

Development of VUV-reflecting mirrors and wavelength shifting films for the CBM-RICH detector

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Materials science for heavy ion physics

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- aim:** electron identification for momenta below 8-10 GeV/c
→ high efficiency, large acceptance, 10^4 combined π -suppr. with TRD
- concept:** gaseous RICH detector
stable, robust
limited R&D efforts
rely to a large extend on components from industry
not too expensive
- people:** Pusan Natl. Univ. (I.K. Yoo et al) – gas system, RICH prototype
PNPI, St. Petersburg (V. Samsonov et al.) – mechanics
IHEP Protvino (S. Sadovsky et al.) – PMT development
HS Esslingen (M. Dürr) – mirror development/ investigations, WLS films
Wuppertal (applied for funding) – photodetector development
GSI (C. Höhne et al.) – concept, simulations, layout, coordination,
photodetector + WLS films R&D, readout electronics

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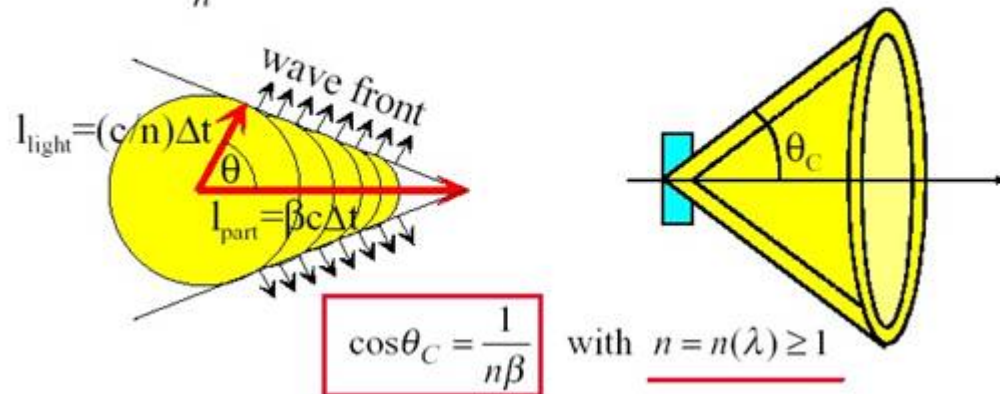
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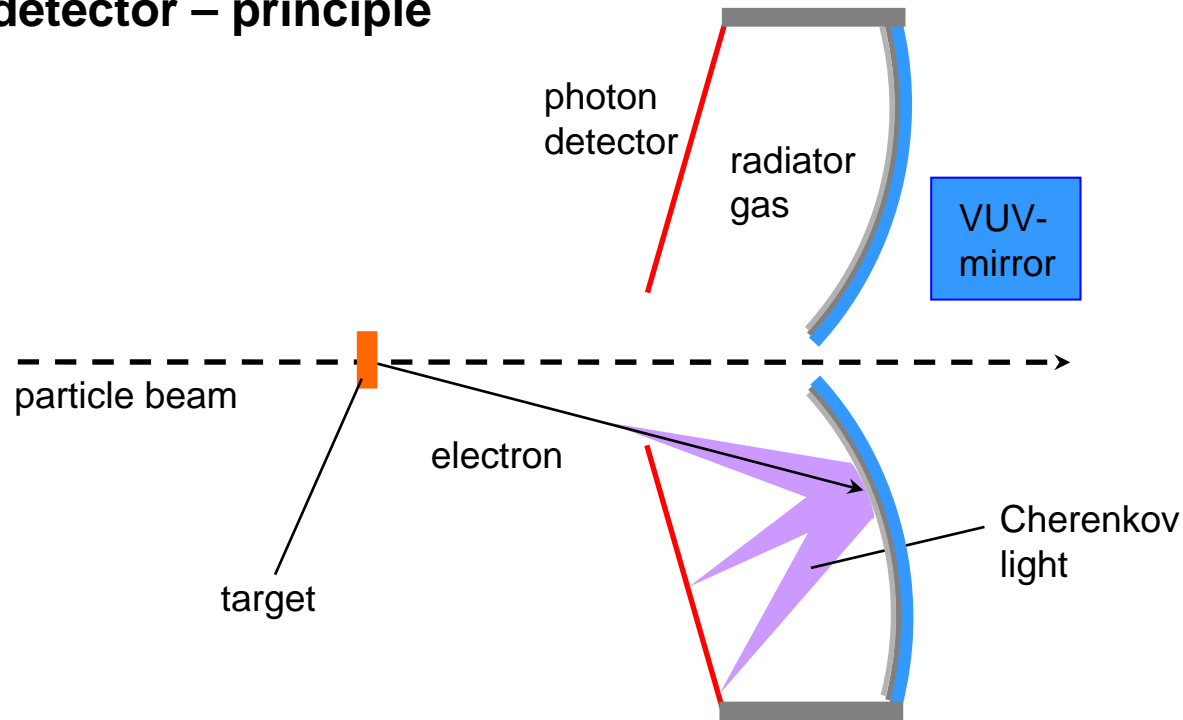
RICH detector – principle

- particles traversing matter with a velocity larger than the velocity of light in that medium (refractive index n) emit Cherenkov radiation
- opening angle of light cone $\cos \theta = \frac{1}{n\beta}$

$$\beta \geq \beta_{thr} = \frac{1}{n} \quad n: \text{refractive index}$$



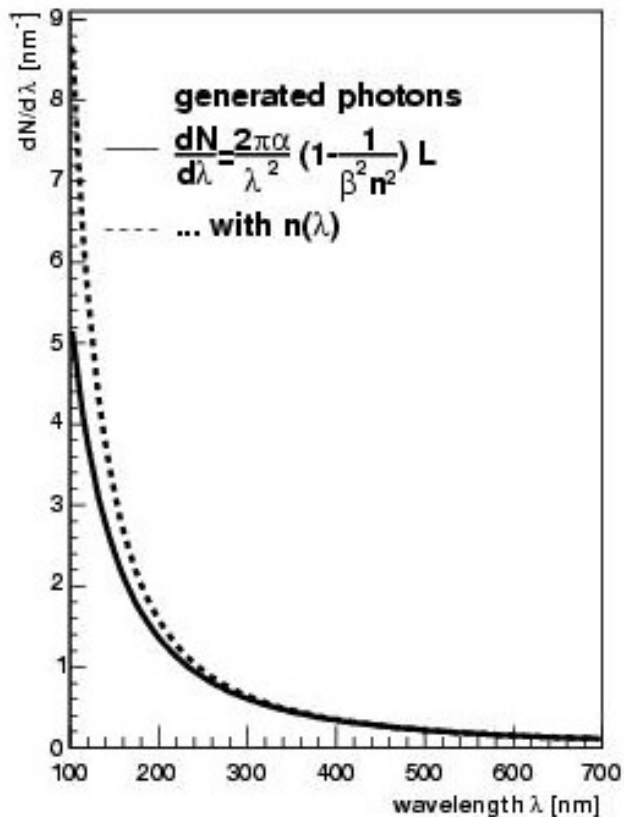
RICH detector – principle



- light cone of cherenkov radiation is projected onto the focal plane where the cone is imaged as rings
- focal plane at $r_0/2$ (r_0 : radius of curvature of mirror)
- ring radius depends on θ

Cherenkov spectrum - wavelength dependence

Spectrum for nitrogen:



- number of photons per ring N_γ
(full e-ring, $r \approx 5.4$ cm)

$$N_\gamma = \int \frac{dN}{d\lambda} d\lambda$$

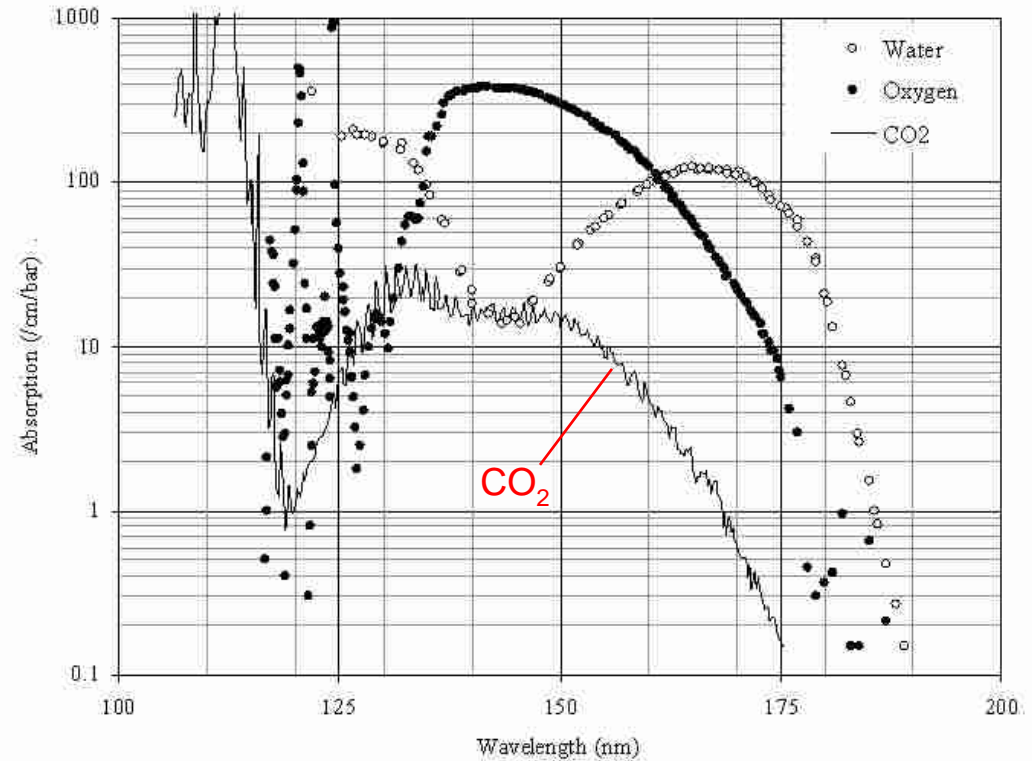
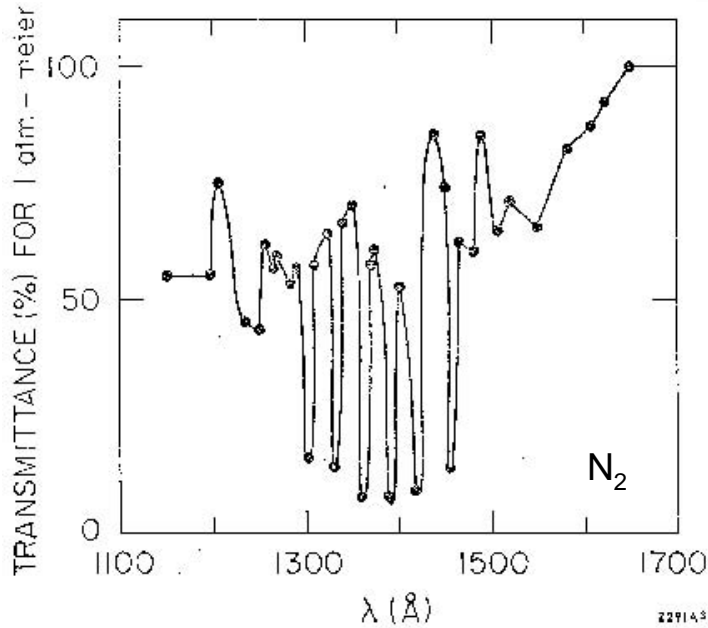
⇒ Short wavelength photons are important

Radiator gas

- as the RICH detector should mainly serve for electron identification, the pion threshold for emission of Cherenkov light should be sufficiently high in momentum

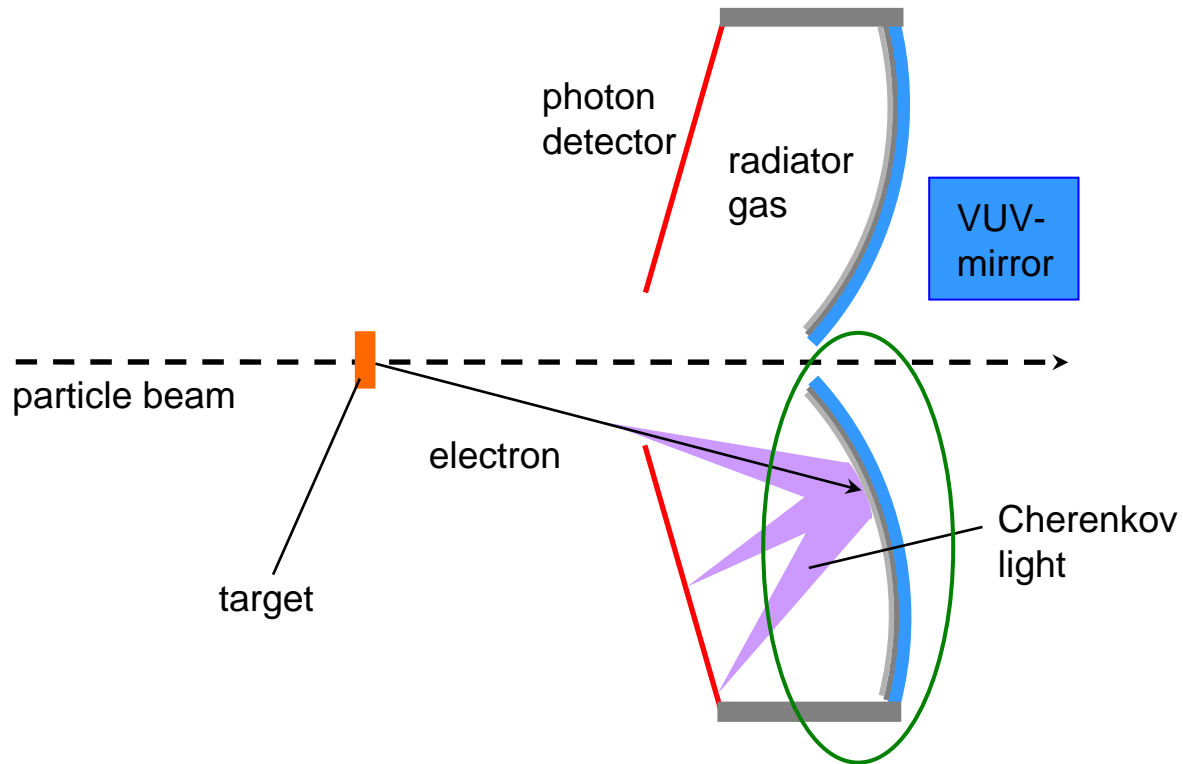
radiator	n	γ_{th}	p_{th}^{π} [GeV/c]	λ_{thresh} [nm]	radiation length [m]	handling?
N ₂	1.000298	41	5.6	~ 150	304	✓
Ar	1.00044	33.6	4.7	~ 130	650	☹
CO ₂	1.00045	33.3	4.65	~ 175	183	✓
CF ₄	1.000488	32	4.47	< 120		chemically aggressive
N ₂ O	1.000509	31.4	4.37			(toxic)
CH ₃ OH (methanol)	1.000546	30.3	4.2			flammable
C ₂ H ₆ (ethane)	1.000706	26.6	3.71	~ 160	340	

