

QCD sum rules for ρ -meson at finite density or temperature

Youngshin Kwon

Heavy Ion Meeting at KNU

October 9, 2010



YONSEI
UNIVERSITY

contents based on :

YK., Procura & Weise [Phys. Rev. C78, 055203 (2008)]

YK, Sasaki & Weise [Phys. Rev. C81, 065203 (2010)]

Outline

- ⊙ Introduction & motivation
- ⊙ QCD sum rules for ρ -meson, in vacuum and in medium
- ⊙ Finite energy sum rules at finite density
- ⊙ Finite energy sum rules at finite temperature
- ⊙ Summary and outlook

Goal: reliable framework of in-medium QCD sum rules for vector mesons
 \Rightarrow constraints for the in-medium spectral properties

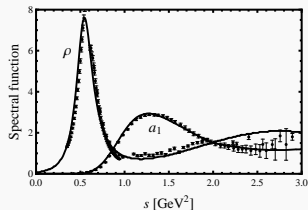
Motivation

⊙ Spontaneous chiral symmetry breaking:

- ▶ quark condensate: $\langle \bar{q}q \rangle \neq 0$
- ▶ Goldstone bosons: π , K , etc.
pion decay constant: $f_\pi \simeq 92.4$ MeV
- ▶ mass splitting of chiral partners
(e.g. $\rho(770)$ - $a_1(1250)$)

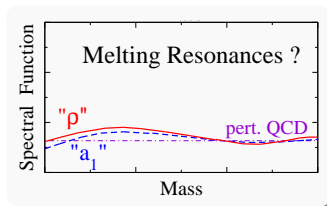
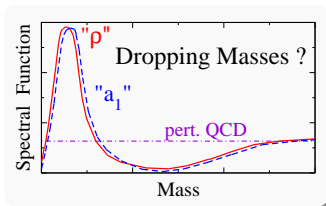
⊙ Chiral symmetry restoration in nuclear medium:

- ▶ degenerate chiral partners \Rightarrow modifications of hadron spectrum



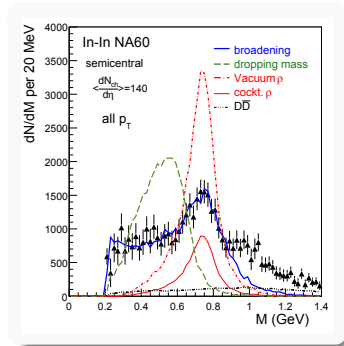
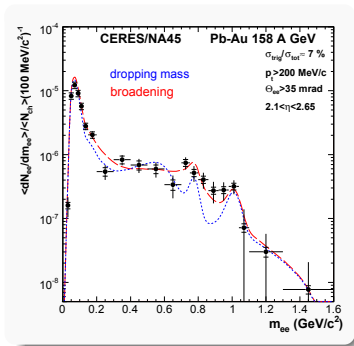
Restoration scenarios in medium

- ⊙ Pole mass shift:
 - ▶ masses of parity partners degenerate in medium.
 - ▶ moving toward each other or going to zero (Brown-Rho).
- Brown & Rho [PRL 66, 2720 (1991)]
- ⊙ Width broadening:
 - ▶ in-medium spectral functions are broadly distributed.
 - ▶ the continuum merges the broadened spectral distributions.



Dilepton spectroscopy

- ⊙ Dilepton production in RHIC ($\gamma^* \rightarrow l^+ l^-$):
 - ▶ EM probe with pure information of the hot and/or dense region
 - ▶ dilepton emission \Leftrightarrow in-medium vector-meson spectroscopy



General review of QCD sum rules (in vacuum)

General review of QCD sum rules

- Current correlation function:

$$\Pi^{\mu\nu}(q) = i \int d^4x e^{iq \cdot x} \langle \mathcal{T} j^\mu(x) j^\nu(0) \rangle$$

