

Introduction of DJBUU : New Transport code

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partially on behalf of DJBUU Project



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a pure Korean word meaning Delightful, Joyful, Happy, ...

in 2011, Korean government approved a **Rare Isotope Accelerator Project**

- My recent works have been related to neutron stars
- NS EoS / Dense Matter
 - NS Binary Evolution / Gravitational Waves

for RAON

better to start from where you have an advantage

Some experience on Heavy Ion Collisions at Stony Brook

- Lee, Wirstam, Zahed, Hansson, PLB 448, 168 (1999)
- Lee, Yamagishi, Zahed, PRC 58, 2899 (1998)
- Lee, Yamagishi, Zahed, NPA 653, 185 (1999)

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Contents

Part I : EM Radiation for RHIC & LHC

Part II : DJBUU for Rare Isotope Collisions

Part I: EM Radiation in Hot QCD Matter

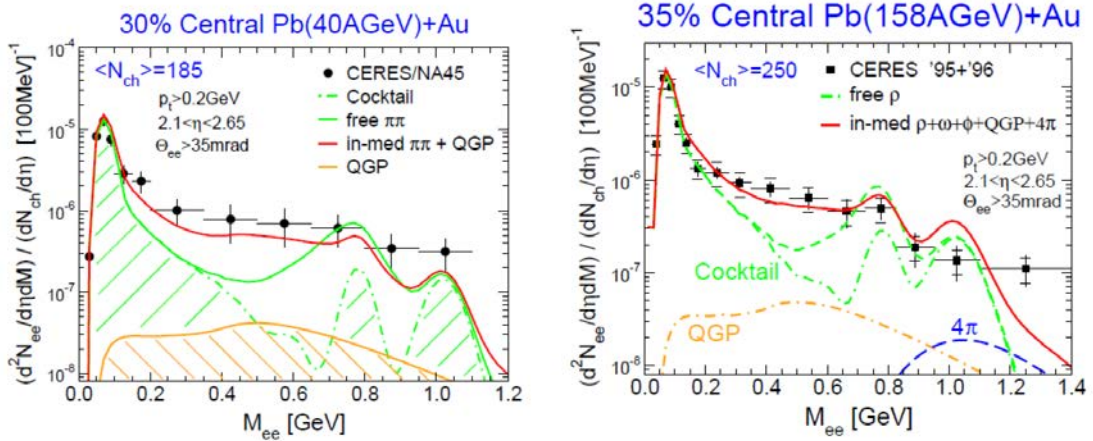
in collaboration with
Y.M. Kim (PNU), D. Teaney, I. Zahed (Stony Brook)
PRC 90, 025204 (2014) & arXiv:1610.06213

Why Photons & Dileptons ?

- No strong interaction
- Can provide direct information on dense medium
- Right time to revisit