Radiator gas - absorption edges

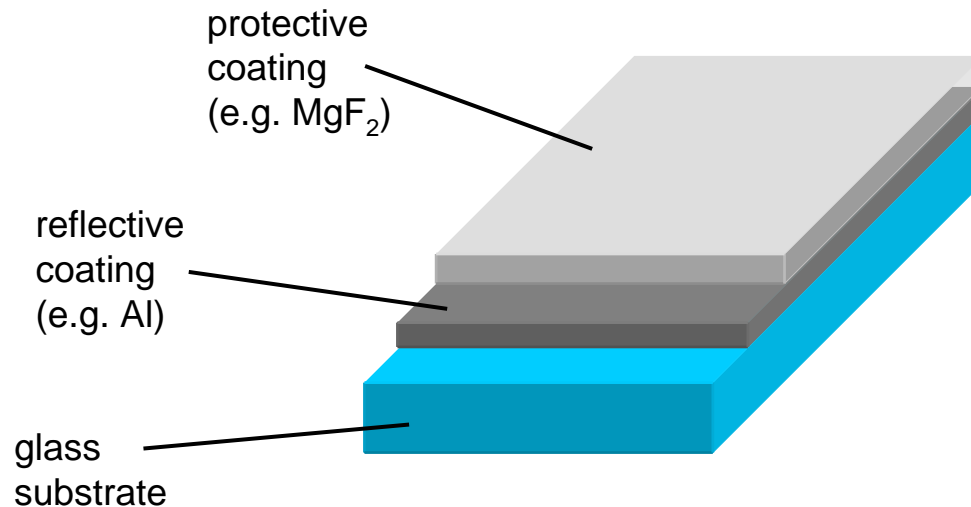


- N₂ absorption edge ~ 150 nm
- CO₂ absorption edge ~ 175 nm

RICH detector – principle



RICH detector – mirror and coatings



- Al has a good reflectivity in the visible and UV spectrum
- protective coating is needed because of Al_2O_3 's absorption edge in the UV

Requirements for RICH mirrors

Physical:

- high reflectivity down to ~ 160 nm
- good surface homogeneity: small error of radius
- moderate material budget (rough number: $< 2.5\%$ X_0 including mirror support, compare to 3 mm float glass: $X_0 = 2.4\%$; to be investigated in detail)

Technical:

- stiffness and precision also with support structure and in support position

Economical

- moderate costs

Our approach

- minimize own R&D
- cooperate with companies as much as possible for mirror substrates and coating

Recent RICH installations and used mirror technology

- HADES: first mirrors from FLABEG, Germany, coating done at TU München
- RICH1 LHCb: carbon fiber mirrors from CMA, USA, coating at SISO, France
- RICH2 LHCb: mirrors (Pyrex) from Compas, Czech Republic, coating at CERN

Companies possibly to cooperate with

- Flabeg, Germany (mirrors)
- Compas, Czech Republic (mirrors)
- Siso, France (coating)

1st trial: Flabeg

- can do now both mirrors and coating
- ⇒ test coating for reflectivity measurements

Specs



photography

dimensions:

Size: $A = 400 \times 400 \text{ mm}^2$

glass: $d = 6 \text{ mm}$

Radius: $r_0 = 3200 \text{ mm}$

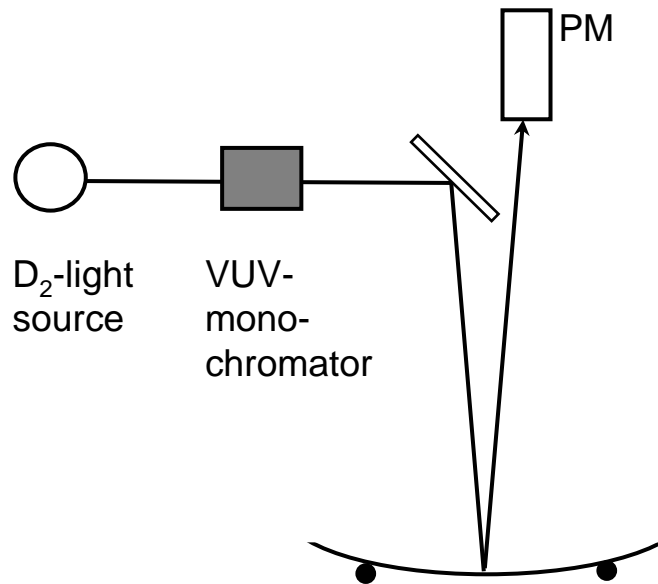
coating:

Al: $d = 70 \text{ nm}$

MgF₂: $d = 90 \text{ nm}$

Set up

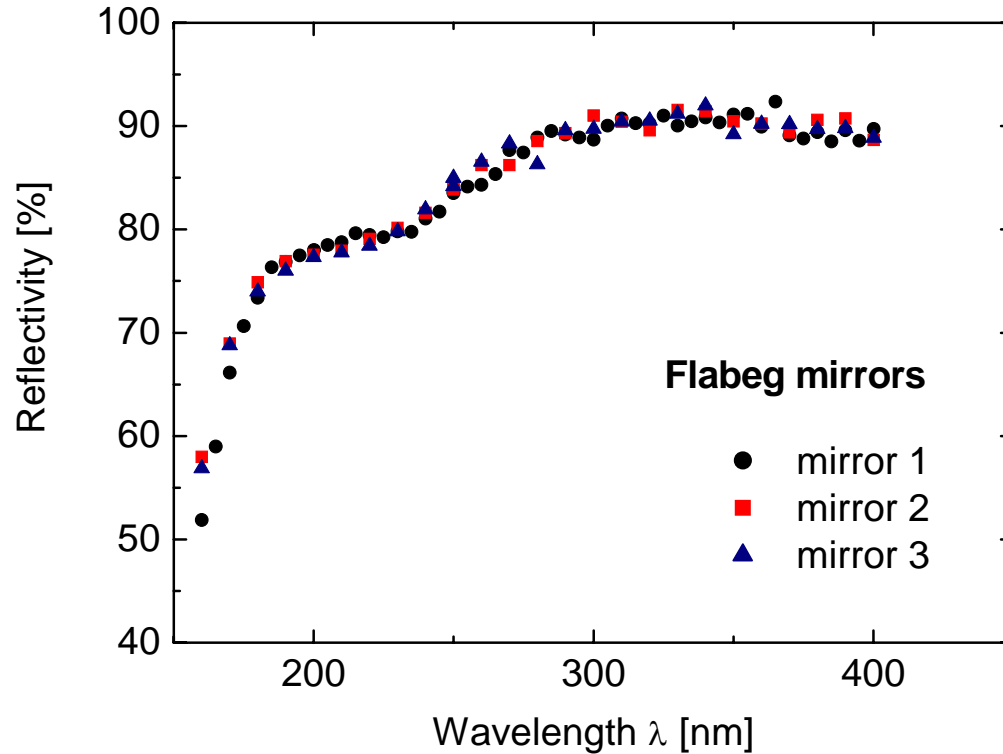
Schematics:



Apparatus at CERN – A.Braem:

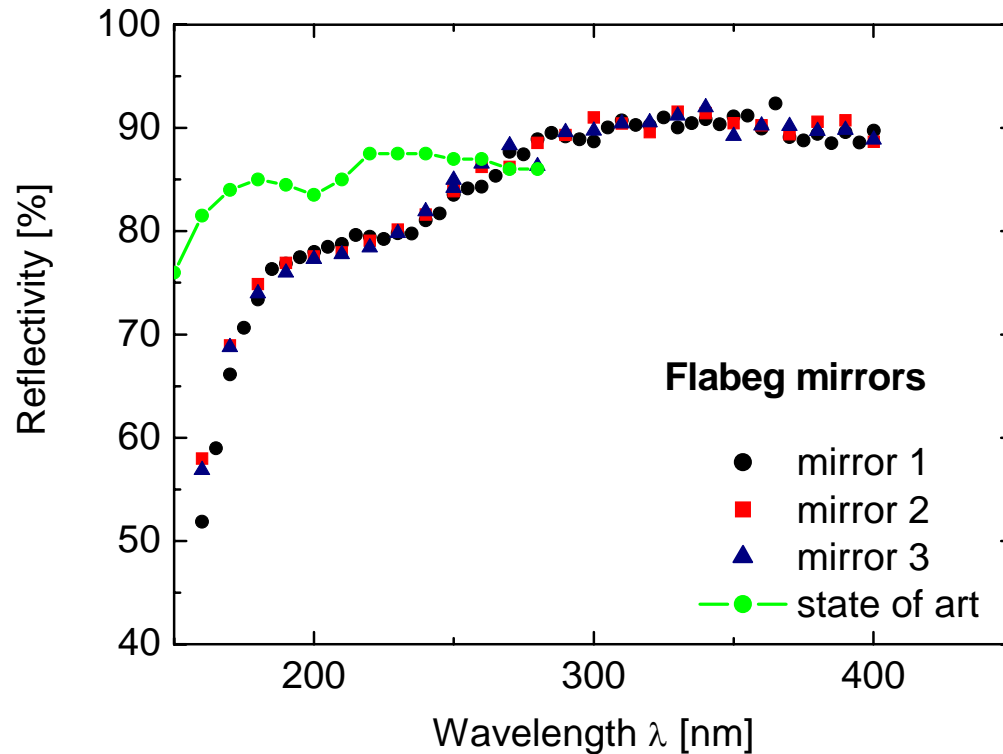


Results



- Very good reflectivity between 400 nm and 270 nm
- First drop around 250 nm
- Second drop at about 180 nm

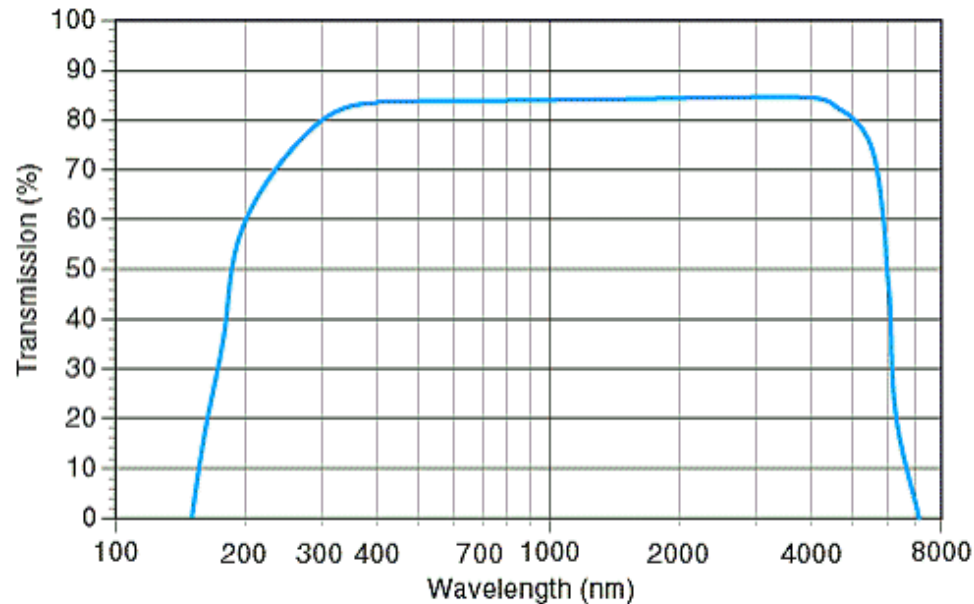
Results – comparison with state of the art



- Very good reflectivity between 400 nm and 270 nm
- First drop around 250 nm
- Second drop at about 180 nm
- state-of-the-art data courtesy of A. Braem

Influence of aluminum oxide

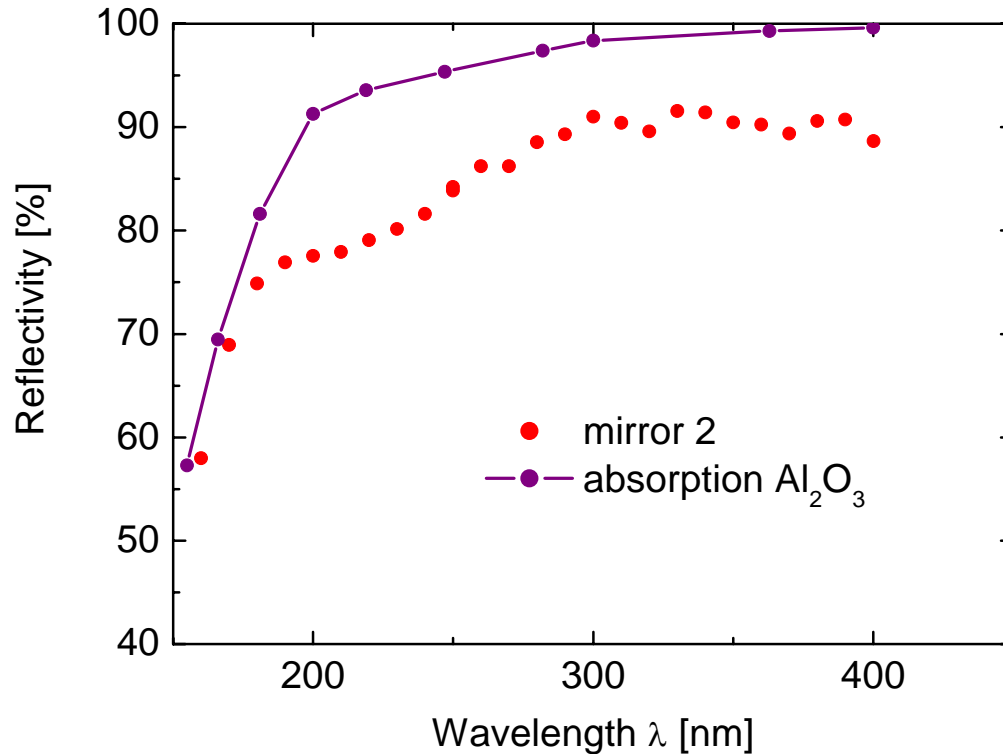
Transmission of Al_2O_3 bulk material



Reason 1:

