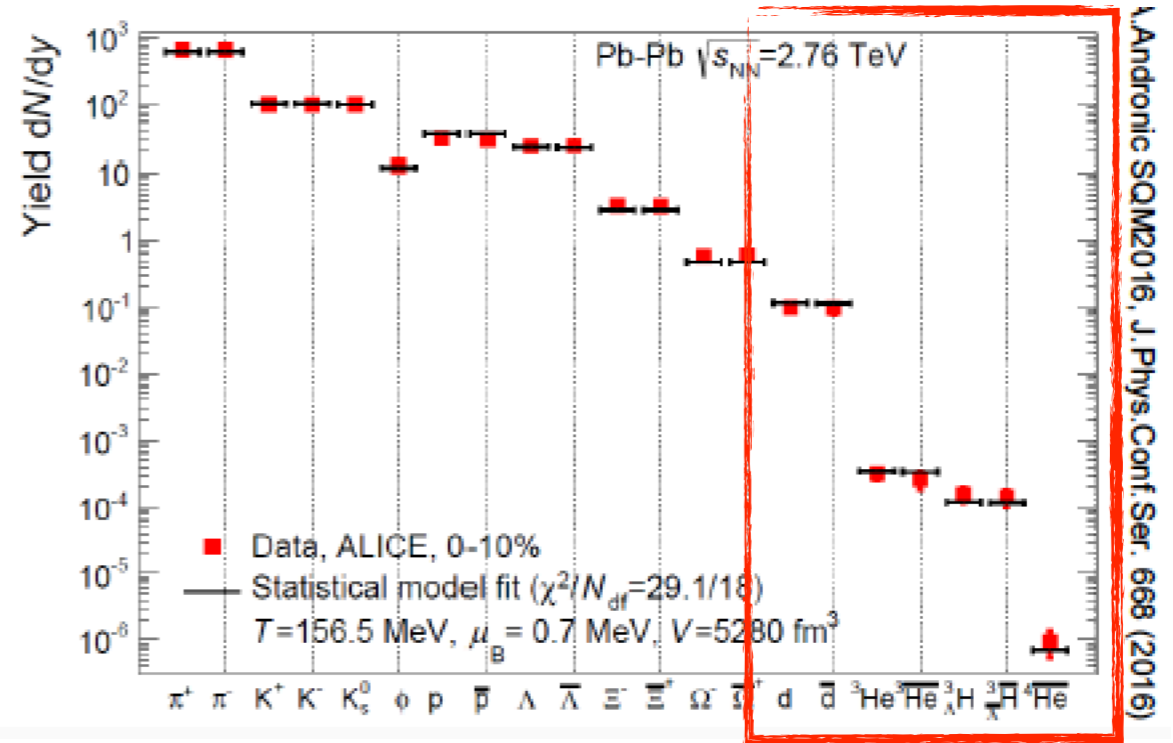
today's mystery, as well!



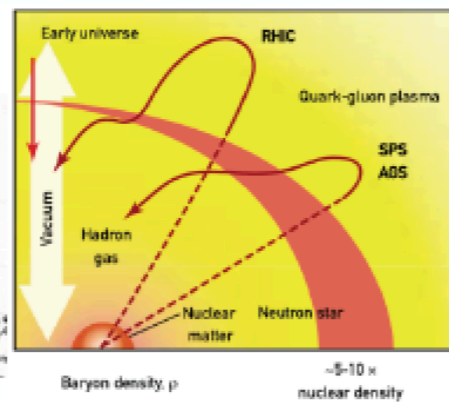
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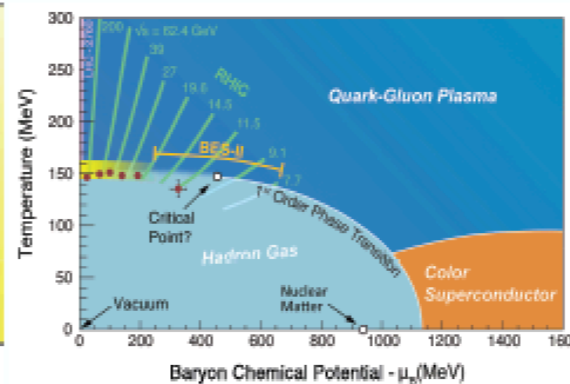
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Fermi (1952)