CERES/NA45 Pb+Au 8.8 & 17.3 GeV

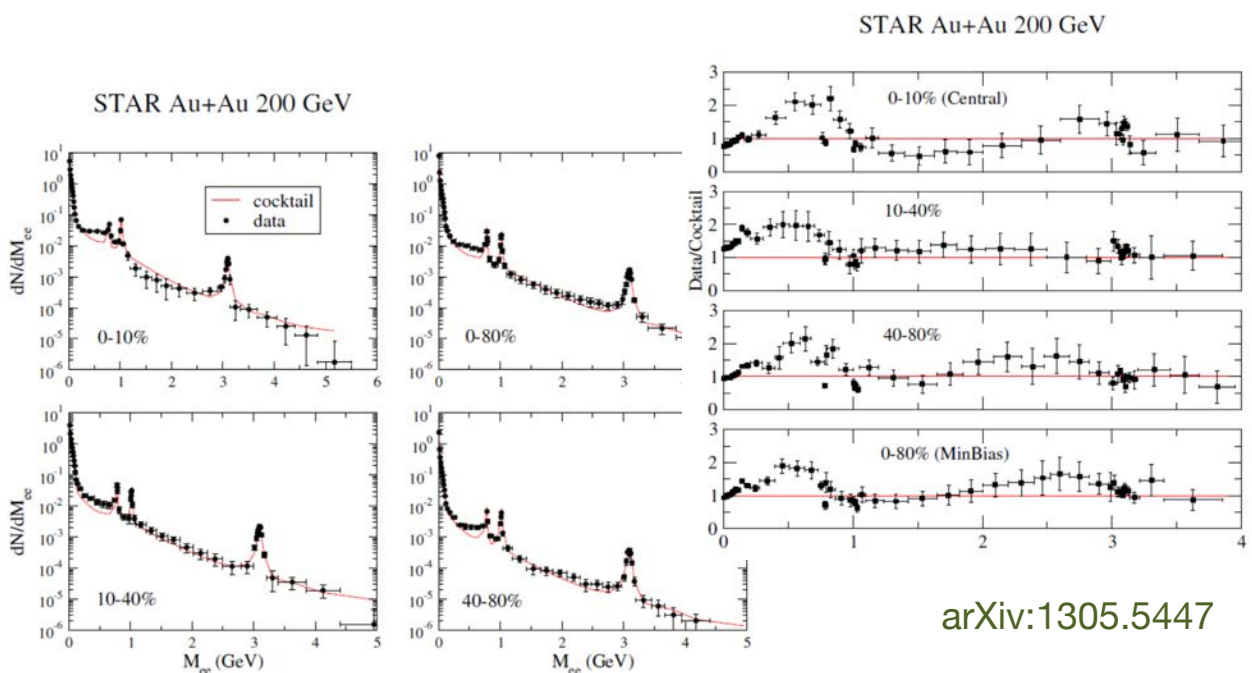
R.Rapp, arXiv:1306.6394



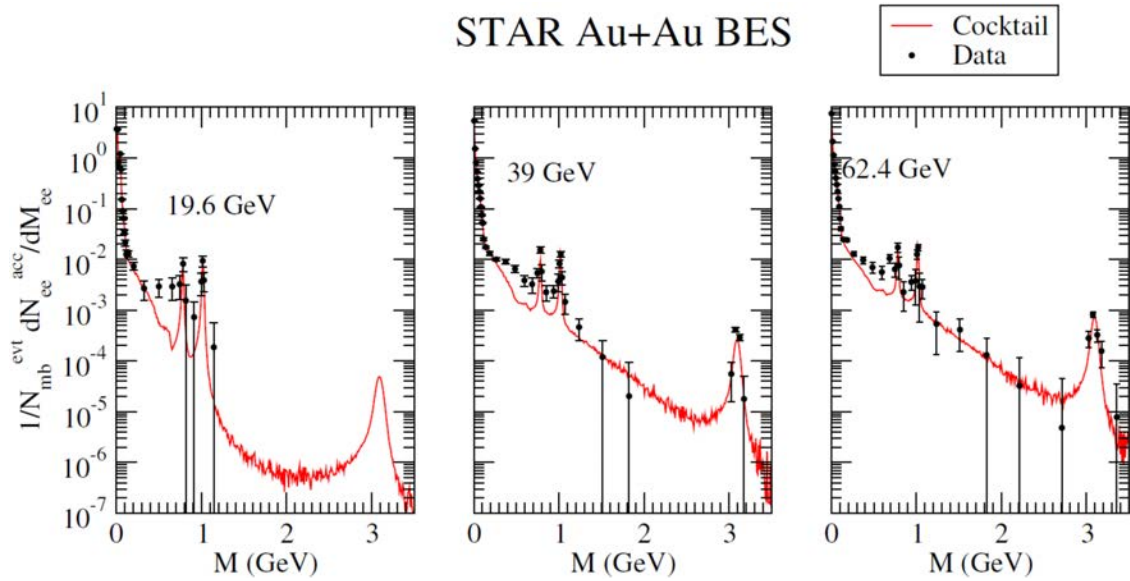
Key question : low-mass dilepton enhancement

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STAR Dilepton Enhancement Au+Au 200 GeV