- absorption of Al_2O_3 formed during the process (absorption edge of bulk material at 200 nm)

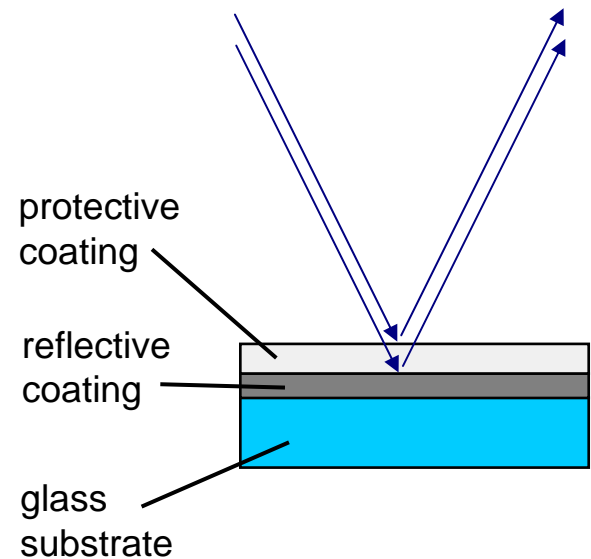
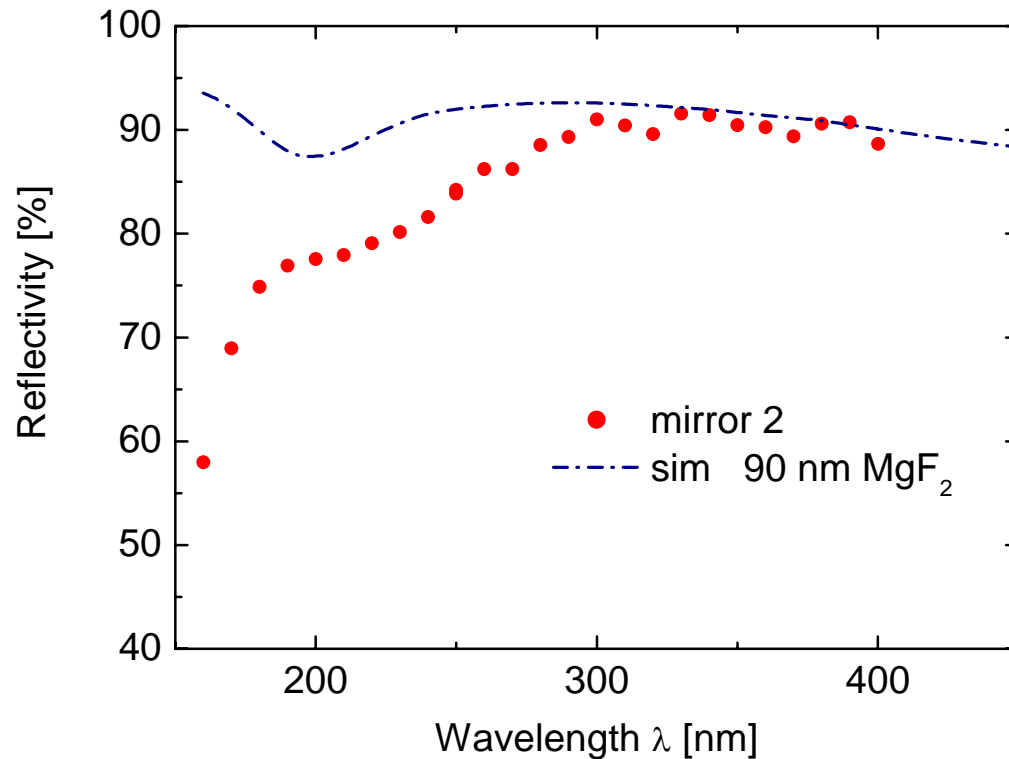
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Influence of MgF_2 layer

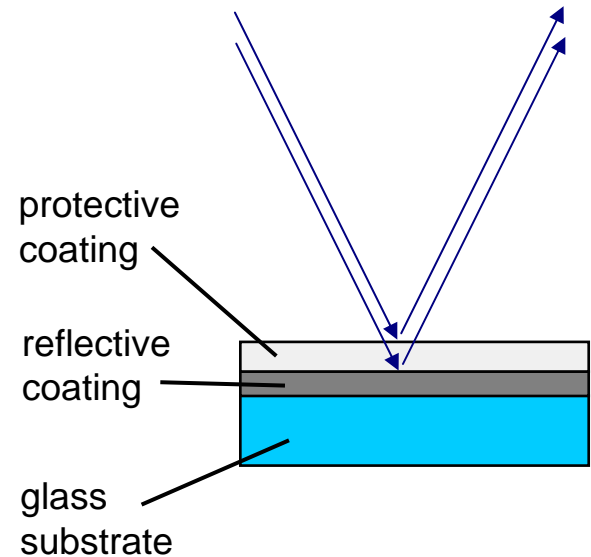
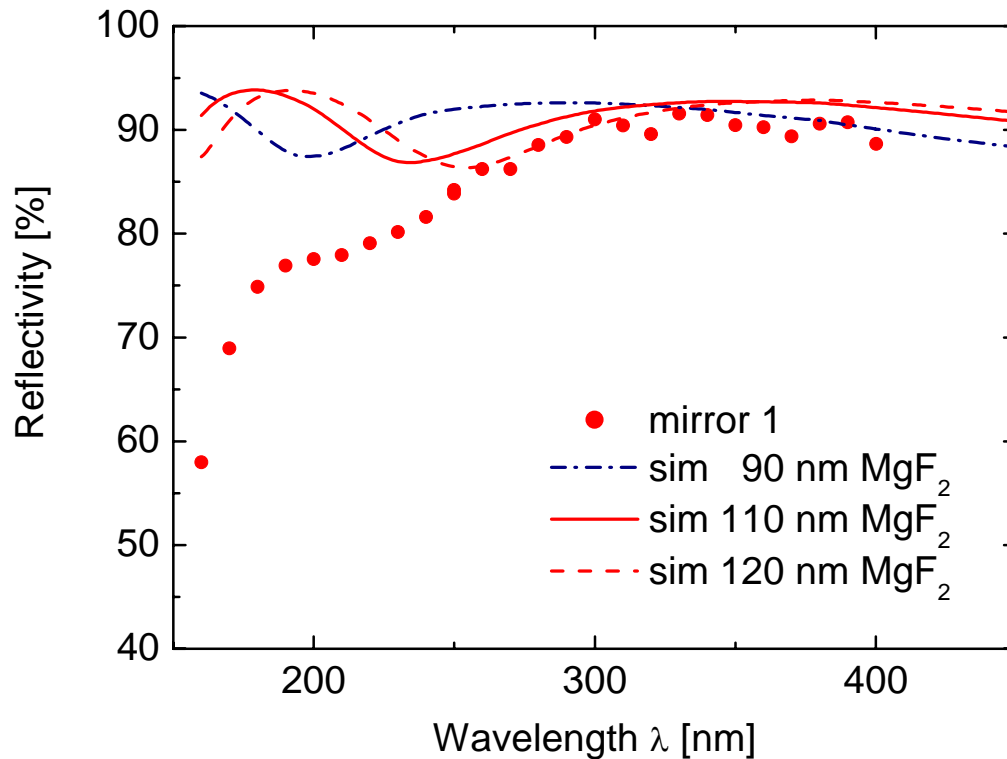


Simulations: A. Braem

Reason 2:

- interference of light reflected at Al-layer with light reflected at MgF_2 -surface

Influence of MgF₂ layer

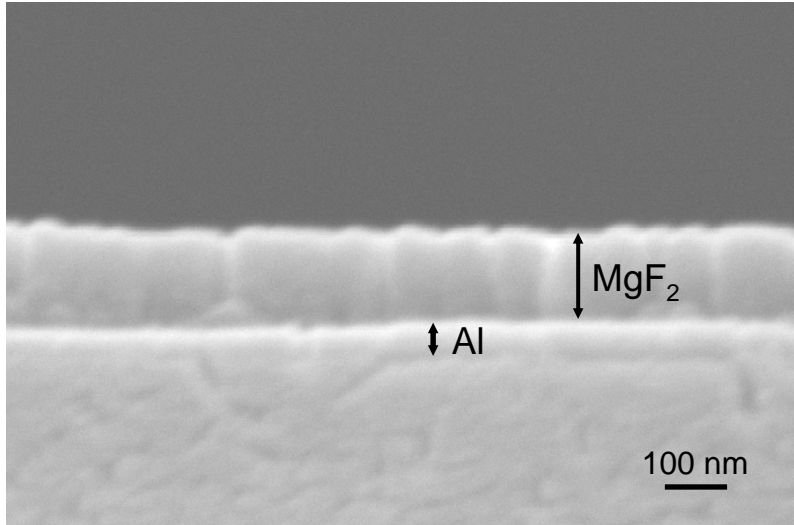


Simulations: A. Braem

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- interference of light reflected at Al-layer with light reflected at MgF₂-surface

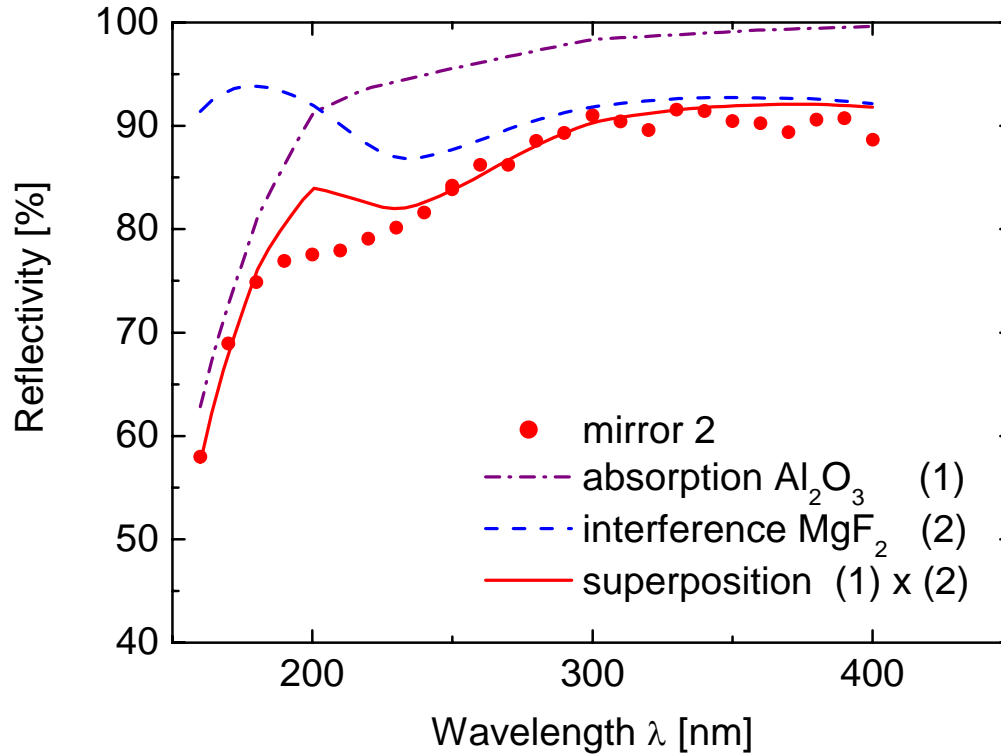
Scanning electron microscopy



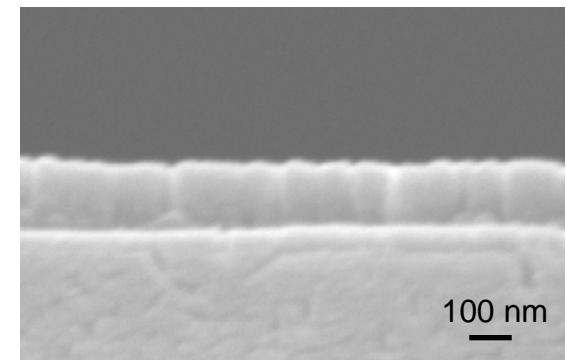
Layer thickness:

- Al-layer: $d \cong 55 \text{ nm}$
- MgF_2 -layer: $d \cong 120 \text{ nm}$

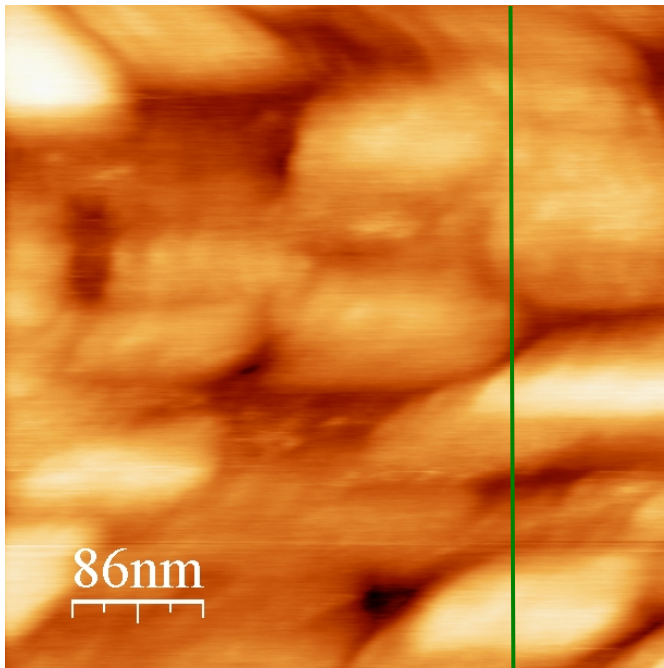
Influence of aluminum oxide and MgF₂ layer