- isovector vector- and axialvector-currents:

$$j_\rho^\mu = \frac{1}{2} (\bar{u} \gamma^\mu u - \bar{d} \gamma^\mu d), \quad j_A^\mu = \frac{1}{2} (\bar{u} \gamma^\mu \gamma_5 u - \bar{d} \gamma^\mu \gamma_5 d)$$

- invariant correlator: $\Pi(q^2) = \frac{1}{3} g_{\mu\nu} \Pi^{\mu\nu}(q)$

- Operator product expansion** (quark & gluon d.o.f.) at large $Q^2 = -q^2$:

$$\frac{12\pi^2}{Q^2} \Pi(Q^2) = -c_0 \ln \left(\frac{Q^2}{\mu^2} \right) + \frac{c_1}{Q^2} + \frac{c_2}{Q^4} + \dots$$

- Spectral representation** (hadronic d.o.f.) at resonance region:

$$\Pi(q^2) = \Pi(0) + q^2 \Pi'(0) + \frac{q^4}{\pi} \int ds \frac{\text{Im} \Pi(s)}{s^2(s - q^2 - i\epsilon)}$$

General review of QCD sum rules

- ⊙ Borel transformation:

$$12\pi^2\Pi(0) + \int_0^\infty ds R(s) e^{-s/M^2} = c_0 M^2 + c_1 + \frac{c_2}{M^2} + \frac{c_3}{2M^4} + \dots$$

- ▶ dimensionless spectral function: $R(s) \equiv -\frac{12\pi}{s} \text{Im} \Pi(s)$

- ⊙ Coefficients c_n :

$$c_0 = \frac{3}{2} \left(1 + \frac{\alpha_s}{\pi}\right) + \dots, \quad c_1 \propto m_q^2 : \text{negligibly small}$$

$$c_2 = \frac{\pi^2}{2} \left\langle \frac{\alpha_s}{\pi} G^2 \right\rangle \pm 6\pi^2 \left(m_u \langle \bar{u}u \rangle + m_d \langle \bar{d}d \rangle \right)$$

$$0.005 \pm 0.004 \text{ GeV}^4$$

Ioffe [PPNP 56, 232 (2006)]

$$\simeq -m_\pi^2 f_\pi^2 = -(0.11 \text{ GeV})^4$$

[Gellman-Oaks-Renner]

General review of QCD sum rules

- ⊙ Borel transformation:

$$12\pi^2\Pi(0) + \int_0^\infty ds R(s) e^{-s/M^2} = c_0 M^2 + c_1 + \frac{c_2}{M^2} + \frac{c_3}{2M^4} + \dots$$

- ▶ dimensionless spectral function: $R(s) \equiv -\frac{12\pi}{s} \text{Im} \Pi(s)$

- ⊙ Coefficients c_n :

$$c_0 = \frac{3}{2} \left(1 + \frac{\alpha_s}{\pi} \right) + \dots, \quad c_1 \propto m_q^2 : \text{negligibly small}$$

$$c_2 = \frac{\pi^2}{2} \left\langle \frac{\alpha_s}{\pi} G^2 \right\rangle \pm 6\pi^2 \left(m_u \langle \bar{u}u \rangle + m_d \langle \bar{d}d \rangle \right)$$

$$c_3 \propto \mp \langle (\bar{q}q)^2 \rangle \text{ uncertain value}$$

General review of QCD sum rules

⊙ Borel transformation:

$$12\pi^2\Pi(0) + \int_0^\infty ds R(s) e^{-s/M^2} = c_0 M^2 + c_1 + \frac{c_2}{M^2} + \frac{c_3}{2M^4} + \dots$$

- ▶ dimensionless spectral function: $R(s) \equiv -\frac{12\pi}{s} \text{Im} \Pi(s)$
- ▶ expand for $s_0 \ll M^2$ and compare term by term