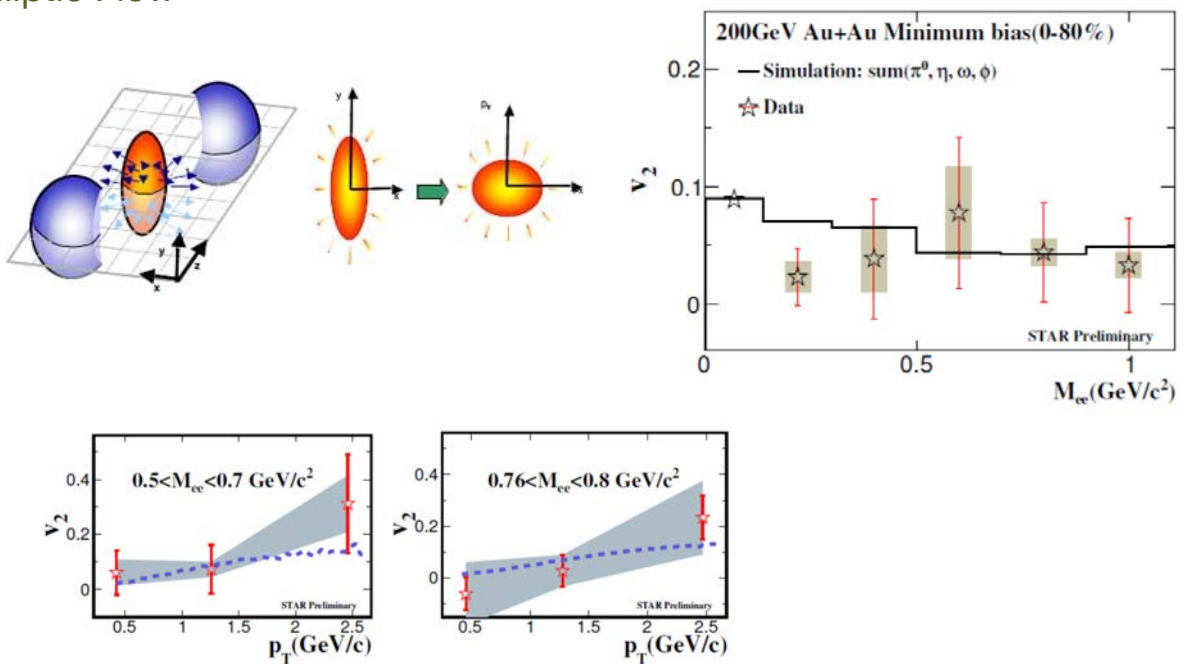


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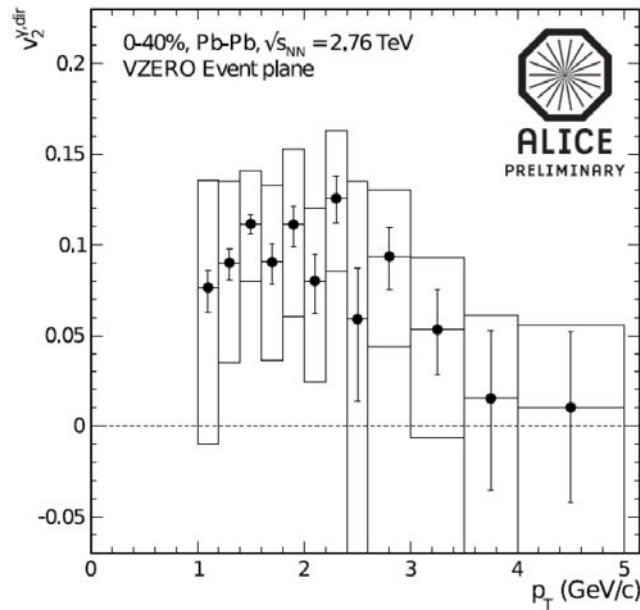
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Elliptic Flow



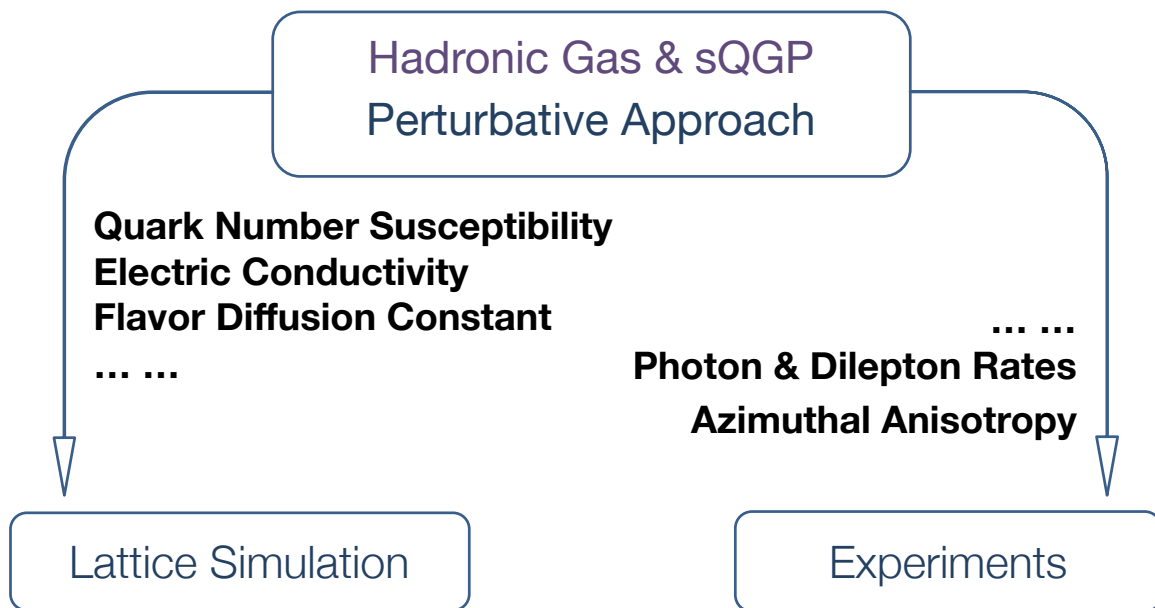
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Elliptic Flow



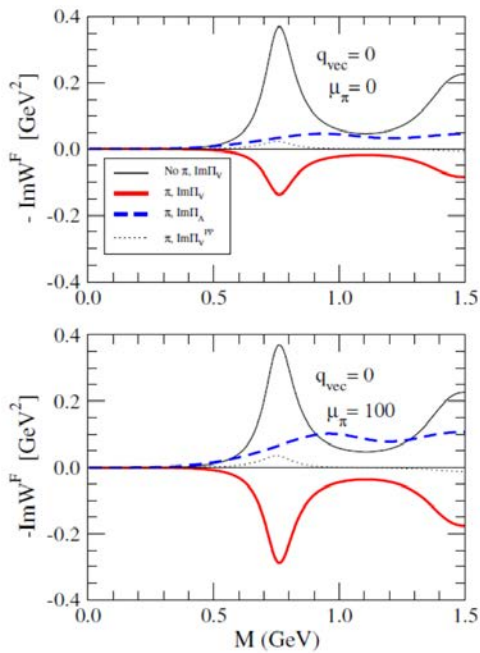
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Theory vs Experiment



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Mixing between vector-axial : Chiral symmetry restoration