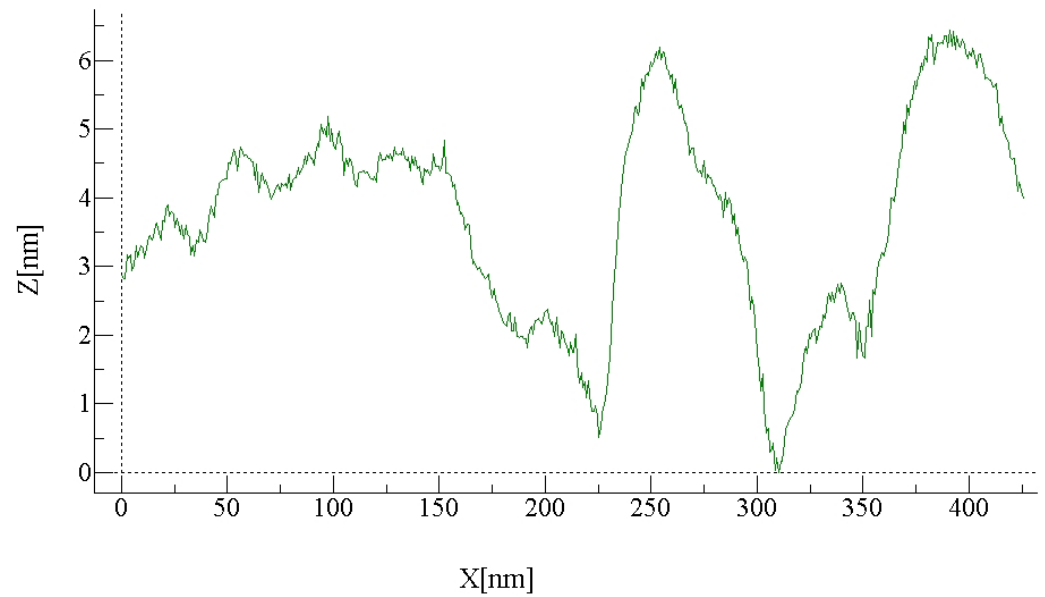
- good overall representation
- still some discrepancies
- influence of surface roughness?



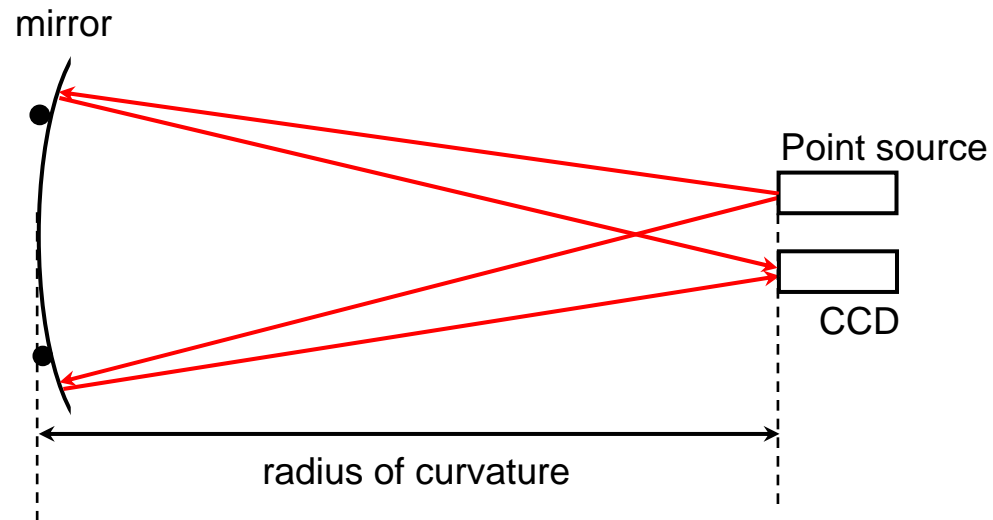
Atomic Force Microscopy



Line scan:

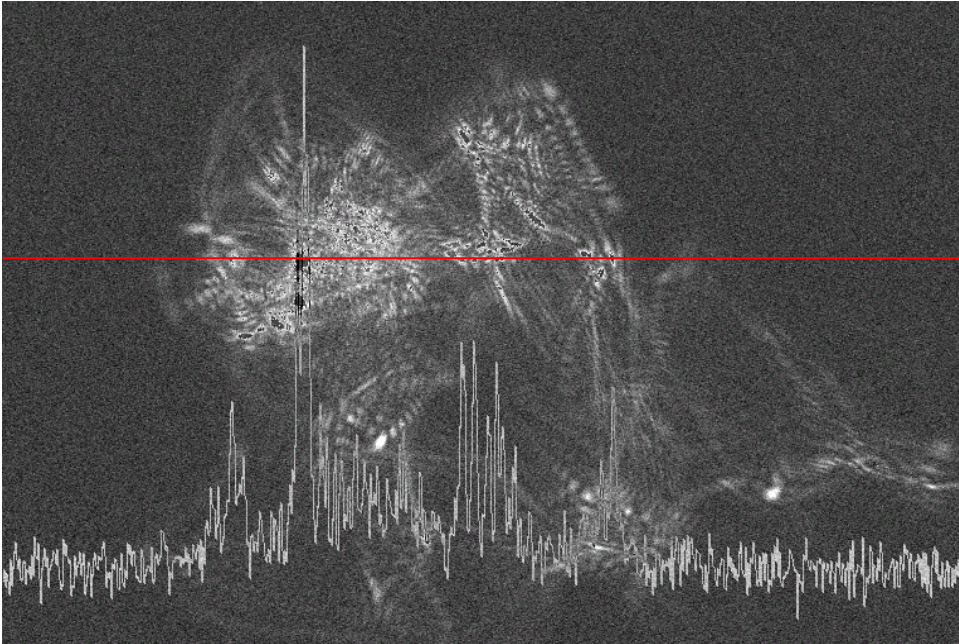


Radius of curvature – measurement set up



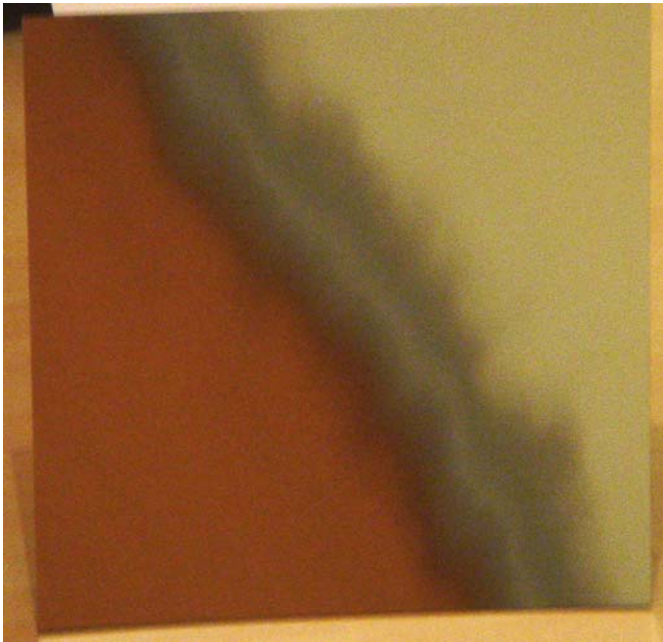
at CERN – Carmelo D'Ambrosio

Results



- very broad feature,
**most of the intensity
in the background**

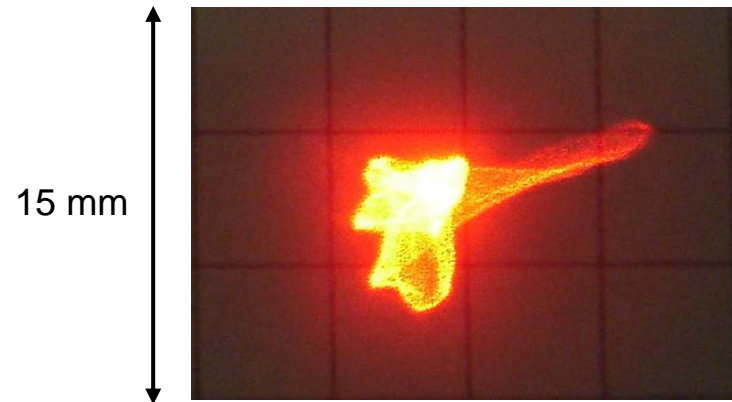
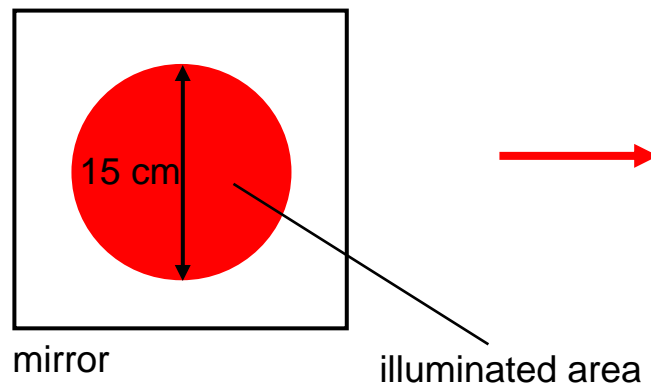
Results



- very broad feature, most of the intensity in the background
- pronounced irregularities on the cm-scale
- possible explanations:
 - rather low-cost fabrication process
 - rather thick glass substrate

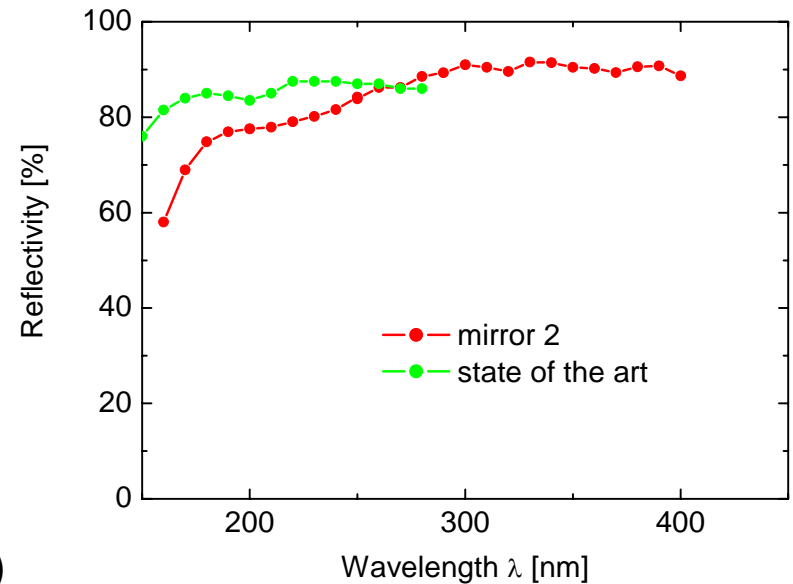
Different substrates

- some more substrates were delivered
- thickness between 3 mm and 6 mm
- preliminary measurements of D_0 did show minor improvements
- Example:



Reflectivity – coating

- good reflectivity for a „first-shot“ trial
- improvements expected from
 - higher evaporation rate
 - better base pressure
(both feasible at Flabeg)



Radius of curvature – glass substrate

- low homogeneity of the surface on the cm-scale
- minor improvements from thinner glass and changed fabrication process

Companies possibly to cooperate with

- Flabeg, Germany (mirrors) ✓
- Compas, Czech Republic (mirrors)
- Siso, France (coating)