⊙ Coefficients c_n :

$$c_0 = \frac{3}{2} \left(1 + \frac{\alpha_s}{\pi} \right) + \dots, \quad c_1 \propto m_q^2 : \text{negligibly small}$$

$$c_2 = \frac{\pi^2}{2} \left\langle \frac{\alpha_s}{\pi} G^2 \right\rangle \pm 6\pi^2 \left(m_u \langle \bar{u}u \rangle + m_d \langle \bar{d}d \rangle \right)$$

$$c_3 \propto \mp \langle (\bar{q}q)^2 \rangle \text{ **uncertain value**}$$

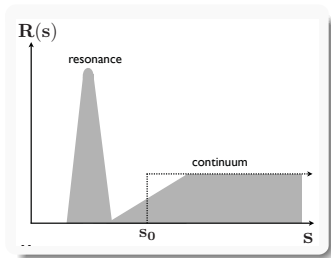
Finite energy sum rules

- Hierarchy of finite energy sum rules for moments of $R(s)$:

$$0^{th} \text{ moment : } \int_0^{s_0} ds R(s) = s_0 c_0 + c_1 - 12\pi^2 \Pi(0)$$

$$1^{st} \text{ moment : } \int_0^{s_0} ds s R(s) = \frac{s_0^2}{2} c_0 - c_2$$

- Spectral distribution (resonance + continuum):



$$R(s) = R_\rho(s) \theta(s_0 - s) + R_c(s) \theta(s - s_0)$$

- Assumption for vector channel;

$$\sqrt{s_0} \simeq 4\pi f_\pi$$

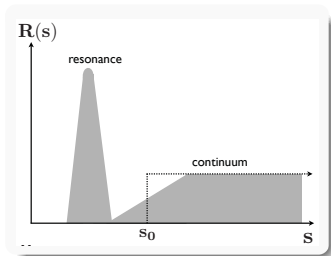
Finite energy sum rules

- ⊙ Hierarchy of finite energy sum rules for moments of $R(s)$:

$$0^{th} \text{ moment : } \int_0^{s_0} ds R(s) = s_0 c_0 + c_1 - 12\pi^2 \Pi(0)$$

$$1^{st} \text{ moment : } \int_0^{s_0} ds s R(s) = \frac{s_0^2}{2} c_0 - c_2$$

- ⊙ Spectral distribution (resonance + continuum):



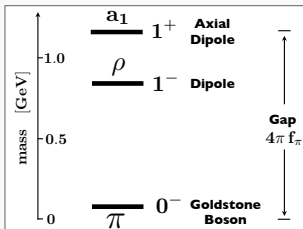
$$R(s) = R_\rho(s) \theta(s_0 - s) + c_0 \theta(s - s_0)$$

- ▶ Assumption for vector channel;

$$\sqrt{s_0} \simeq 4\pi f_\pi$$

Consistency with current algebra

- Identification of $\sqrt{s_0}$ with $\Lambda_{\text{CSB}} \simeq 4\pi f_\pi$:



► KSRF relation

Kawarabayashi & Suzuki [PRL 16, 255 (1966)]

Riazuddin & Fayyazuddin [PR 147, 1071 (1966)]

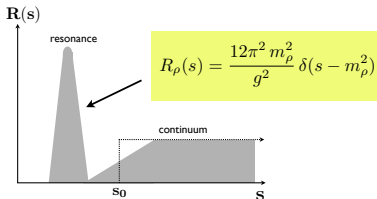
► Weinberg sum rules

Weinberg [PRL 18, 507 (1967)]

$$\Rightarrow m_{a_1} = \sqrt{2} m_\rho = 4\pi f_\pi$$

Consistency with current algebra

- Identification of $\sqrt{s_0}$ with $\Lambda_{\text{CSB}} \simeq 4\pi f_\pi$:



► KSRF relation

Kawarabayashi & Suzuki [PRL 16, 255 (1966)]
 Riazuddin & Fayyazuddin [PR 147, 1071 (1966)]