2002



2015



2020+

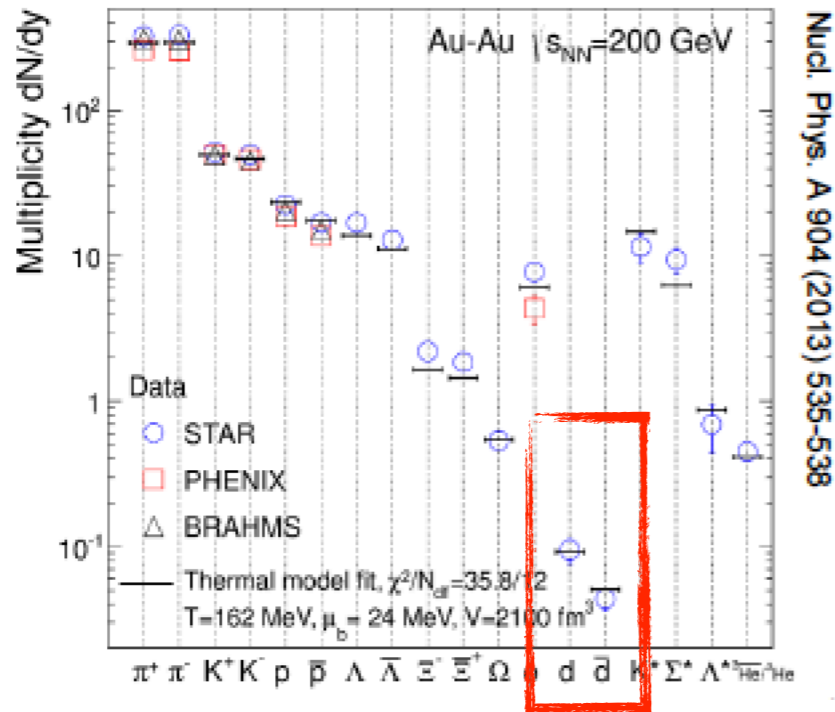
RHIC starts running

BESII

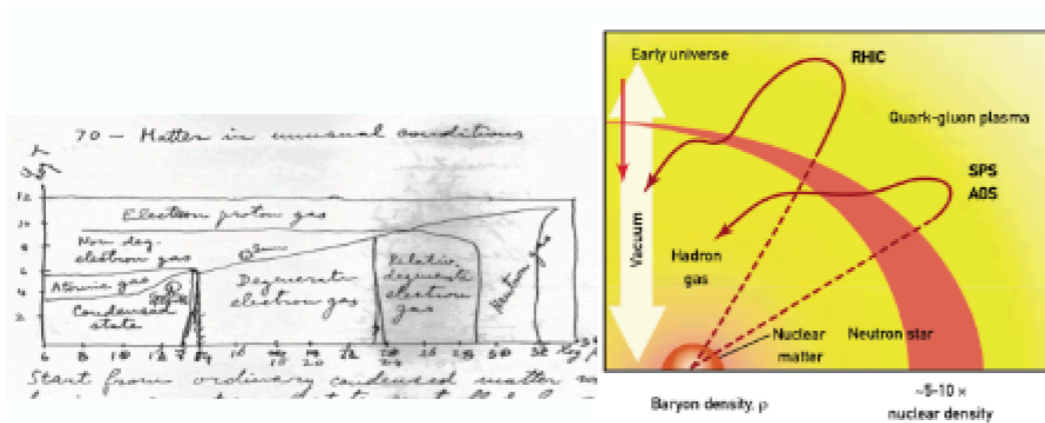
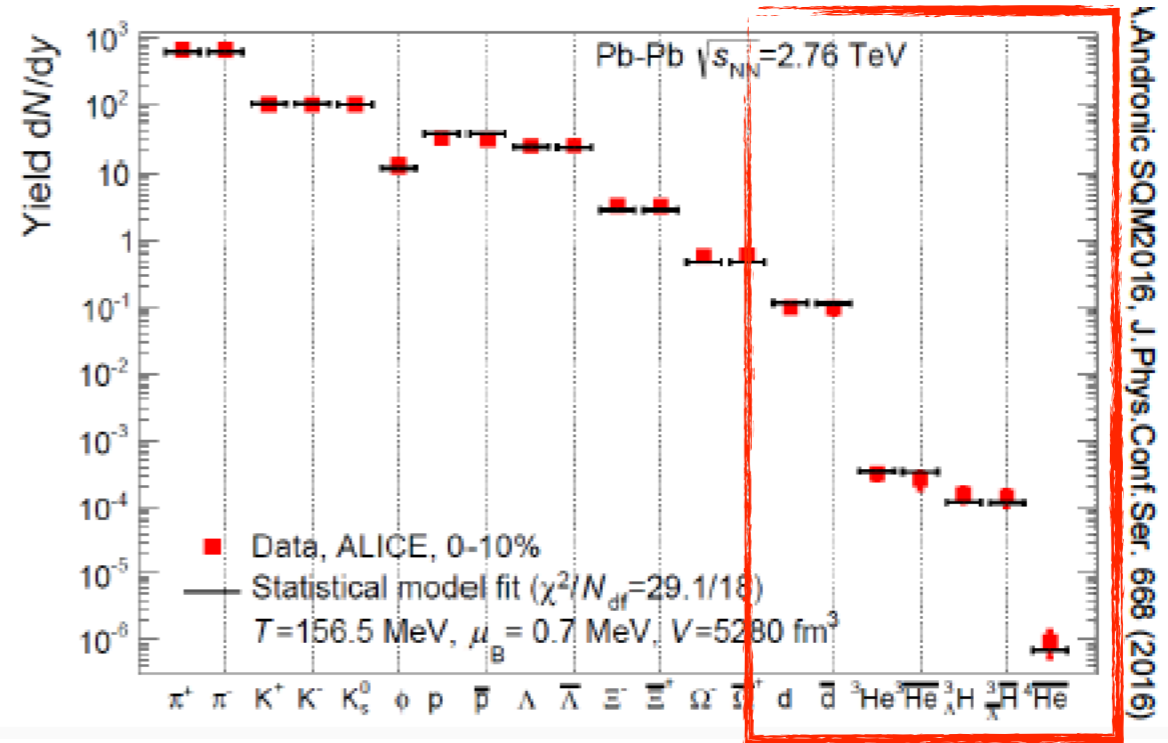