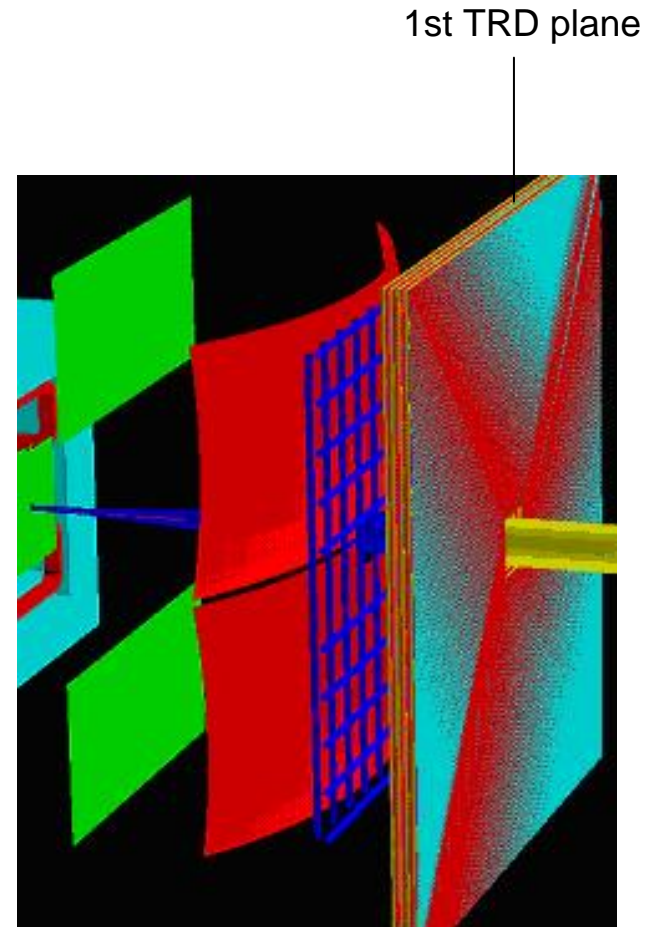
Compas

- delivered substrates for LHCb, can do coating themselves

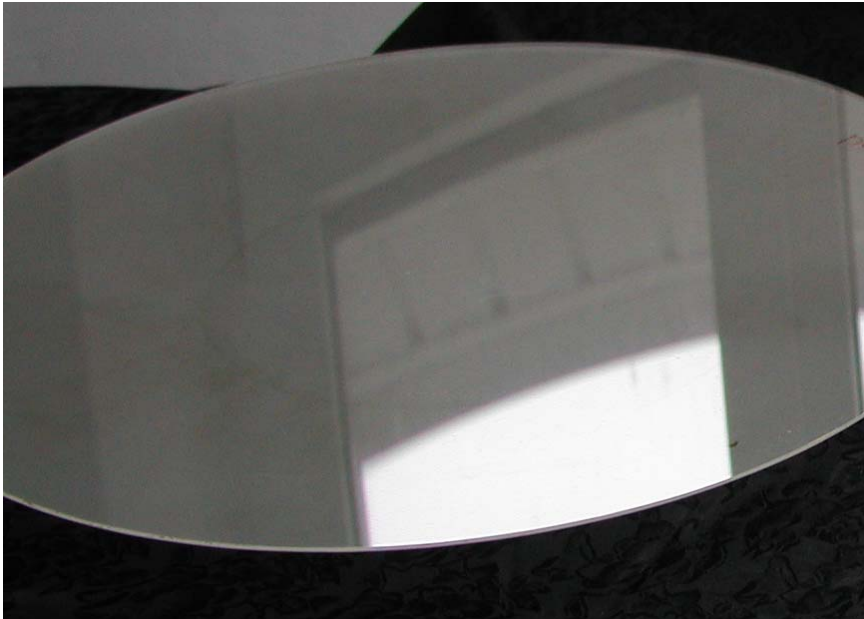
RICH detector – material budget

- 2.9 m radiator (nitrogen) = 0.95 % X_0
- entrance/ exit window from kapton foil ≤ 0.5 % X_0
- radiation length of glass 12-14 cm
- radiation length of Al 8.9 cm
- material budget of 6 mm glass = 4.6 % (e.g. RICH2 @ LHCb)
- typically: material budget of support ≤ 50 % of mirror but non-uniformly distributed

⇒ material budget of RICH detector mainly concentrated in mirror + support



Specs according to simulations at that time



photography

dimensions:

Size: $R = 300 \text{ mm}$

glass: $d = 3 \text{ mm}$

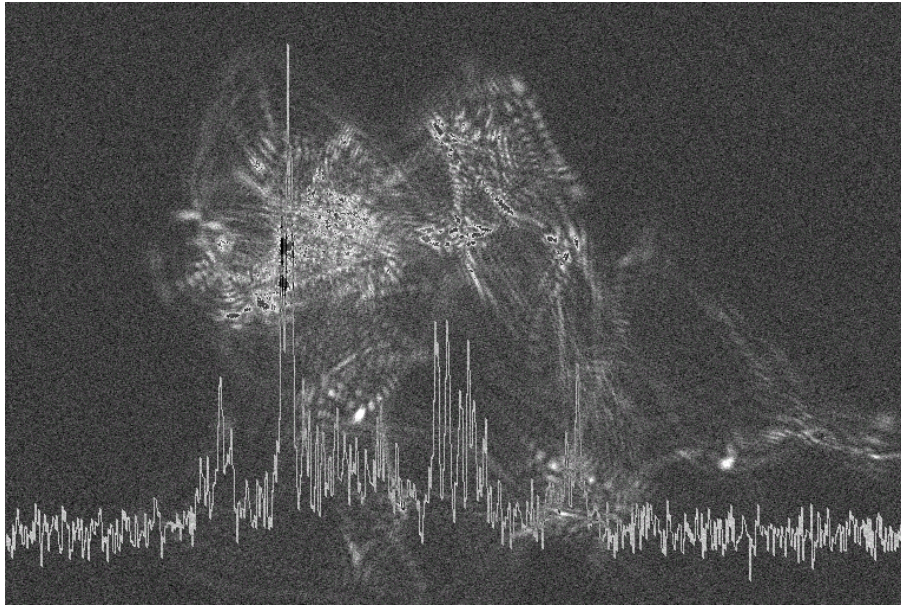
Radius: $r_0 = 3000 \text{ mm}$

coating:

Al: $d = ?$

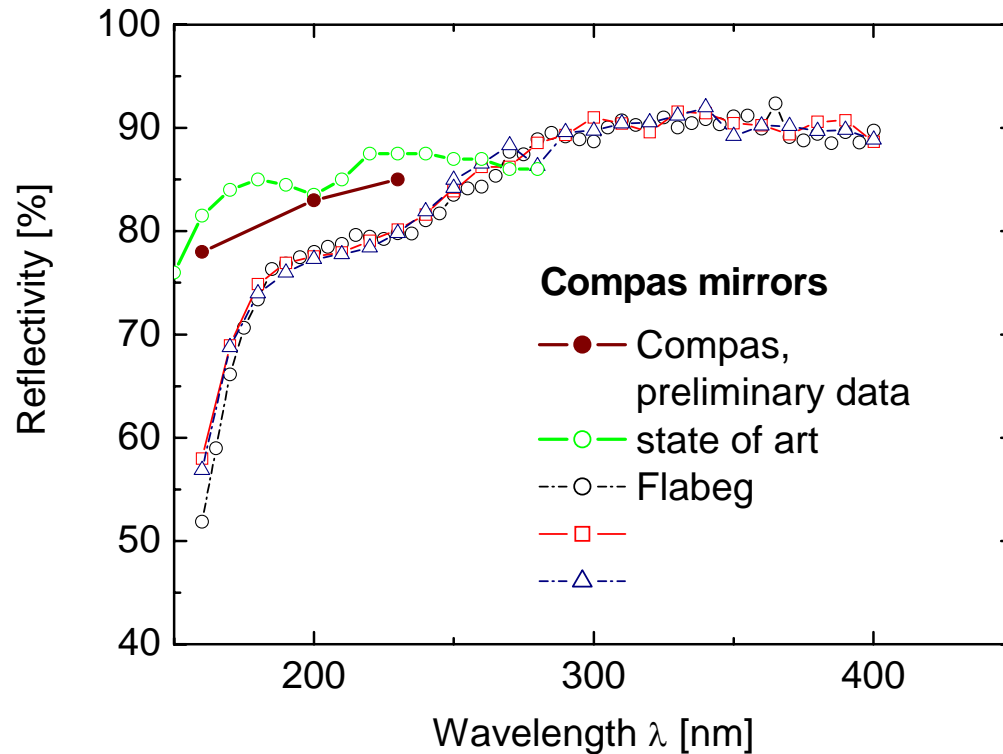
MgF_2 : $d = ?$

Results



- $D_0 \cong 2 \text{ mm}$ (90 % intensity)

Results – comparison with Flabeg and state of the art

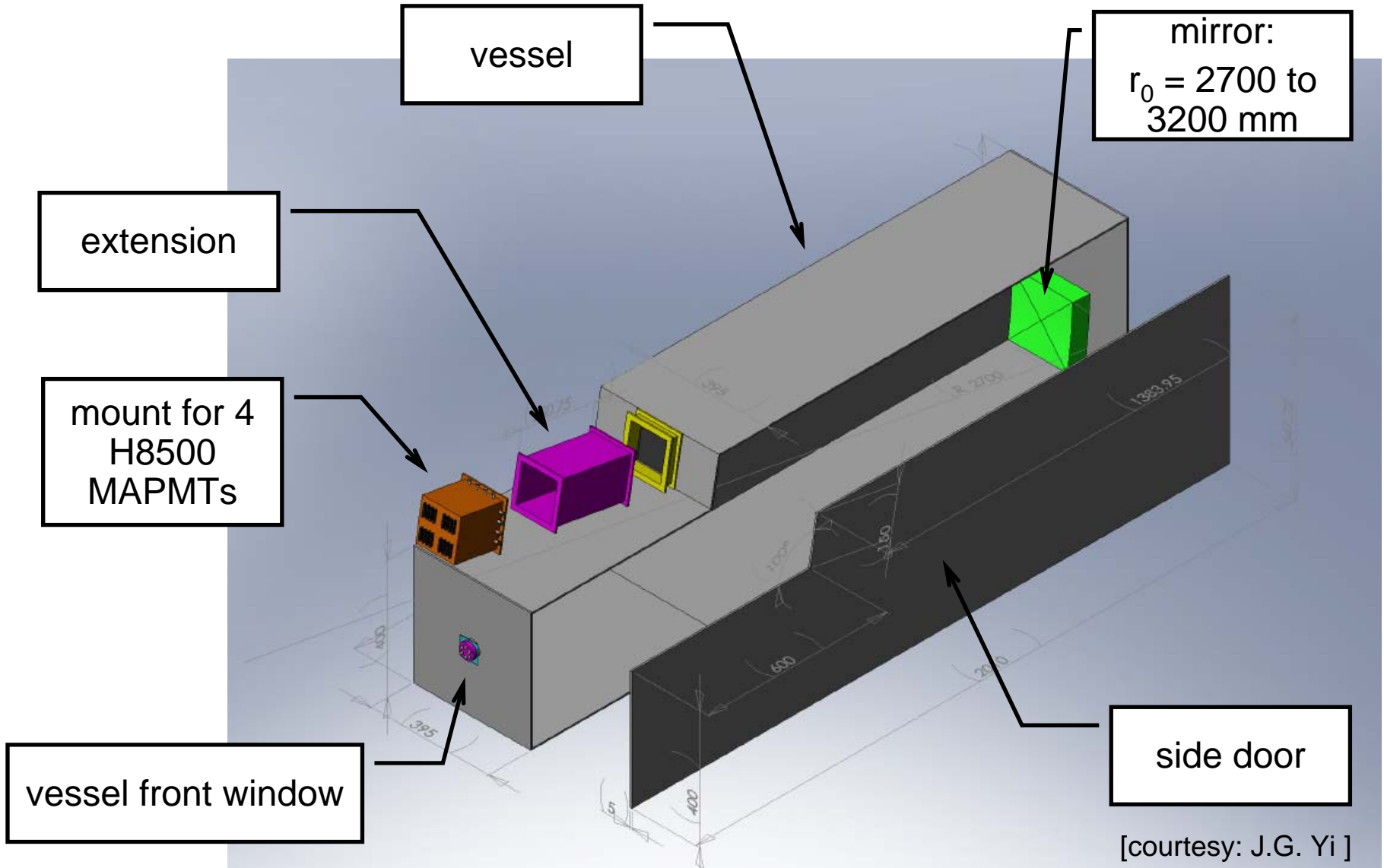


- good reflectivity in the UV-region
- more data points needed
- to be confirmed by CERN measurements

Results and outlook

- thickness of 3 mm and good D_0 possible
- good reflectivity (to be confirmed!)
- test in Pusan's mini-RICH?
- work on design of mirror support

Outlook: RICH prototype PNU



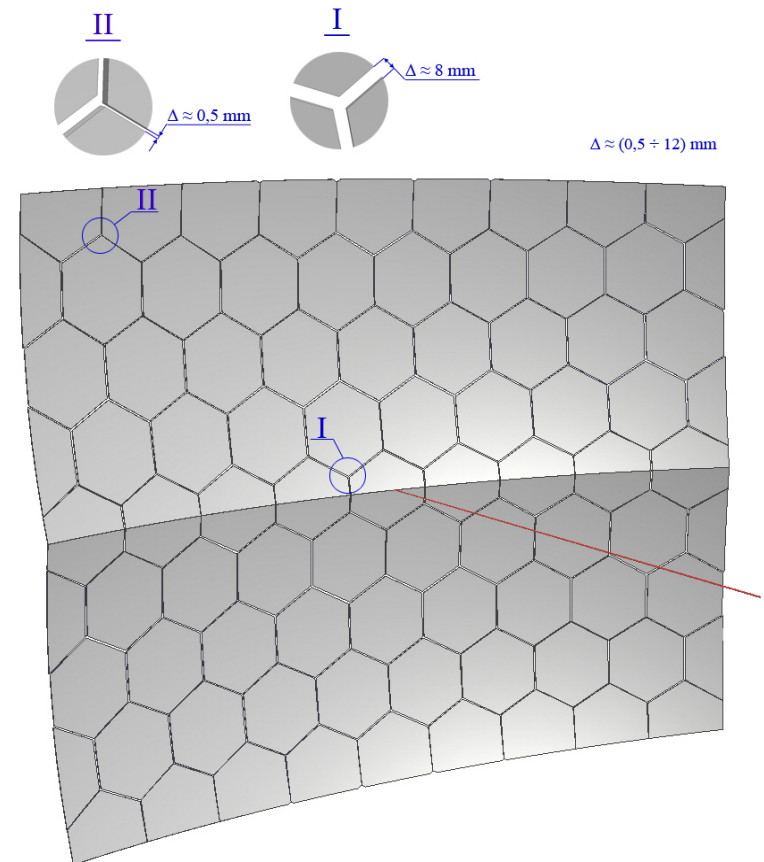
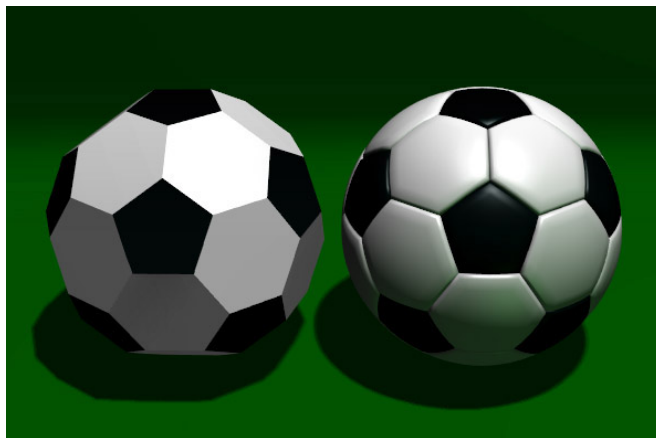
[courtesy: J.G. Yi]

Shape of tiles

Hexagonal versus quadratic:

- impossible to divide a spherical surface into hexagons exactly
- approximation by irregular gaps between them (0.5-12mm)