► Weinberg sum rules

Weinberg [PRL 18, 507 (1967)]
 $\Rightarrow m_{a_1} = \sqrt{2} m_\rho = 4\pi f_\pi$

0th moment:

$$\int_0^{s_0} ds R_\rho(s) = \frac{3}{2} s_0$$

$$\Rightarrow m_\rho^2 = 2g^2 f_\pi^2$$

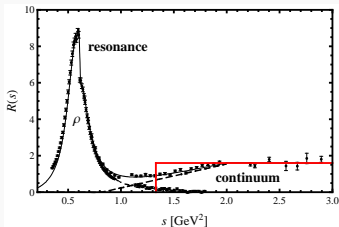
1st moment:

$$\int_0^{s_0} ds s R_\rho(s) = \frac{3}{4} s_0^2$$

$$\Rightarrow g = 2\pi$$

Vacuum sum rule analysis

- ⊙ Input: $R_\rho(s)$ from chiral effective field theory + vector mesons (VMD)



$$\int_0^{s_0} ds R_\rho(s) = s_0 c_0 + c_1$$

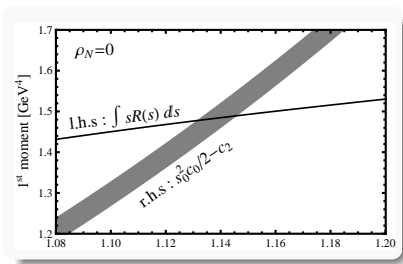
$$\int_0^{s_0} ds s R_\rho(s) = \frac{s_0^2}{2} c_0 - c_2$$

$$\sqrt{s_0} = 1.14 \pm 0.01 \text{ GeV} \simeq 4\pi f_\pi$$

$$\bar{m}_\rho \equiv \sqrt{\frac{\int ds s R(s)}{\int ds R(s)}} = 0.78 \pm 0.01 \text{ GeV}$$

Vacuum sum rule analysis

- Input: $R_\rho(s)$ from chiral effective field theory + vector mesons (VMD)



$$\int_0^{s_0} ds R_\rho(s) = s_0 c_0 + c_1$$

$$\int_0^{s_0} ds s R_\rho(s) = \frac{s_0^2}{2} c_0 - c_2$$

$$\sqrt{s_0} = 1.14 \pm 0.01 \text{ GeV} \simeq 4\pi f_\pi$$

$$\bar{m}_\rho \equiv \sqrt{\frac{\int ds s R(s)}{\int ds R(s)}} = 0.78 \pm 0.01 \text{ GeV}$$

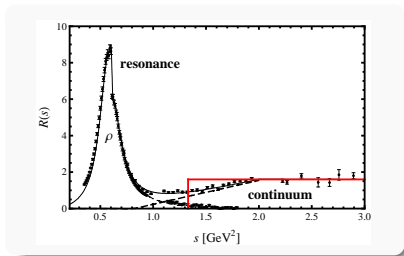
Sensitivity to threshold modeling

- replace the Heaviside step function with a ramp function:

$$R(s) = R_\rho(s) \theta(s_2 - s) + R_c(s) W(s)$$

with the weight function $W(s)$

$$W(x) = \begin{cases} 0 & \text{for } x \leq s_1 \\ \frac{x - s_1}{s_2 - s_1} & \text{for } s_1 \leq x \leq s_2 \\ 1 & \text{for } x \geq s_2 \end{cases}$$