As pion chemical potential increase

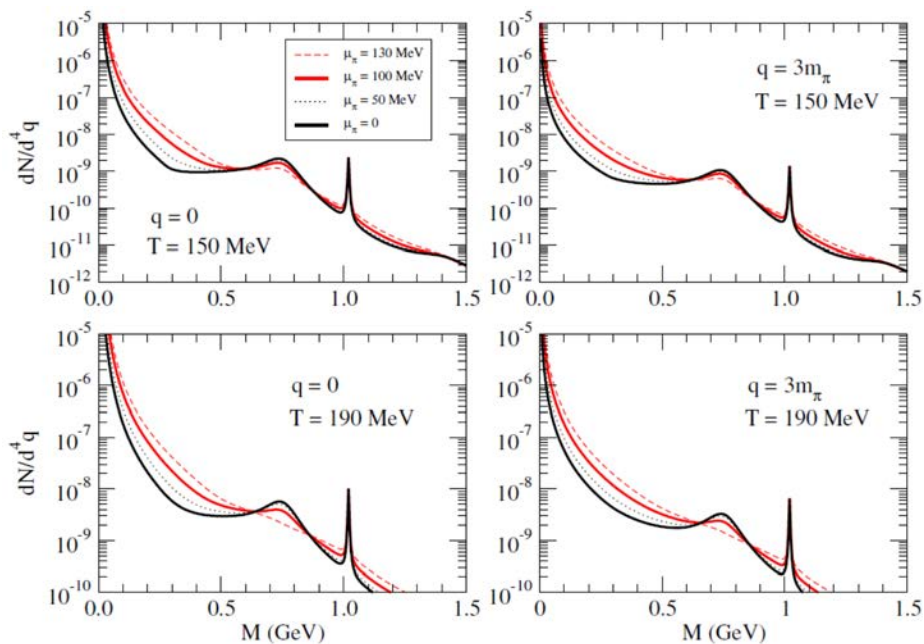
- **Reduction of Vector Contribution** due to the cancellation (no pion + pion contribution)
- **Enhancement of Axial Contribution**

Lee & Zahed PRC 90, 025204 (2014)

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Dilepton Rates up to two pion

Low-mass enhancement due to mixing between vector & axial



Lee & Zahed PRC 90, 025204 (2014)

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Conclusion of Part I

- **Low-mass dilepton enhancement (PRC 90, 025204, 2014)**
 - partial restoration of chiral symmetry
 - mixing between vector & axial correlators
- **Charged particle elliptic flow (arXiv:1610.06213)**
 - pion is better than (anti-)proton
 - hadronic rates dominate the photon emissivity rates
 - still misses ALICE high q_T data
- **Future Plan**
 - inclusion of nucleons for STAR BES

Part II

DJBUU project

- What is DJBUU
Dae Jeon Boltzmann-Uehling-Uhlenbeck
- Current collaboration members
S. Jeon (McGill, chair) at RISP during 2015.10~2016.01
Y. Kim, J. Hong, K. Kim (RISP)
M.K.Kim, Y.M. Kim, C.-H. Lee (PNU)
- Supported by RISP

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Contents of Part II

1. Boltzmann-Uehling-Uhlenbeck (BUU)
2. Daejeon Boltzmann-Uehling-Uhlenbeck (DJBUU)
3. What has been done before with RBUU
 - Sn + Pb @ 200 MeV/u
 - Xe + Pb @ 200 MeV/u
 - Au + Au @ 200 MeV/u
4. What we are doing now
 - Single Au density profile
 - Au + Au @ 100 MeV/u
 - Pauli Blocking factor
5. Summary and Plan

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Nuclear transport theory

- Treats non-equilibrium process
- One-body phase-space distribution in heavy-ion collision
 - mean field + two-body collision + Pauli blocking
- Boltzmann-Vlasov type :
 - evolution of the one-body phase-space density under the influence of a mean field
- molecular-dynamics type :
 - nucleon coordinates and momenta under the action of a many-body Hamiltonian
- test particle
 - solve nonlinear integro-differential eq. (= BUU eq.)

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Relativistic mean field theory

$$\begin{aligned}
 \mathcal{L} = & \bar{\psi} [i\partial\!\!\!/ - (m_N - g_\sigma\sigma) - g_\omega\omega - g_\rho\vec{\tau} \cdot \vec{\rho}] \psi \\
 & + \frac{1}{2} (\partial_\mu\sigma\partial^\mu\sigma - m_\sigma^2\sigma^2) - \frac{1}{3}a\sigma^3 + \frac{1}{4}b\sigma^4 \\
 & + \frac{1}{2}m_\omega^2\omega_\mu\omega^\mu - \frac{1}{4}(\partial_\mu\omega_\nu - \partial_\nu\omega_\mu)(\partial^\mu\omega^\nu - \partial^\nu\omega^\mu) \\
 & + \frac{1}{2}m_\rho^2\vec{\rho}_\mu \cdot \vec{\rho}^\mu - \frac{1}{4}(\partial_\mu\vec{\rho}_\nu - \partial_\nu\vec{\rho}_\mu) \cdot (\partial^\mu\vec{\rho}^\nu - \partial^\nu\vec{\rho}^\mu).
 \end{aligned}$$

Parameter	Set I	Set II
f_σ (fm ²)	10.33	same
f_ω (fm ²)	5.42	same
f_ρ (fm ²)	0.95	3.15
f_δ (fm ²)	0.00	2.50
A (fm ⁻¹)	0.033	same
B	-0.0048	same

Phys. Rev. C, 65, 045201 (2002)

- Describe nuclear structure (phenomenological)
- Applied from finite nuclei to neutron star
- Mean field theory including σ , ω , ρ meson fields (exclude δ -meson field in DJBUU simulation)

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Boltzmann-Uehling-Uhlenbeck eq.

$$\begin{aligned} & \left[\tilde{k} \cdot \partial^{(x)} + \left(\tilde{k}_\nu F^{\mu\nu} + m^* \partial^\mu m^* \right) \partial_\mu^{(\tilde{k})} \right] f(x, \tilde{k})|_{\tilde{k}=k_1+\Sigma} \\ &= \frac{1}{2} \int \mathcal{D}_{k_2, k_3, k_4} W(k_1 k_2 | k_3 k_4) \delta^4(k_1 + k_2 - k_3 - k_4) \\ & \quad \times [f_3 f_4 (1 - f_1)(1 - f_2) - f_1 f_2 (1 - f_3)(1 - f_4)]. \end{aligned}$$