Yesterday is history, tomorrow is mystery and today is a present

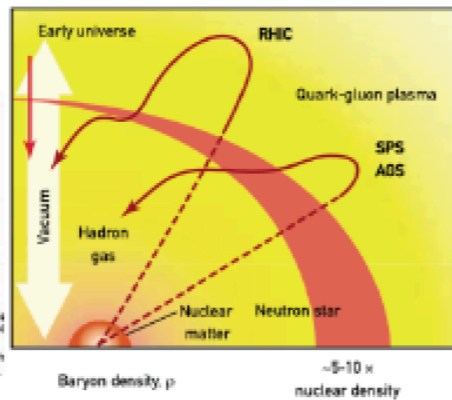
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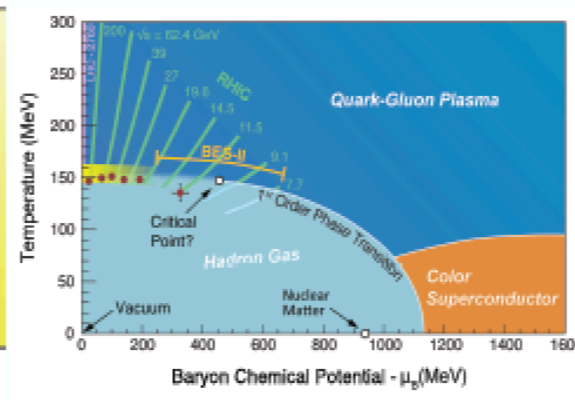
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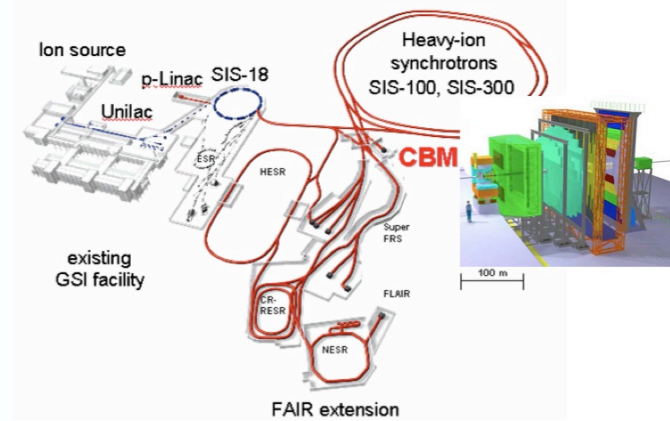
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The Compressed Baryonic Matter experiment

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