- No dependence on details of the threshold modeling

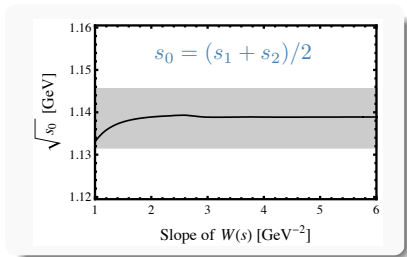
Sensitivity to threshold modeling

- replace the Heaviside step function with a ramp function:

$$R(s) = R_\rho(s) \theta(s_2 - s) + R_c(s) W(s)$$

with the weight function $W(s)$

$$W(x) = \begin{cases} 0 & \text{for } x \leq s_1 \\ \frac{x - s_1}{s_2 - s_1} & \text{for } s_1 \leq x \leq s_2 \\ 1 & \text{for } x \geq s_2 \end{cases}$$



- No dependence on details of the threshold modeling

Finite energy sum rules at finite density

Medium-modification of the sum rules

- ⊙ The existence of nuclear matter causes breaking Lorentz invariance:
 - ▶ two invariant correlator: longitudinal and transverse parts.
 - ▶ choosing a preferred reference frame of the medium ($\mathbf{q} = 0$), longitudinal and transverse correlators coincide.

$$\Pi^L(\omega, \mathbf{q} = 0) = \Pi^T(\omega, \mathbf{q} = 0) \equiv \Pi(\omega, \mathbf{q} = 0)$$

- ⊙ New operators with spin appear due to the broken Lorentz invariance:
 - ▶ the first moment of in-medium FESR involves twist-2 operator (e.g. $\langle \bar{q}\gamma_\nu D_\mu q \rangle$) to be considered.
- ⊙ Medium-dependence in the OPE side contributes only to the condensates:
 - ▶ non-perturbative contributions in OPE appear to be clearly separated into the condensates.
 - ▶ medium effects are non-perturbative.

Density-dependence of OPE

Hatsuda & Lee [Phys. Rev. C 46, R34 (1992)]

- Expectation value: vacuum \rightarrow ground state of nuclear matter

$$\langle 0|O|0\rangle \equiv \langle O\rangle_0 \rightarrow \langle O\rangle_{\rho_N} = \langle N|O|N\rangle$$

- In-medium coefficients: $c_n \rightarrow c_n + \delta c_n$

$$\delta c_2 = -3\pi^2 \left(\frac{4}{27} M_N^{(0)} - 2\sigma_N - A_1 M_N \right) \rho_N$$

density dependence of
gluon condensate

$$(M_N^{(0)} \approx 0.88 \text{ GeV})$$

density dependence of
quark condensate

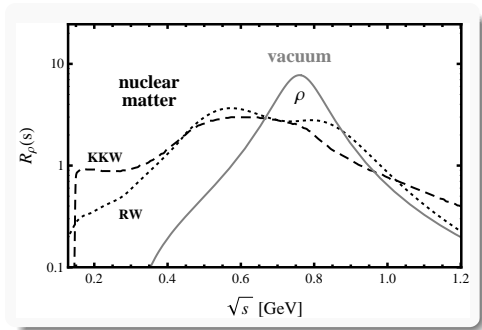
$$(\sigma_N \approx 45 \text{ MeV})$$

first moment of parton
distribution from DIS

$$(A_1 \approx 1.24)$$

Spectral functions at finite density

- ⊙ ρ -meson spectral functions in nuclear medium ($\rho_N = \rho_0 = 0.17 \text{ fm}^{-3}$):



- ▶ KKW: SU(3) chiral dynamics with vector meson dominance
Klingl, Kaiser & Weise [Nucl. Phys. A624, 527 (1997)]
- ▶ RW: particle-hole excitations ($\Delta(1232)$ -h and $N^*(1520)$ -h)
Rapp & Wambach [Adv. Nucl. Phys. 25, 1 (2000)]