- Boltzmann eq. - **collision term**
- Uehling-Uhlenbeck eq. - **Pauli-blocking factor**
- Time evolution of the one-body phase-space density under the mean field potential

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Numerical realization

RBUU code : simulate Heavy-ion collisions

- 1995, first developed in Munich (C. Fuchs)
- 1996-2000, density-dep. RMF models, DBHF approaches (T. Gaitanos, C. Fuchs)
- 2002-2005, isospin effects in the production thresholds (G. Ferini, T. Gaitanos)
- 2005-2010, in-medium isospin effects in cross sections&kaon pot. (V. Prassa, T. Gaitanos)
- 2014, improvement in stability (RISP)

Test particle method > Parallel ensemble method

single particle phase-space distribution function represented by N covariant Gaussian test particles

$$g(x, x_i) = \frac{1}{(\pi\sigma^2)^{3/2}} e^{[(x_\mu - x_{i\mu})^2 - (x_\mu - x_{i\mu})u_i^\mu]^2 / \sigma^2}$$

$$g(k^*, k_i^*) = \frac{1}{(\pi\sigma_k^2)^{3/2}} e^{[k_\mu^{*2} - (k_\mu^* u_i^\mu)^2] / \sigma_k^2}$$

$$\blacktriangleright f(x, k^*) = \frac{1}{N} \sum_{i=1}^{AN} g(x, x_i) g(k^*, k_i^*)$$

$\sigma = 1.4\text{fm}$
 $\sigma_k = 0.346\text{fm}^{-1}$
 $N=100$

slide from Y.J. Lee

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Baryon density $\rho_B(x) = \frac{1}{N} \sum_{i=1}^{AN} g(x, x_i) u_{i0}$

Isospin asymmetry $\rho_I(x) = \frac{[\rho_n(x) - \rho_p(x)]}{\rho_B(x)}$

Local temperature by a fit of the **RBUU momentum distribution** to the **Fermi-Dirac distribution**, assuming **local thermodynamic equilibrium**

$$n(x, \mathbf{k}) = \frac{(2\pi)^3}{4} \frac{1}{N(\pi\sigma\sigma_k)^3} \sum_{i=1}^{AN} e^{[(x_\mu - x_{i\mu})^2 - ((x_\mu - x_{i\mu})u_i^\mu)^2]/\sigma^2} e^{[k^{*2} - (k_\mu^* u_i^\mu)^2]/\sigma_k^2}$$

Nucl. Phys. A, 589, 732, (1995)

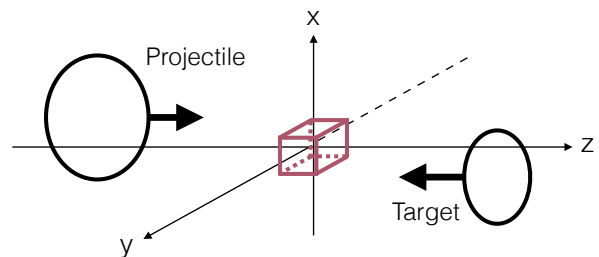
$$n(x, \mathbf{k}, T) = \frac{1}{1 + \exp[-(\mu^* - k_\mu^* u^\mu)/T]} \quad (k_0^* = \sqrt{\mathbf{k}^{*2} + m^{*2}})$$

... T is the only free parameter, μ^* is T-dependent parameter,
Determine μ^* from the **baryon number conservation**

$$j_0(x) = 4 \int \frac{d^3k}{(2\pi)^3} n(x, \mathbf{k}, T)$$

Consider the point $x=0$

... practically, a **(5x5x5)fm³ cubical box** around the center



slide from Y.J. Lee

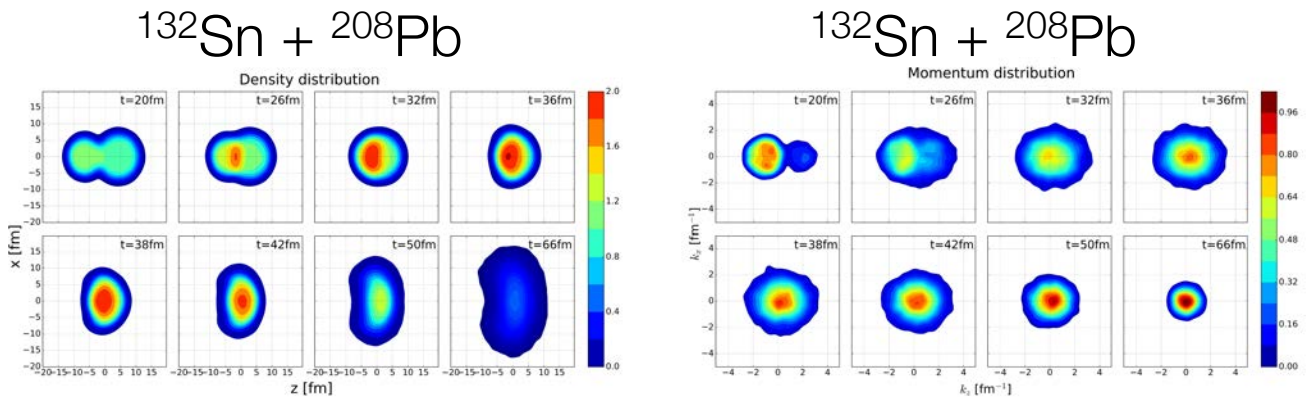
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RBUU results (what has been done before)