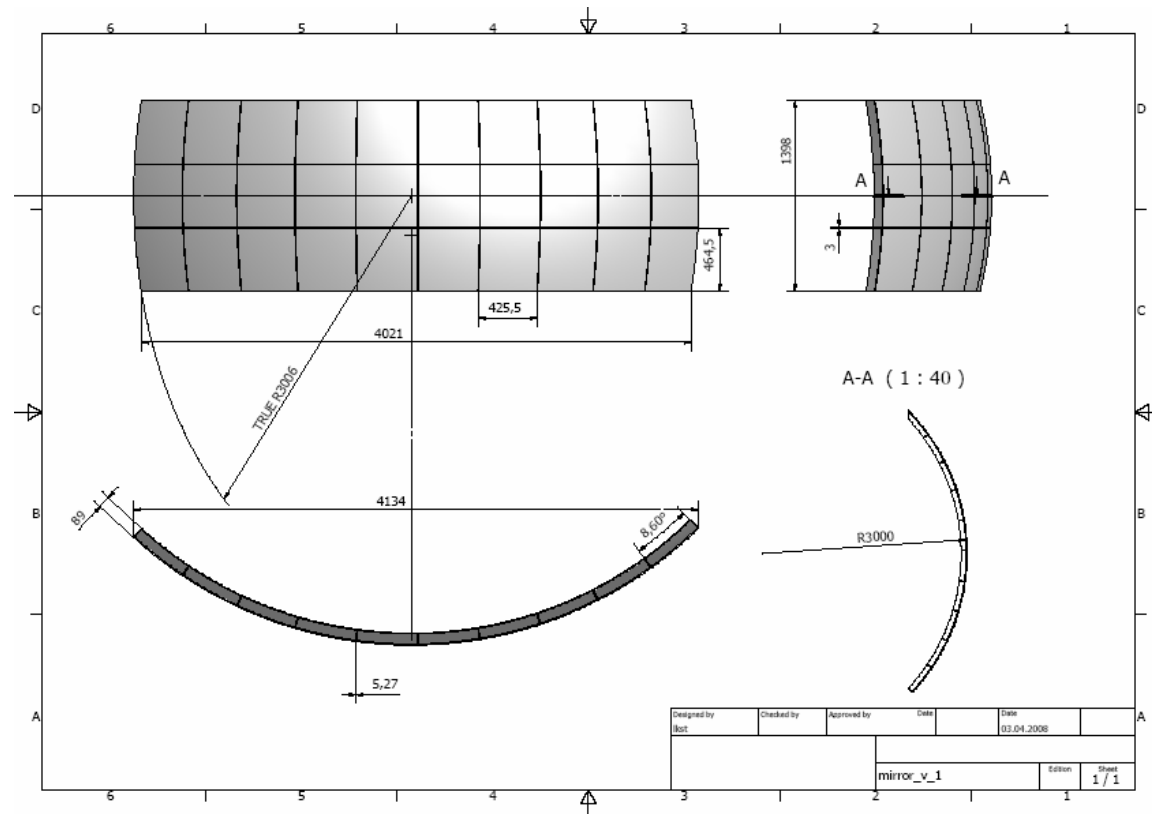
→ different size of hexagons or large gaps (due to small r_0)



[E. Vznuzdaev, PNPI St. Petersburg]

Shape of tiles

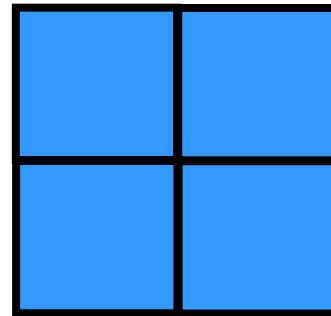
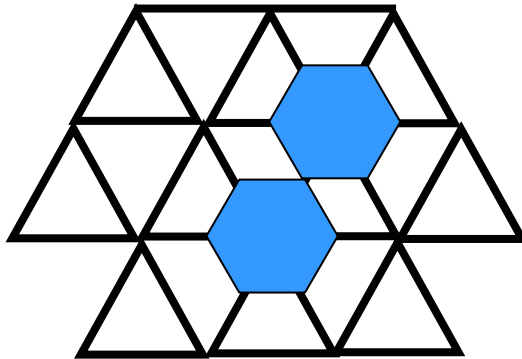
- possible solution for CBM: approx. 30 tiles of 40 x 40 cm² for each mirror half



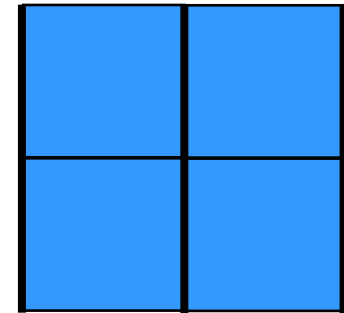
[E. Vznuzdaev, PNPI St. Petersburg]

Mirror support structure

- mirror support structure depends on shape of mirror tiles
- hexagon tiles:
one mount in center of mirror
adjustment around 2 axis
- rectangular tiles:
typically 3 adjustable mounts
(shift along z-axis possible)



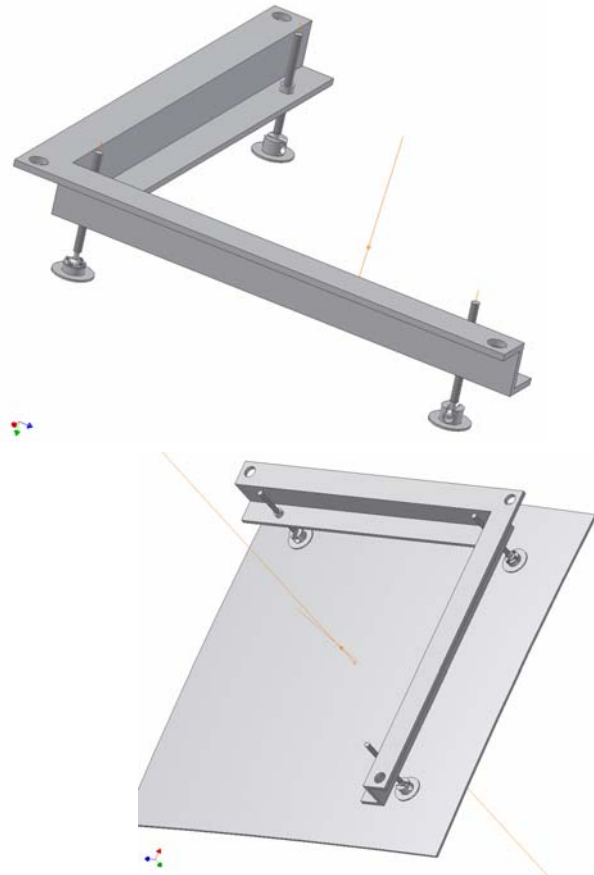
or



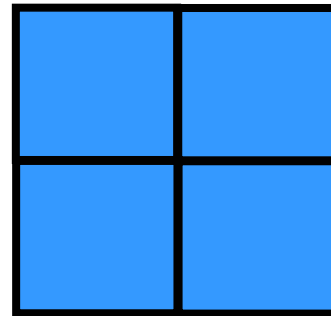
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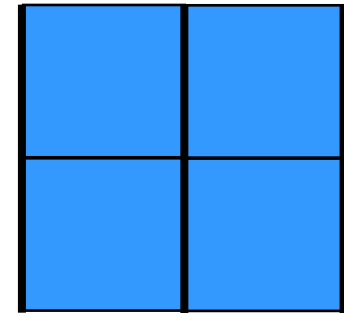
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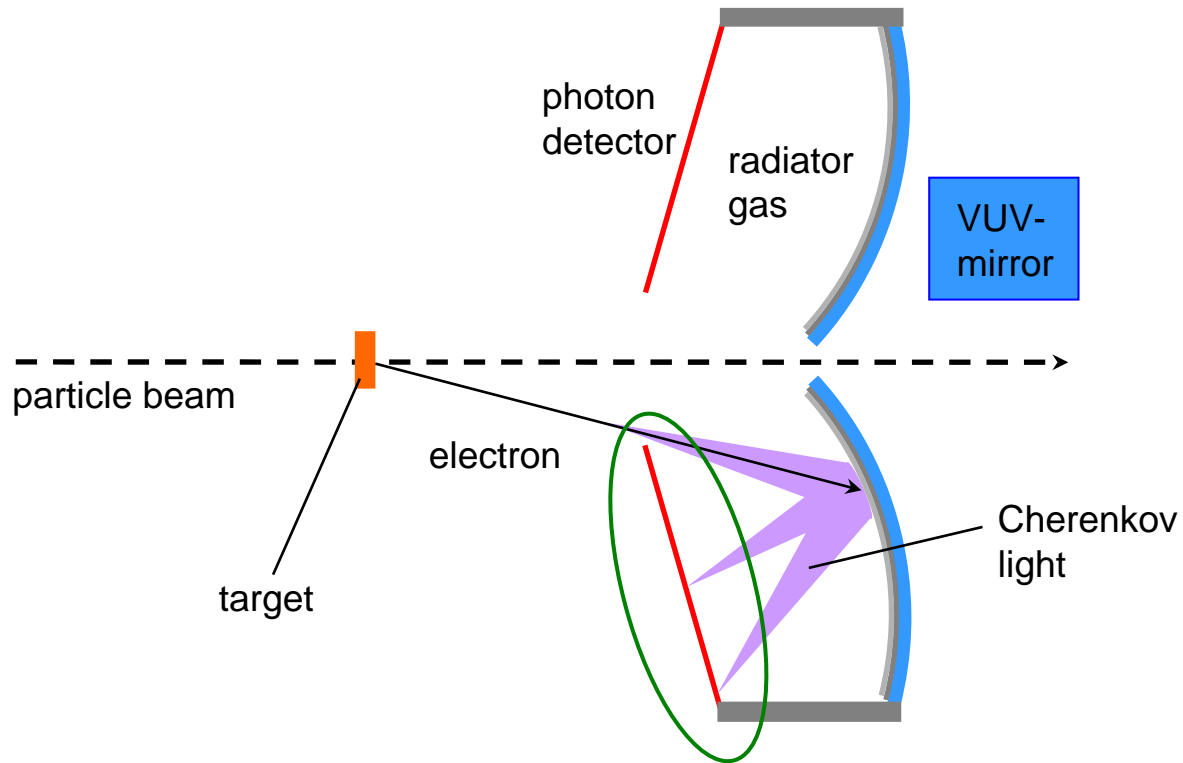


or ...

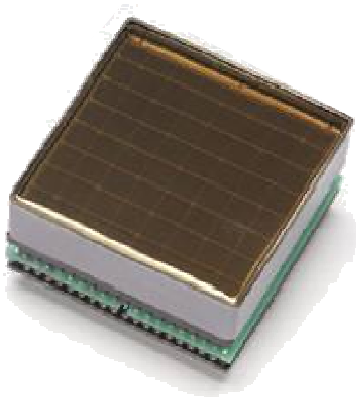
requirements:

- no additional stress on the mirrors
- stabilization against gravity (→ distortions)

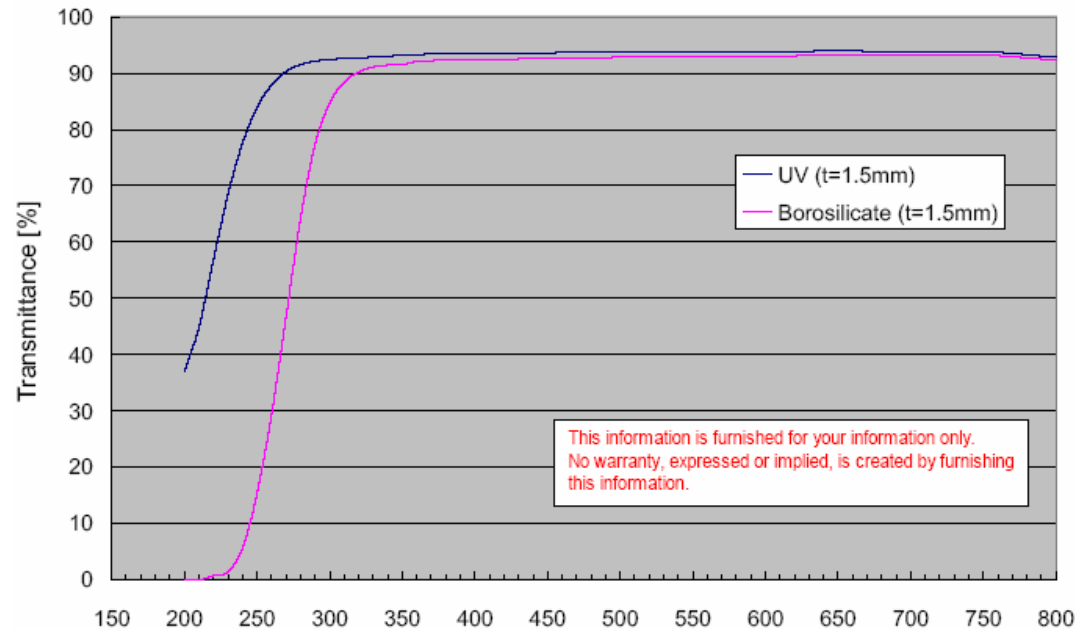
RICH detector – principle



- quantum efficiency limited by transmission/absorption of window
- examples: borosilicate – UV extended – quartz
- Hamamatsu H8300-03:

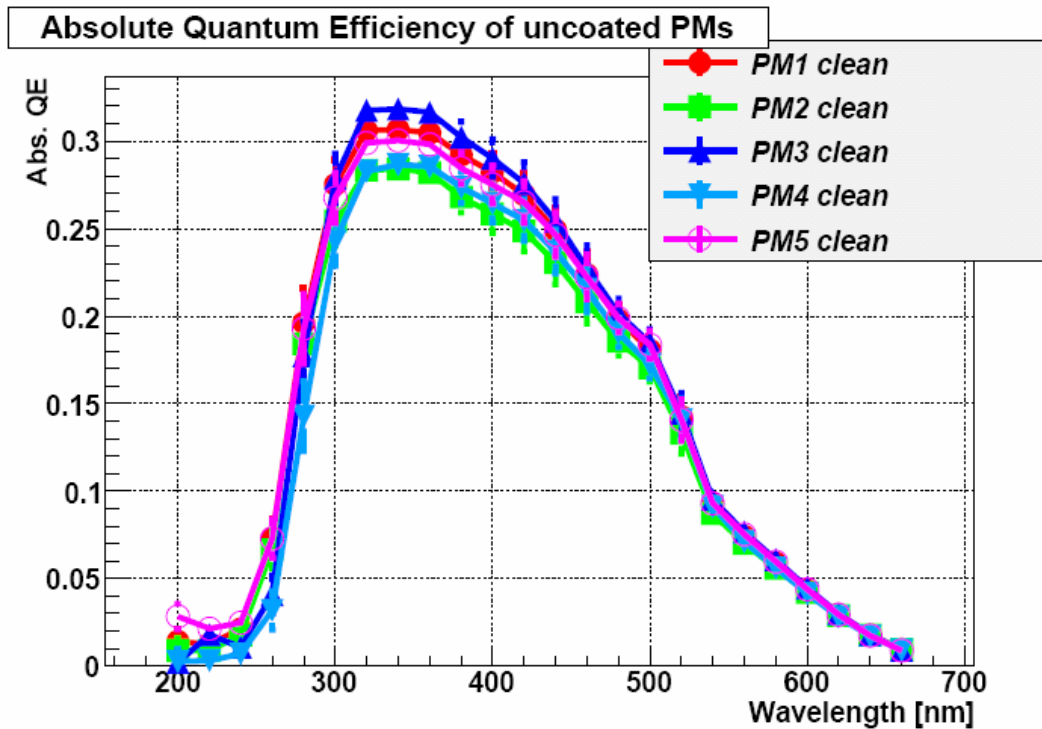


Spectral transmittance of the window



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this information.