Results for ρ -meson at finite densityIn vacuum: $\sqrt{s_0} \approx 1.14 \text{ GeV} \approx 4\pi f_\pi$

$$\bar{m}^2 \equiv \frac{\int_0^{s_0} ds s R(s)}{\int_0^{s_0} ds R(s)}$$

In-medium KKW spectrum:

$$\sqrt{s_0^*} \approx 1.00 \pm 0.02 \text{ GeV}$$

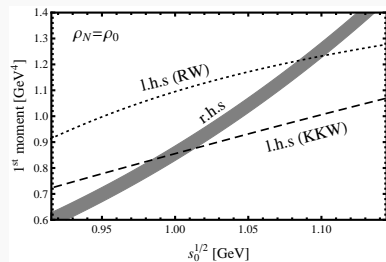
$$\sqrt{\frac{s_0^*}{s_0}} = \frac{f_\pi^*}{f_\pi} \approx 0.87 \approx \frac{\bar{m}^*}{\bar{m}} : \text{BR-scaling}$$

In-medium RW spectrum:

$$\sqrt{s_0^*} \approx 1.09 \pm 0.01 \text{ GeV}$$

$$\sqrt{\frac{s_0^*}{s_0}} \approx \frac{\bar{m}^*}{\bar{m}} \approx 0.96$$

Kwon, Procura & Weise [PRC 78, 055203 (2008)]



Finite energy sum rules at finite temperature

Temperature-dependence of OPE

Hatsuda, Koike & Lee [Nucl. Phys. B 394, 221 (1993)]

- Thermal expectation value:

$$\langle O \rangle_0 \rightarrow \langle O \rangle_T = \frac{\text{Tr } O \exp(-H/T)}{\text{Tr } \exp(-H/T)}$$

- In-medium coefficients: $c_n \rightarrow c_n + \delta c_n$

$$\delta c_2 = -\frac{3}{2} \left(\frac{2}{9} \mp 3 + A_1 \right) m_\pi^2 T^2 \int_{m_\pi/T}^{\infty} dy \frac{\sqrt{y^2 - \left(\frac{m_\pi}{T}\right)^2}}{e^y - 1}$$

T -dependence of
gluon condensate

T -dependence of
quark condensate

first moment of parton
distribution from DIS

Vector & axialvector mixing with temperature

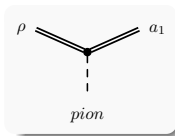
- ⊙ Mixing of vector and axialvector:

$$R_V(s, T) = R_V(s, 0) (1 - \epsilon(T)) + R_A(s, 0) \epsilon(T)$$

$$R_A(s, T) = R_A(s, 0) (1 - \epsilon(T)) + R_V(s, 0) \epsilon(T)$$

Eletsky & Ioffe [PRD 47, 3083 (1993), PRD 51, 2371 (1995)]

- ▶ the mixing parameter $\epsilon(T)$ is given by the thermal pion loop:



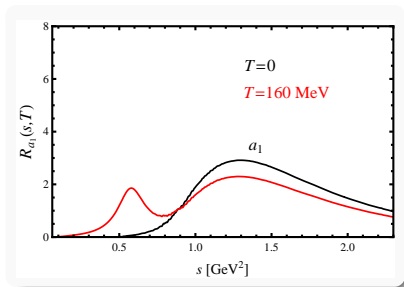
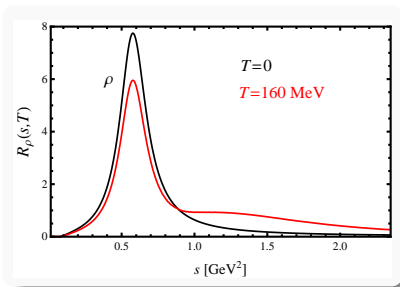
$$\epsilon(T) = \frac{2}{f_\pi^2} \int \frac{d^3k}{\omega (2k)^3} \frac{1}{e^{\omega/T} - 1} \xrightarrow{m_\pi \rightarrow 0} \frac{T^2}{6 f_\pi^2}$$

$$\text{where } \omega^2 = k^2 + m_\pi^2 .$$

- ▶ At critical temperature where $\epsilon \simeq \frac{1}{2}$, R_V and R_A become identical.