JKPS,69, 1430 (2016), Y. Lee, C.-H. Lee, T. Gaitanos, Y. Kim

NPSM (2017), Myungkuk Kim, Y. Lee, Y. M. Kim, C.-H. Lee

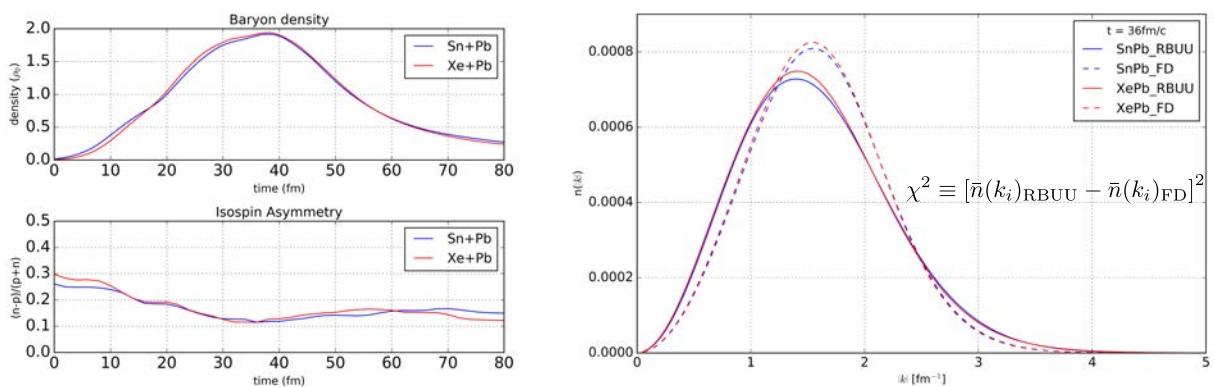
Result 1 : $^{132}\text{Sn} + ^{208}\text{Pb}$ @ 200 MeV/u



- Nuclear structure : near the magic number 126
- Neutron star internal structure
- Dense matter : symmetry energy and collective flow
- $^{140}\text{Xe} + ^{208}\text{Pb}$: similar evolution history for density and momentum

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Central density and effective temperature

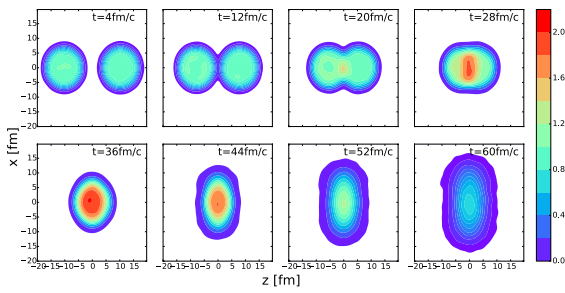


- Central density reaches up to $2\rho_0$ ($\sim 0.16 \text{ fm}^{-3}$) at $t = 36 \sim 38 \text{ fm}/c$
- Symmetry energy \propto density (excluding δ -meson field)
- T (MeV) - 27.1 and 27.8 @ 36 fm/c
- μ_B (MeV) - 721.7 and 706.3 @ 36 fm/c

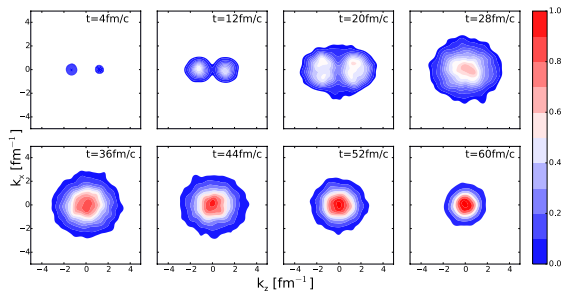
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Result 2 : $^{197}\text{Au} + ^{197}\text{Au}$ @ 200 MeV/u

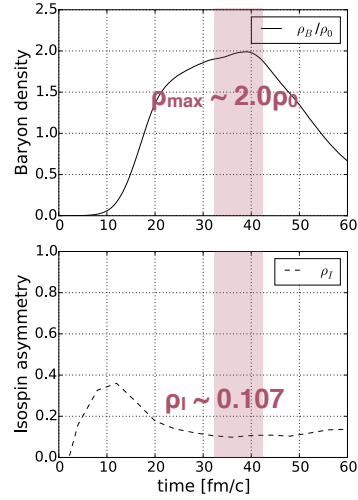
▼ Time evolution of density distribution