Performance and limitations

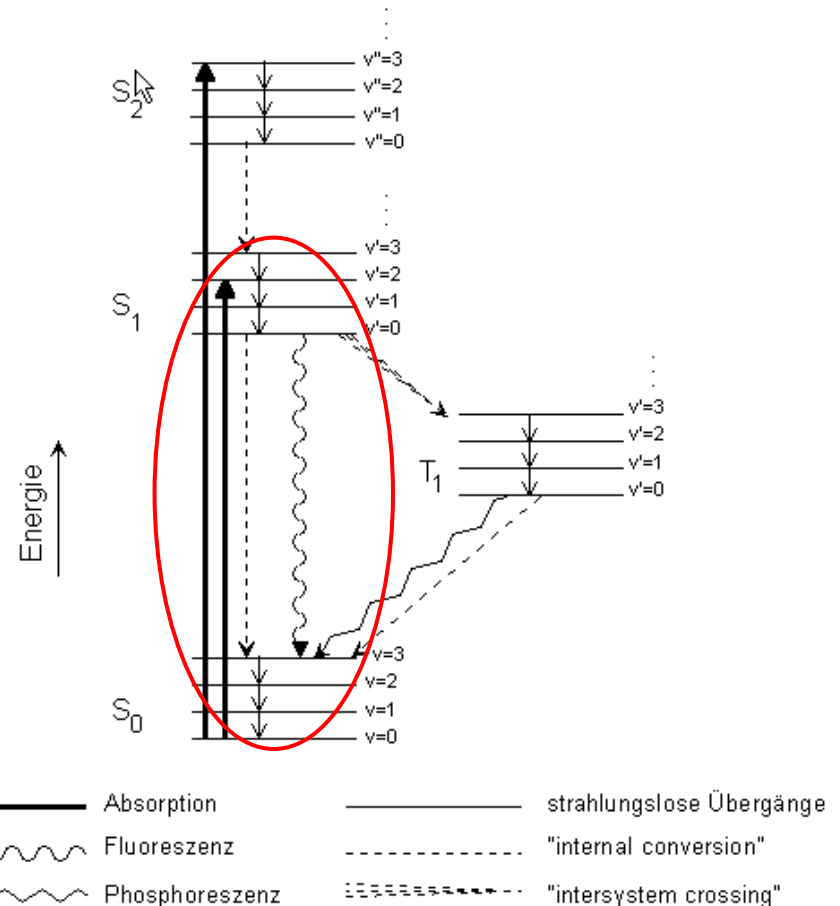


[Data: P.Koczon]

- PM: Photonis XP3102
- good quantum efficiency around 300 nm to 400 nm
- low efficiency below 300 nm

Wavelength shifting films – principle and application

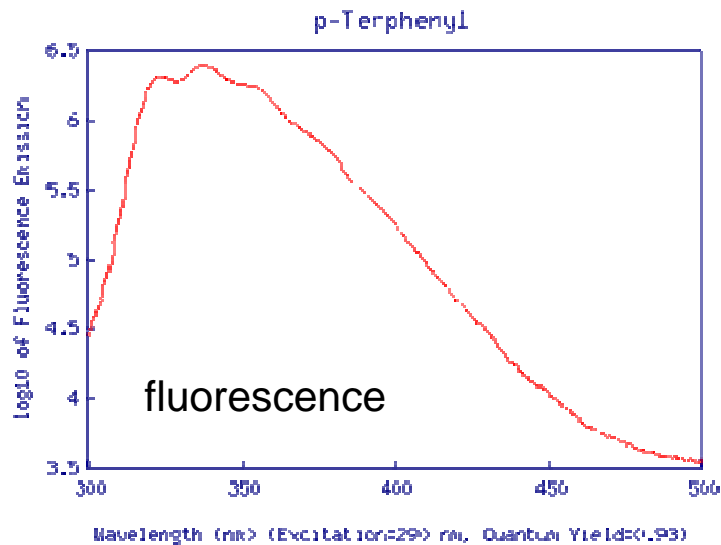
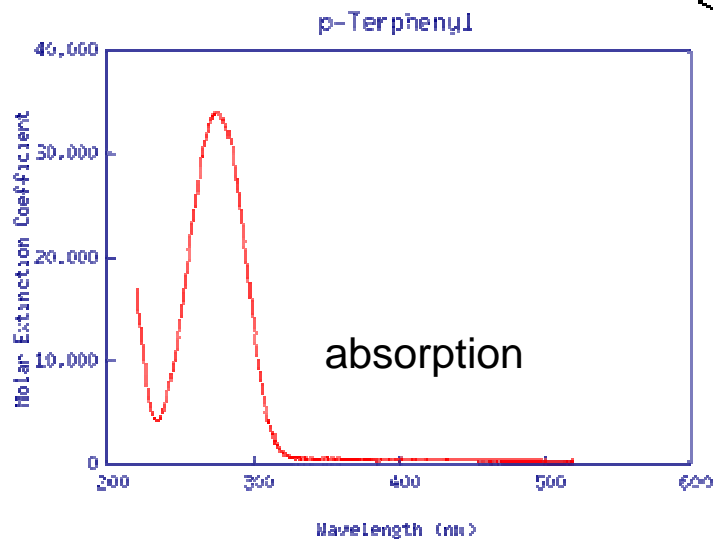
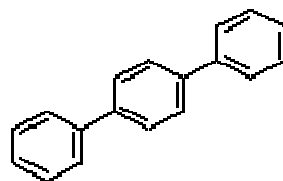
- Organic molecules absorbing in the short (UV) wavelength region
- Strong fluorescence in visible region



Wavelength shifting films – principle and application

- Organic molecules absorbing in the short (UV) wavelength region
- Strong fluorescence in visible region

Example: p-Terphenyl

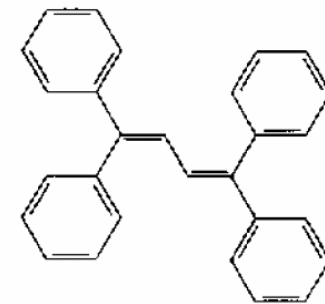
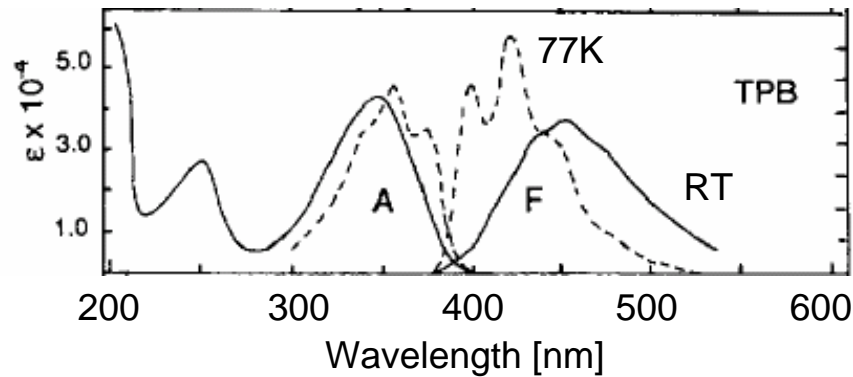


<http://omlc.ogi.edu/spectra/PhotochemCAD/html/p-terphenyl.html>

Wavelength shifting films – principle and application

- Organic molecules absorbing in the short (UV) wavelength region
- Strong fluorescence in visible region

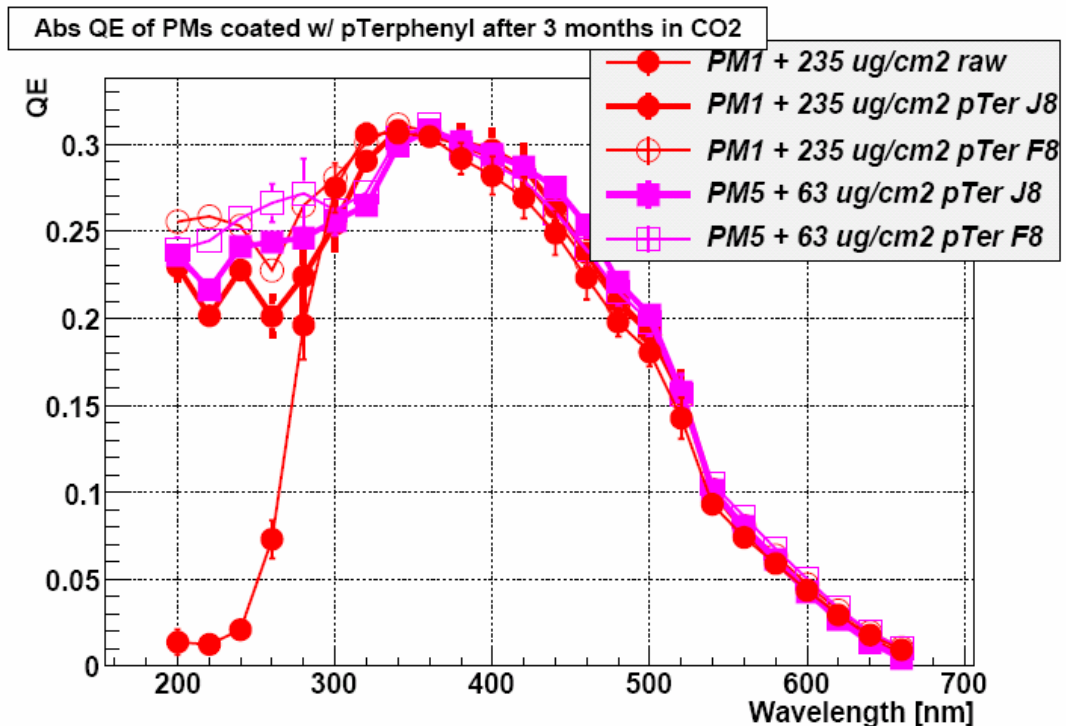
Example: TPB = Tetraphenyl Butadiene



Wavelength shifting films – principle and application

- Depending on material used, improved photomultiplier performance in the short wavelength region

P-Terphenyl:



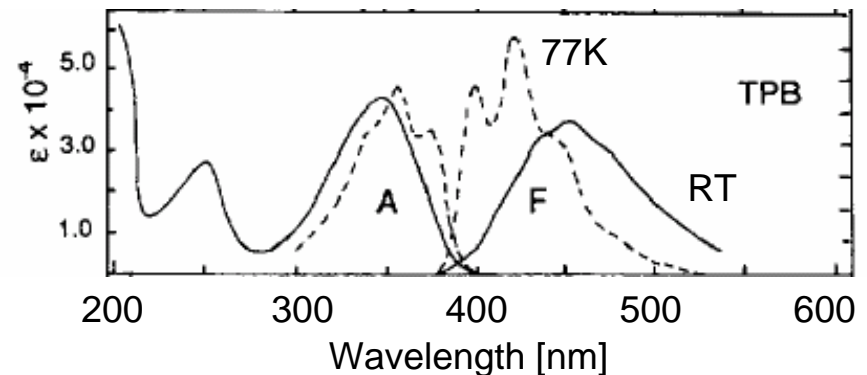
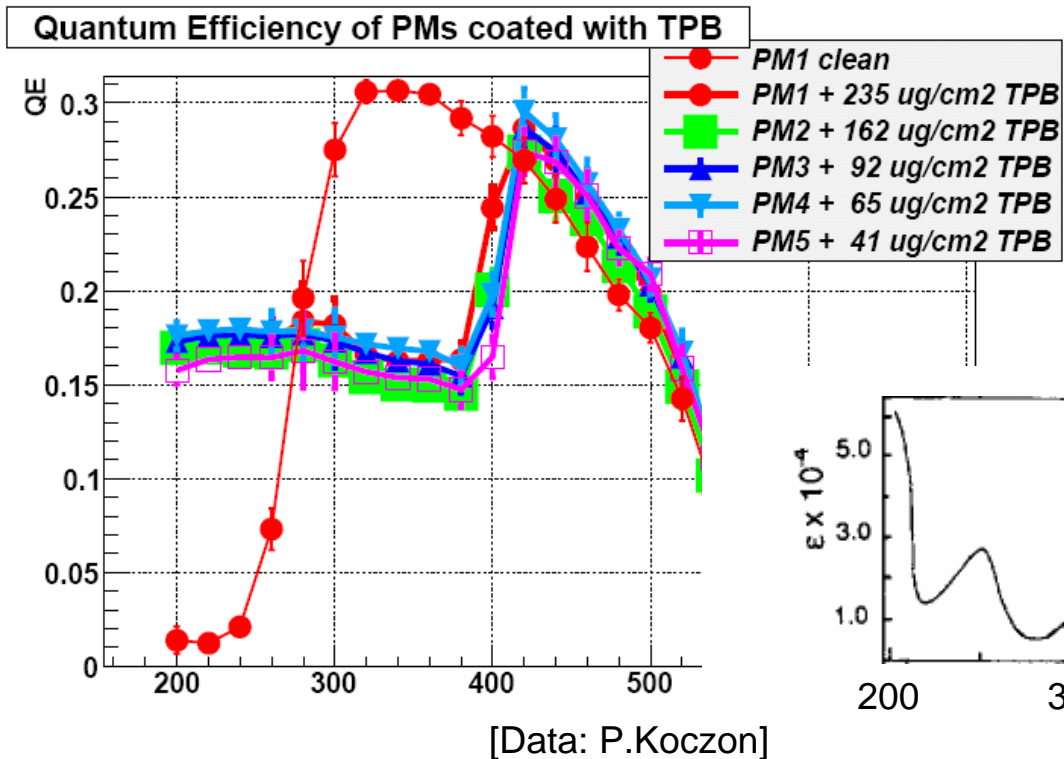
- good long-term stability when stored in CO₂ (under dark)