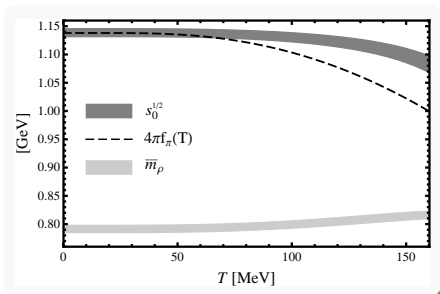
Mixing of finite-width spectrum

- ⊙ Spectral functions with finite decay width:



Mixing of finite-width spectrum

- Sum rule result for vector channel:



- Average ρ -meson mass:

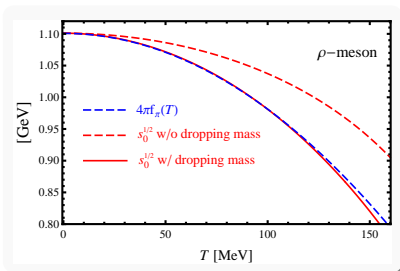
$$\bar{m}_\rho^2 = \frac{\int_0^{s_0} ds s R_\rho(s)}{\int_0^{s_0} ds R_\rho(s)}$$

- Comparison with ChPT:

$$f_\pi(T) = f_\pi \left(1 - \frac{1}{2} \epsilon(T) \right)$$

Simple test beyond V - A mixing

- ⊙ Dropping pole mass in addition to the V - A mixing:



The simplest ansatz (zero width):

$$R_\rho(s, 0) = F_\rho^2 \delta(s - m_\rho^2)$$

$$R_a(s, 0) = F_a^2 \delta(s - m_a^2)$$

$$R_\rho(s, T) = R_\rho(s, 0) (1 - \epsilon) + R_a(s, 0) \epsilon$$

Brown-Rho scaling hypothesis:

$$m_\rho^2 \rightarrow m_\rho^2 \left(1 - \frac{1}{2} \epsilon(T)\right)^2$$

⇒ better agreement : $\sqrt{s_0} = 4\pi f_\pi(T)$

About four-quark condensates

- Sum rules for 0th and 1st moments: RHS quantities are accurately determined (pQCD and leading condensates)
- Sum rules for 2nd moment: involving four-quark condensates

$$\int_0^{s_0} ds s^2 R(s) = \frac{s_0^3}{3} + c_3$$

$$c_3 = -6\pi^3 \alpha_s \left[\langle (\bar{u}\gamma_\mu\gamma_5\lambda^a u - \bar{d}\gamma_\mu\gamma_5\lambda^a d)^2 \rangle + \frac{2}{9} \langle (\bar{u}\gamma_\mu\lambda^a u + \bar{d}\gamma_\mu\lambda^a d) \sum_{q=u,d,s} \bar{q}\gamma^\mu\lambda^a q \rangle \right]$$

- Ground state saturation ($\kappa = 1$)

$$\langle (\bar{q}\gamma_\mu\gamma_5\lambda^a q)^2 \rangle = -\langle (\bar{q}\gamma_\mu\lambda^a q)^2 \rangle = \frac{16}{9} \kappa \langle \bar{q}q \rangle^2$$

valid approximation? \Rightarrow Always $\kappa > 3$ and large uncertainties.

Summary

- ⊙ The sum rules for the **lowest two moments** of the ρ -meson spectral function involve perturbative contributions and only leading condensates as small corrections:
accuracy both in vacuum and in medium
- ⊙ Chiral gap scale: $4\pi f_\pi$ meaningful both in vacuum and in-medium.
- ⊙ For broad spectral distributions, “mass shift” vs. “broadening” discussion must be specified in terms of first moment.
- ⊙ Brown-Rho scaling as a statement involving the lowest two moments in the window of low-mass enhancement.
- ⊙ Further step: extension to **nonvanishing three-momentum**.

Thank you for your attention!