▼ Time evolution of momentum distribution at x=0

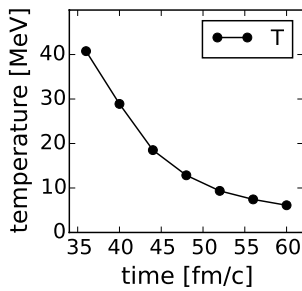


▼ Time evolution of baryon density and isospin asymmetry at the center

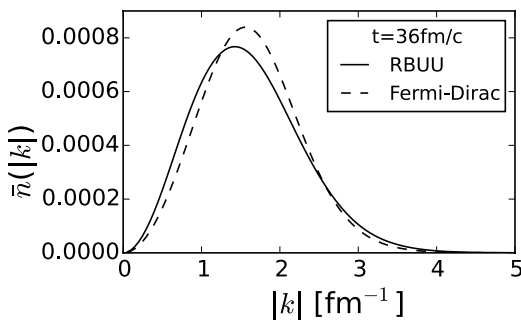


(ρ_0 : saturation density, $\sim 0.16\text{fm}^{-3}$)

slide from Y.J. Lee 29



time [fm/c]	36	40	44	48	52	56	60
T [MeV]	40.8	28.9	18.5	12.6	9.3	7.4	6.1
μ^* [MeV]	771.7	802.9	813.4	842.3	867.2	885.9	899.1



◀ 1-dimensional RBUU & fitted Fermi momentum distribution at the center at time $t=36\text{fm}/c$



$$\chi^2 = 2.41 \times 10^{-7}$$

$$(\chi^2 \equiv [\bar{n}(k_i)_{\text{RBUU}} - \bar{n}(k_i)_{\text{FD}}]^2, i : \text{all grid points})$$

► The first application of a transport model to HIC at low energy.

slide from Y.J. Lee 30

What we have done with RBUU

- Simulation of HIC by using the RBUU transport code for estimation of properties of nuclear matter that is expected to be created in 
- Baryon density, isospin asymmetry, local temperature for $^{132}\text{Sn} + ^{208}\text{Pb}$ and $^{140}\text{Xe} + ^{208}\text{Pb}$ @ 200A MeV
: $\rho_{\text{max}} \sim 2\rho_0$, $\rho_I \sim 0.1$, $T_{\text{loc}} \sim 27$ MeV.
- Quantitative discussion on the possibility of QCD phase transition, liquid-gas phase transition, pion condensation, ...etc. at 
- Simulations with more neutron-rich isotopes/deformed nuclei.
- Ongoing development of advanced transport calculation code for better results.

slide from Y.J. Lee 31

DJBUU

Daejeon Boltzmann-Uehling-Uhlenbeck eq.

- c / c++ language
- openMP (Open Multi-Processing) implemented
- easy to follow & modify
- simulated mainly in Mac OSX

```
E_beam_NN_for_Heavy_Ion_Collision_in_GeV 0.2
Record_p_n_densities_at_the_center_1_for_on_0_for_off 1
Radius_for_density_calc_in_fm 3.0
Record_interval 10
Record_particle_states_1_for_on_0_for_off 1
Record_Interval 10
Turn_on_Coulomb_1_for_on_0_for_off 0
Number_of_grid_points_in_x 100
SigmaNN_CutOff_in_mb 50
Uncertainty_param_dxdp 0.6
```

RBUU code

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- 2002-2005, isospin effects in the production thresholds (G. Ferini, T. Gaitanos)
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- 2014, improvement in stability (RISP)

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DJBUU preliminary results

- time evolution of single Au density profile
- average density contours in Au + Au collision @ 100 MeV/u ($b=7$ fm)
- Pauli blocking

NPSM, RI Science Special Issue, M. Kim, CHL, Y.Kim, S. Jeon

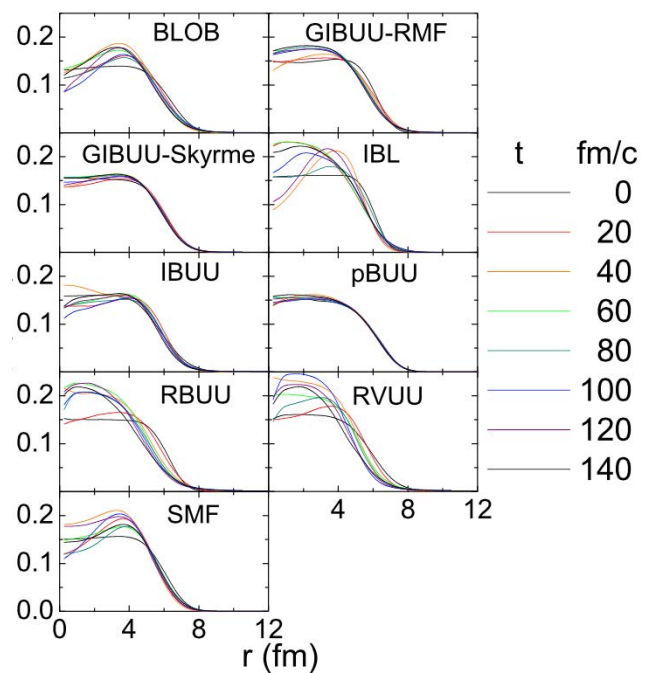
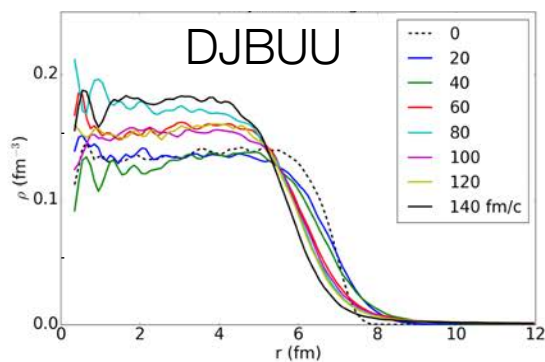
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DJBUU vs RBUU result

- 현재 직접적인 비교 불가 할 것으로 판단됩니다.
(comments by **Myungkuk Kim**)