[Data: P.Koczon]

Wavelength shifting films – principle and application

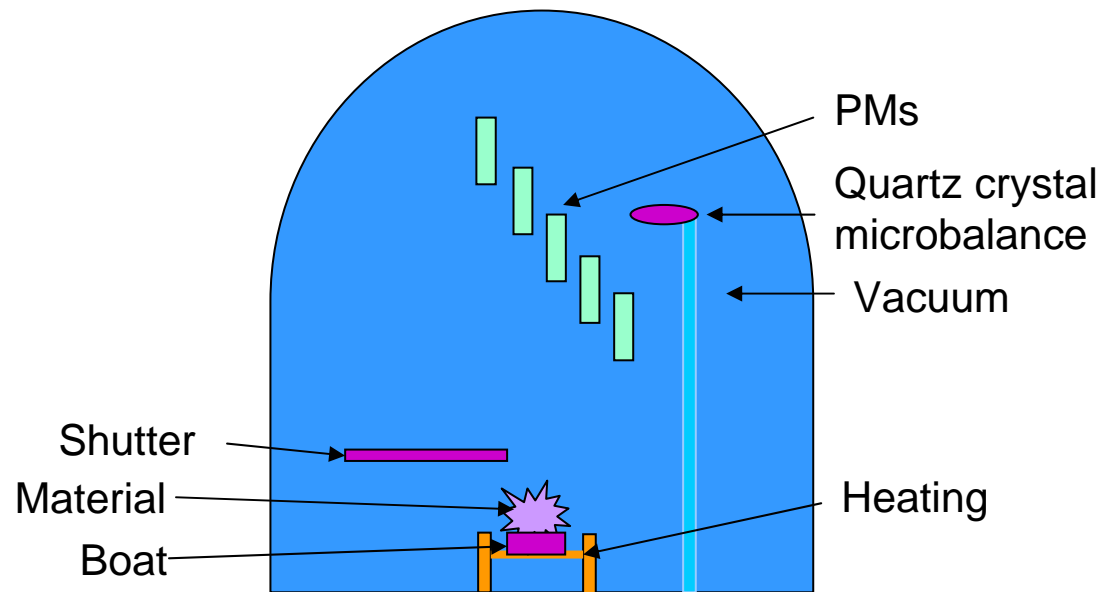
- Depending on material used, improved photomultiplier performance in the short wavelength region

TPB:



Application techniques

- Evaporation (used so far):
 - good optical properties (no solvents, no binders)
 - Inferior mechanical stability

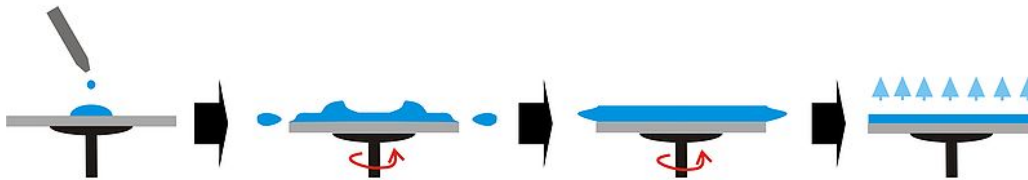


[Setup Andre Braem, CERN]

Application techniques

- Spin coating / dip coating
 - needs solvents → possible influence on optical properties
 - good mechanical stability

Spin coating:

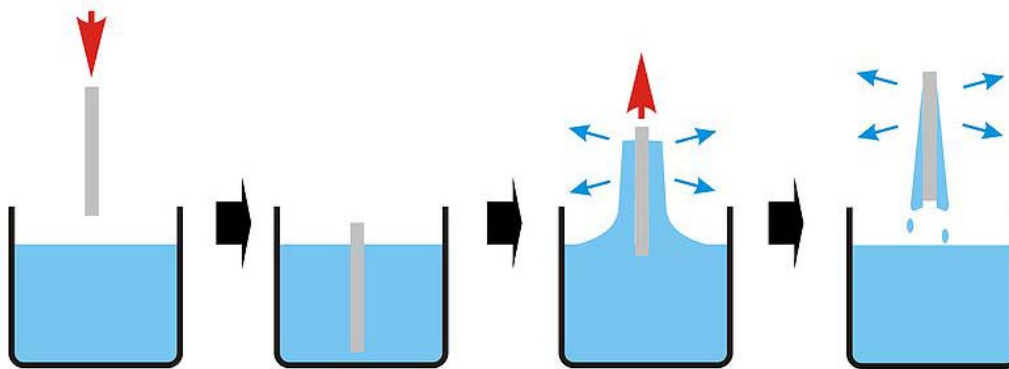


Application techniques

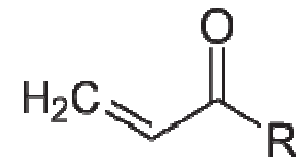
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- needs solvents → possible influence on optical properties
- good mechanical stability

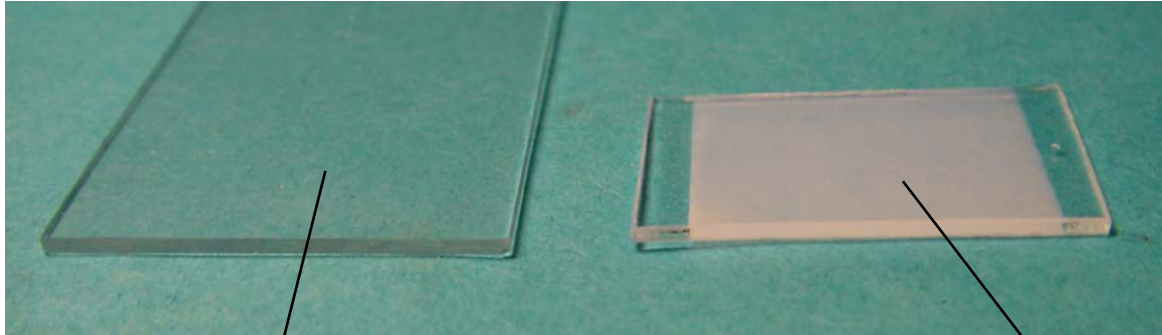
Dip coating:



- speed of removal determines thickness of films
- concentrations in use: 1g / L
- binder in use: paraloid (acrylate), 2 g /L



Comparison evaporated films – dip-coated films



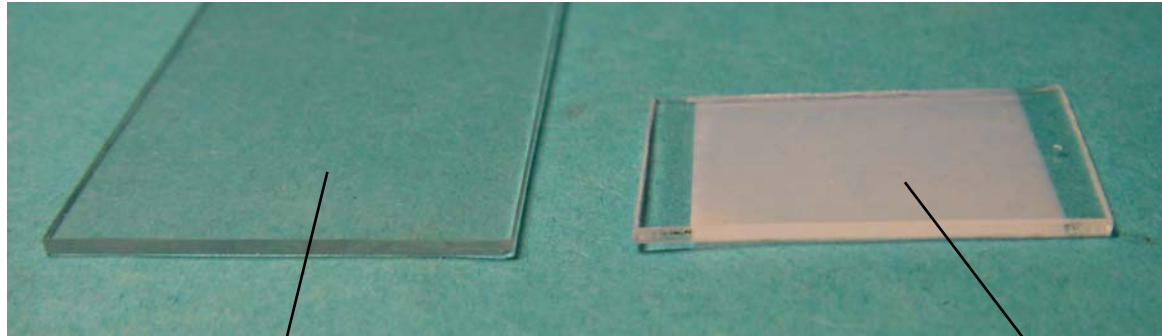
Dip-coated film, 6 cm/min

Evaporated film, 100 $\mu\text{g}/\text{cm}^2$

Dip-coated film:

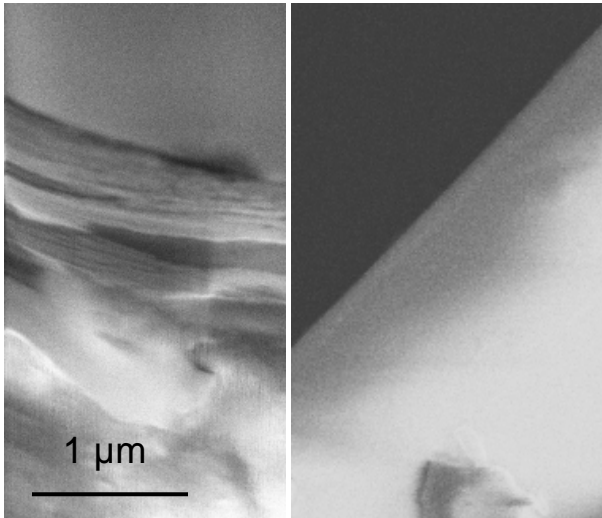
- scratch proof
- transparent

Comparison evaporated films – dip-coated films

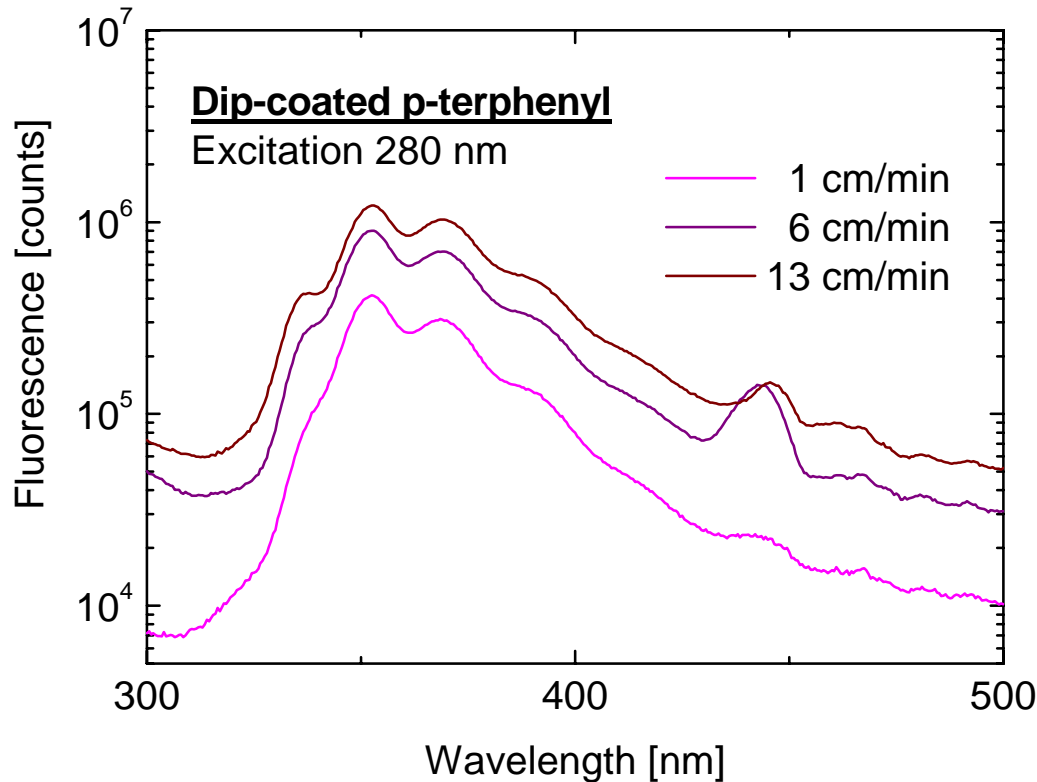


- Evaporation leads to microcrystals in the μm -regime
- ⇒ scattering of visible light

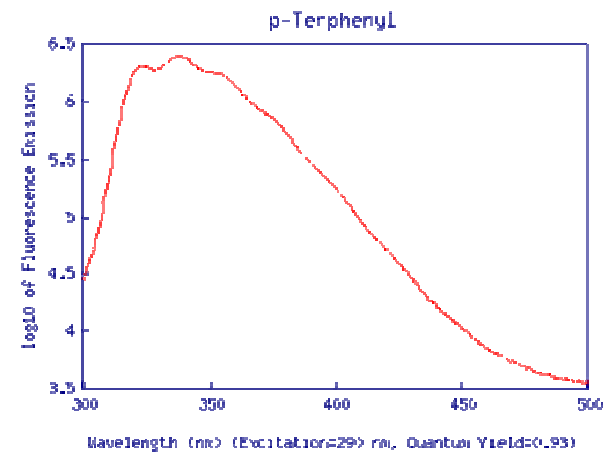
SEM images:



Fluorescence of dip-coated films



- Fluorescence spectrum comparable to reference
- thickness dependence (not all photons get absorbed)



To-do:

- Comparison evaporated/dip-coated films both with respect to fluorescence and efficiency of photomultipliers
- Test of photomultipliers
- Test of WLS-films on multi-anode structure (maybe in mini-RICH?)
- ... and much more ...

Thank you for your attention!