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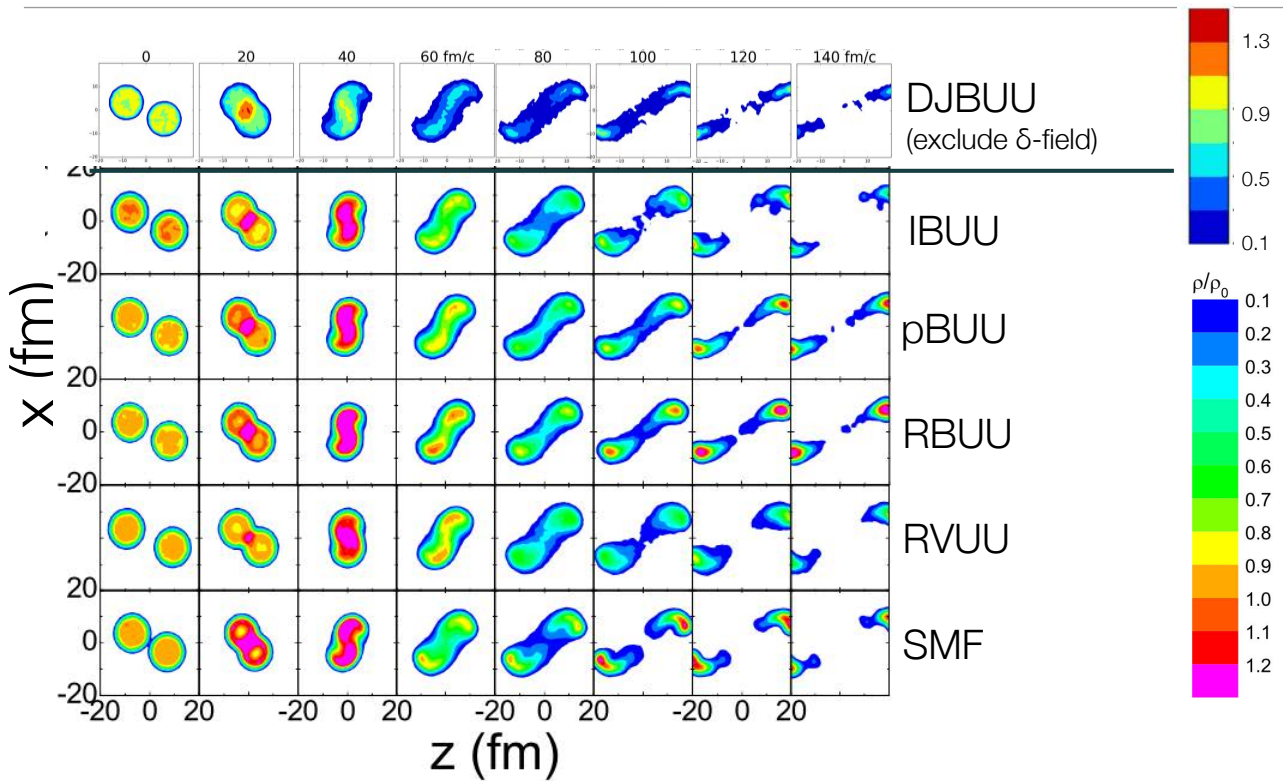
Single Au density profile (preliminary)



- Check the stability and initial condition
(3 run, 100 test particle)
- Initial density $\sim 0.15 \text{ fm}^{-3}$
(from relativistic Thomas-Fermi approx.)
- fluctuation $0.13 \sim 0.17 \text{ fm}^{-3}$

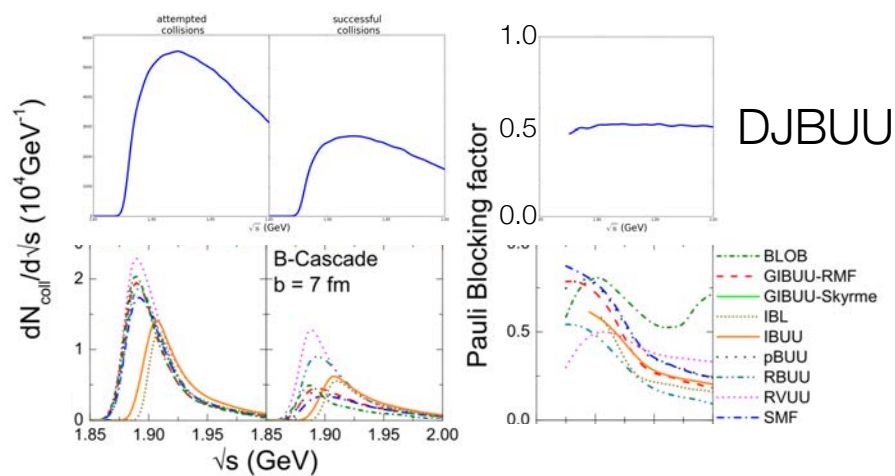
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Au + Au collision @ 100 MeV/u ($b=7$ fm) (preliminary)



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Pauli Blocking factor (for one case) (preliminary)



- B : 100 MeV/u
- Cascade - turn off the mean field (collision term)
- Pauli Blocking factor = 1 - (successful / attempted)

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Part II: Summary & Plan

- **DJBUU**
 - new BUU-type code is developed.
 - DJBUU code is stable, but still preliminary.
 - requires fine tuning.
- **Plan**
 - test new equation of states